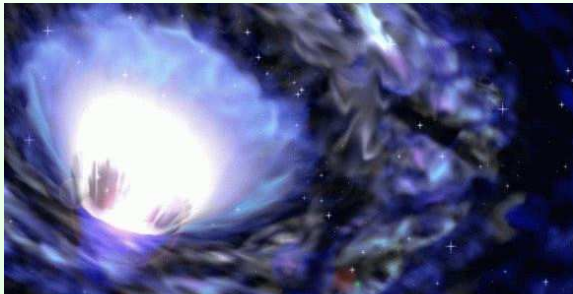


What Should Undergraduates Know About Gravitation?

Dr. Russell Herman

Department of Mathematics & Statistics, UNCW



Abstract

In July 2006 forty five physics faculty from nearly as many universities, including one from UNCW, met at Syracuse University to discuss the importance of teaching general relativity to undergraduates. This discussion came on the heels of a year long celebration of Albert Einstein's miracle year in 1905 which marked the beginnings of the theory of relativity and the ever increasing interest in gravitation and cosmology. There were numerous talks and posters on topics that could be used to present general relativity to undergraduates. These included gravitational radiation and LIGO, deviations from Newtonian gravitation and GPS, black holes, and experimental tests of general relativity. Leading authors of new undergraduate texts on general relativity contributed their thoughts on what students can learn about relativity. One of the goals of the workshop was to produce sample syllabi for courses aimed at three different audiences: general interest, physics intensive and mathematics intensive. In this talk we will discuss some of the ideas that came out of this workshop and, of course, talk a bit about general relativity and the growing need to introduce it into the physics curriculum.

Outline

- 1 TGRU Workshop
 - Web Site
 - Why? Who? What?
- 2 Special Relativity
 - Postulates
 - Spacetime Diagrams
- 3 General Relativity
 - Principles
 - Flat Space
 - Curved Space
- 4 Course Syllabi
 - General Interest
 - Physics First
 - Math Intensive
- 5 Textbooks
- 6 Poster
- 7 Summary

Teaching General Relativity to Undergraduates

AAPT TGRU site

<http://www.aapt-doorway.org/TGRU/>

The workshop was supported by the **LIGO Project**, the **Center for Gravitational Wave Physics** at Penn State, the **American Association of Physics Teachers**, and the **Syracuse University Department of Physics**.

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Information at Web Site

- Talks
- Posters
- Articles - Pedagogy/Research
- Links to Texts
- Course Design/Syllabi

Why Blend SR and GR into the Curriculum?

A Few Answers ...

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- **Laser Interferometer Gravitational-Wave Observatory**
- $E = mc^2$ is more recognized than $F = ma$.

Who, What?

Who Should Do It?

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- Do not need relativity experts.

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Approaches - Math or Physics First?

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- Math First
 - develop math foundations (tensors and differential geometry).

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- Do not need relativity experts.
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- Gravitational Waves
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Approaches - Math or Physics First?

- Math or Physics First?
- Math First
 - develop math foundations (tensors and differential geometry).
 - Einstein's Equations and then Applications.

Who, What?

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- Black holes
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Approaches - Math or Physics First?

- Math or Physics First?
- Physics First
 - Applications first and then mathematical background.
 - Understanding gravity as curved spacetime via examples.

Who, What?

Who Should Do It?

- Do not need relativity experts.
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What Topics?

- Black holes
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Approaches - Math or Physics First?

- Math or Physics First?
- General Interest
Possibly the only physics course.

Before I forget ...

What Should Undergraduates Know About Gravity?

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- Gravity is a universal interaction.

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- Gravity is a universal interaction.
 - Newtonian physics - between all masses

Before I forget ...

What Should Undergraduates Know About Gravity?

- Gravity is a universal interaction.
 - Relativistic - between all forms of energy

Before I forget ...

What Should Undergraduates Know About Gravity?

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- Gravity is unscreened and always attractive

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- Gravity is the weakest fundamental interaction.

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What Should Undergraduates Know About Gravity?

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GRAVITY = GEOMETRY

$$G_{\mu\nu} = T_{\mu\nu}$$

Relativity at a Young Age

Start working **NOW**
on your entry
for the

SECOND ANNUAL GREATER PHILADELPHIA

SCIENCE FAIR

MAY 6 thru MAY 14

Sponsored by
THE PHILADELPHIA INQUIRER

Special Relativity

Postulates

Special Relativity

Postulates

- Speed of light is constant

Special Relativity

Postulates

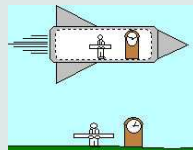
- Speed of light is constant
- Physics is the same for all inertial observers

Special Relativity

Postulates

- Speed of light is constant
- Physics is the same for all inertial observers

Consequences



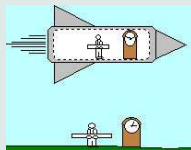
Special Relativity

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- Speed of light is constant
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Consequences

- Simultaneity



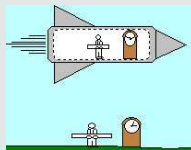
Special Relativity

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Consequences

- Simultaneity
- Time Dilation



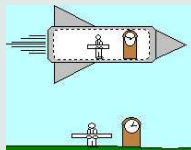
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Consequences

- Simultaneity
- Time Dilation
- Length Contraction



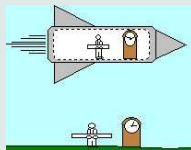
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- Simultaneity
- Time Dilation
- Length Contraction
- $E = mc^2$ and more!



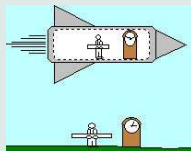
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Noted Concerns

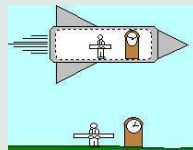
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Noted Concerns

- Reference Frames vs Coordinate Systems

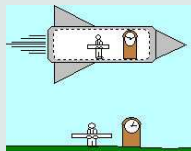
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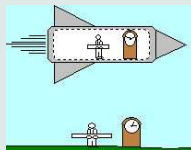
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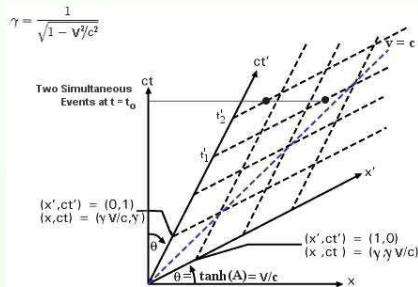
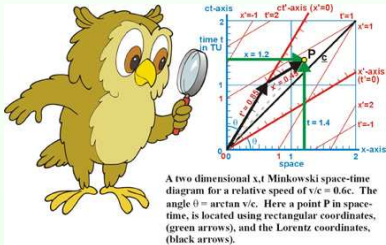
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Noted Concerns

- Reference Frames vs Coordinate Systems
- Simultaneity Misconceptions
- No Spacetime Diagrams

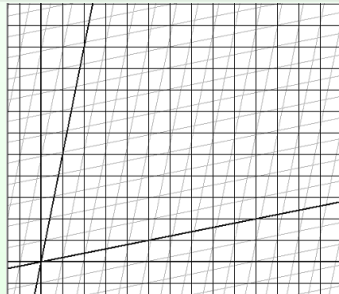
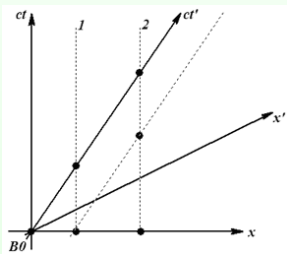
Minkowski Spacetime



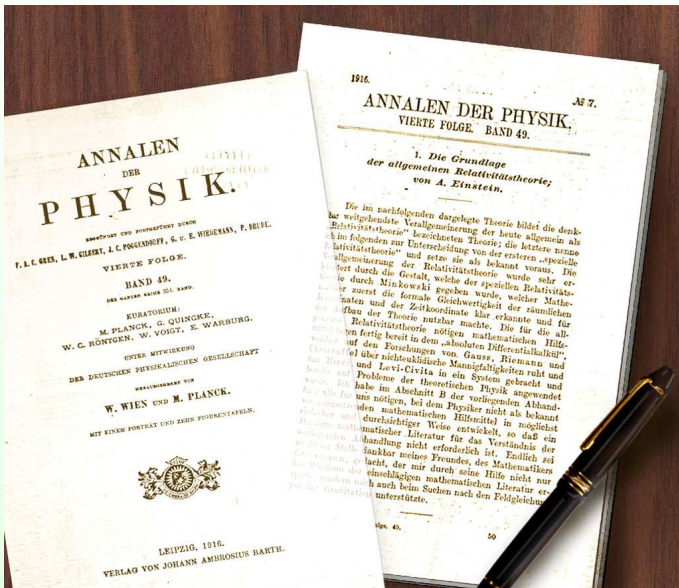
Spacetime Diagram Application

Train in Tunnel Problem

A relativistic train of rest length 240 meters travels at $0.6c$ through a tunnel which has rest length 360 meters. In the figure below the world lines for the tunnel openings are drawn as line 1 and 2 and the world line of the front of the train is the third dotted line. Let S_{tunnel} be the tunnel with coordinates (x, t) and let S_{train} be the train coordinates (x', t') . We set the origin as the event B_0 , the back of the train location just as the front end enters opening 1.



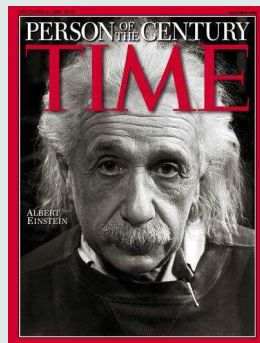
General Relativity - 1916



The Problem of the Century

from the Person of the Century - Smarr, 2000

- 1910s General Theory; Schwarzschild
- 1920s Equation of Motion Posed
- 1930s Two Body Problem Posed
- 1940s Cauchy Problem Posed
- 1950s Numerical Relativity Conceived
- 1960s Geometrodynamics; First NR Attempts
- 1970s Head-On Spacetime Roughed Out
- 1980s NR Becomes a Field
- 1990s Head-On Nailed; 3D Dynamics Begins
- 2000s 3D Dynamics Nailed; Gravitational Wave Astronomy



General Relativity

Postulates

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- **General Principle of Relativity** - The laws of physics are the same for all observers.

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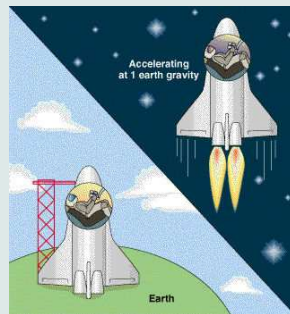
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- **Spacetime Curvature Caused by Stress-Energy** - Described by Einstein field equations.

General Relativity

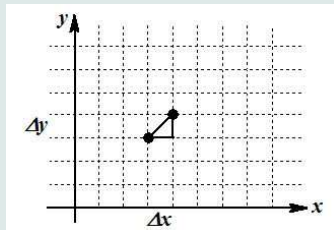
The Equivalence Principle

Experiments in (sufficiently small) freely falling laboratory, over a short time, give results that are indistinguishable from those experiments in an inertial frame in empty space



Spacetime

Flat Spaces - Cartesian

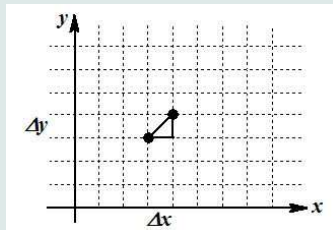


Spacetime

Flat Spaces - Cartesian

- **Line element**

$$dS^2 = dx^2 + dy^2$$



Spacetime

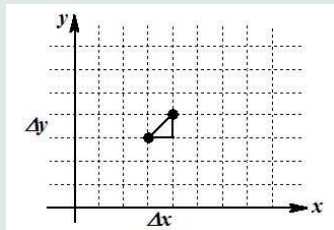
Flat Spaces - Cartesian

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$$dS^2 = dx^2 + dy^2$$

- $S = \int_A^B L\left(\frac{dx}{d\sigma}, \frac{dy}{d\sigma}, x, y\right) d\sigma$

$$L = \sqrt{\left(\frac{dx}{d\sigma}\right)^2 + \left(\frac{dy}{d\sigma}\right)^2}$$



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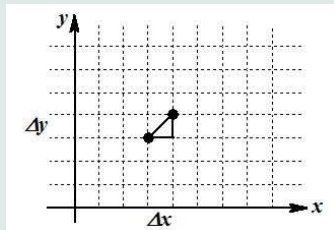
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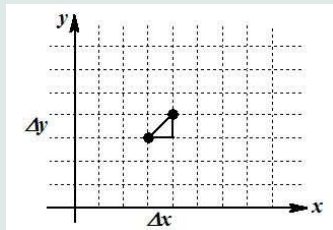
- **Variational Principle**

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- **Euler-Lagrange Eq.**

$$\frac{d}{d\sigma} \left(\frac{\partial L}{\partial \dot{x}} \right) = \frac{\partial L}{\partial x},$$

$$\frac{d}{d\sigma} \left(\frac{\partial L}{\partial \dot{y}} \right) = \frac{\partial L}{\partial y}$$



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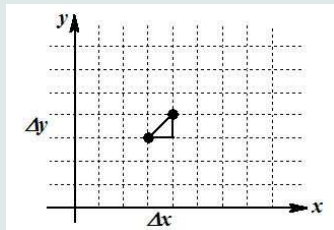
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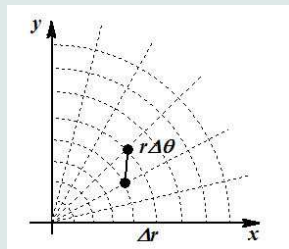
- **Geodesics**

$$\frac{d^2 x}{dS^2} = 0, \quad \frac{d^2 y}{dS^2} = 0$$



Spacetime

Flat Spaces - Polar

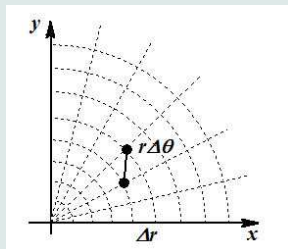


Spacetime

Flat Spaces - Polar

- **Line element**

$$dS^2 = dr^2 + (r d\theta)^2$$



Spacetime

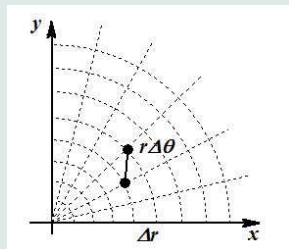
Flat Spaces - Polar

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$$dS^2 = dr^2 + (r d\theta)^2$$

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$$\int_A^B \sqrt{\left(\frac{dr}{d\sigma}\right)^2 + r^2 \left(\frac{d\theta}{d\sigma}\right)^2} d\sigma$$



Spacetime

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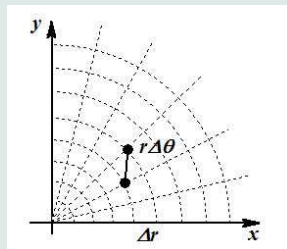
- **S =**

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$$\frac{d}{dS} \left(\frac{\partial L}{\partial \dot{r}} \right) = \frac{\partial L}{\partial r},$$

$$\frac{d}{dS} \left(\frac{\partial L}{\partial \dot{\theta}} \right) = \frac{\partial L}{\partial \theta}$$



Spacetime

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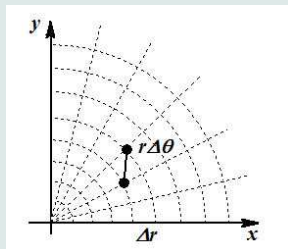
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$$\frac{d}{dS} \left(\frac{\partial L}{\partial \dot{\theta}} \right) = \frac{\partial L}{\partial \theta}$$

- **Geodesics**

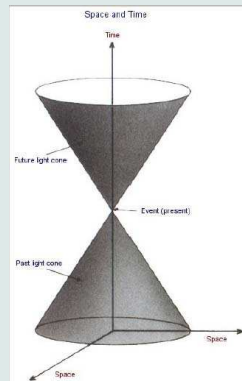
$$\frac{\partial^2 r}{\partial S^2} = r \left(\frac{d\theta}{dS} \right)^2,$$

$$\frac{d}{dS} \left(r^2 \frac{d\theta}{dS} \right) = 0.$$



Spacetime

Minkowski Space

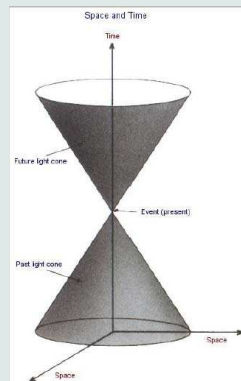


Spacetime

Minkowski Space

- **Line element**

$$ds^2 = -(cdt)^2 + dx^2 + dy^2 + dz^2$$



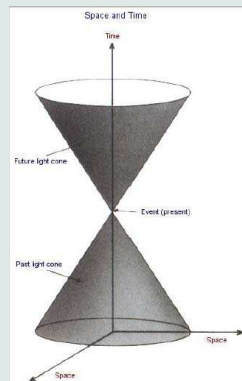
Spacetime

Minkowski Space

- **Line element**

$$ds^2 = -(cdt)^2 + dx^2 + dy^2 + dz^2$$

- **Proper time** $d\tau^2 = -ds^2/c^2$



Spacetime

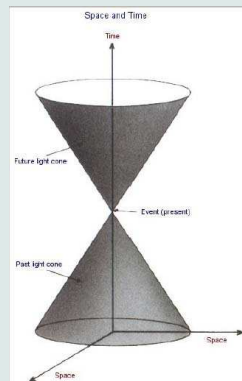
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- $\tau_{AB} = \int_{\tau_A}^{\tau_B} \sqrt{1 - V^2(t')/c^2} dt'$ or
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Spacetime

Minkowski Space

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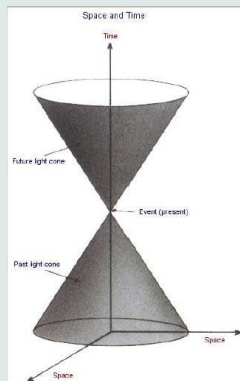
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- **Variational Principle** $\delta\tau = 0$



Spacetime

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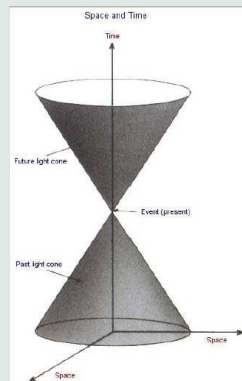
$$d\tau = dt \sqrt{1 - V^2/c^2}$$

- **Variational Principle** $\delta\tau = 0$

- **Euler-Lagrange Eq.**

$$\frac{d}{dS} \left(\frac{\partial L}{\partial \dot{x}} \right) = \frac{\partial L}{\partial x},$$

$$L = \sqrt{-\eta_{\alpha\beta} \frac{dx^\alpha}{d\sigma} \frac{dx^\beta}{d\sigma}}$$



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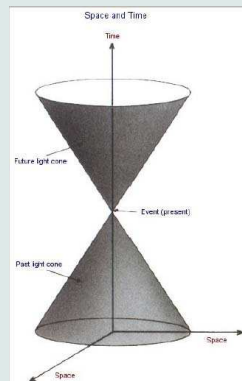
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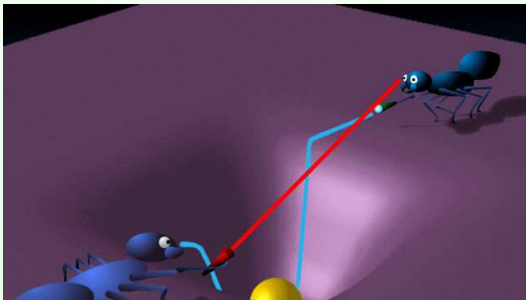
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- **Geodesics** $\frac{d^2 x^\alpha}{d\tau^2} = 0$



Spacetime

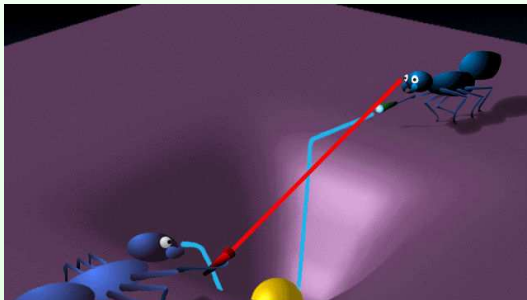
Curved Spaces



Spacetime

Curved Spaces

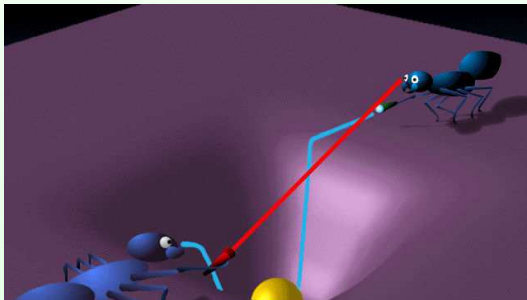
- **Line element** $ds^2 = g_{\alpha\beta} dx^\alpha dx^\beta$



Spacetime

Curved Spaces

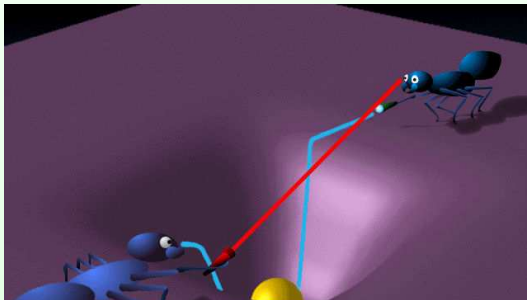
- **Line element** $ds^2 = g_{\alpha\beta} dx^\alpha dx^\beta$
- $T_{AB} = \int_{T_A}^{T_B} \sqrt{-g_{\alpha\beta} \frac{dx^\alpha}{d\sigma} \frac{dx^\beta}{d\sigma}} d\sigma$



Spacetime

Curved Spaces

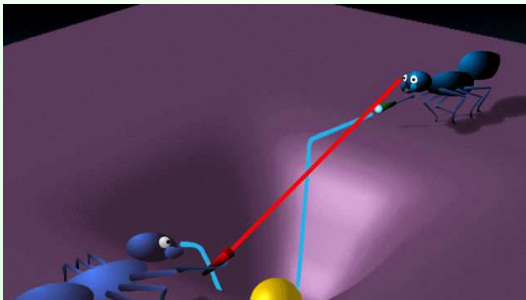
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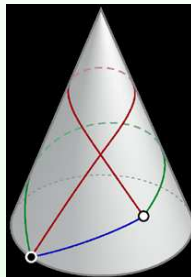
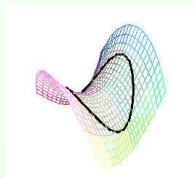
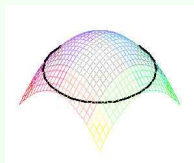
Spacetime

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- **Geodesics** $\frac{d^2 x^\alpha}{d\tau^2} + \Gamma_{\beta\gamma}^\alpha \frac{dx^\beta}{d\tau} \frac{dx^\gamma}{d\tau} = 0$

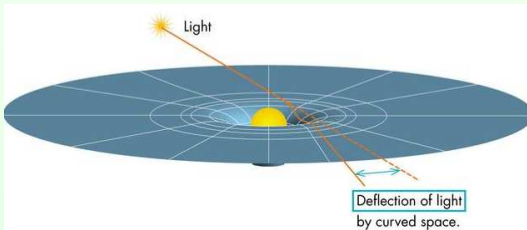


Geodesics



Spacetime

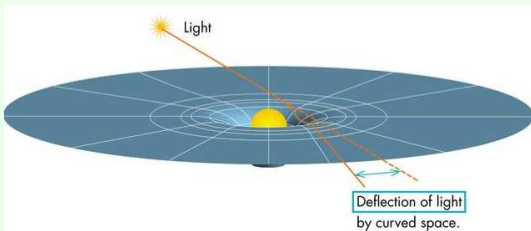
Schwarzschild



Spacetime

Schwarzschild

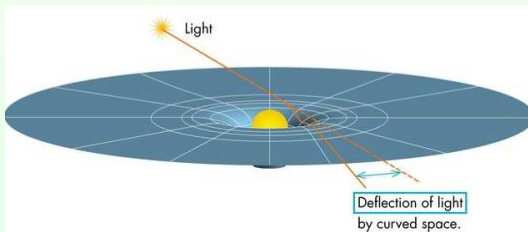
- **Line element** $ds^2 = -\left(1 - \frac{2M}{a}\right) dt^2 + \left(1 - \frac{2M}{a}\right)^{-1} dr^2 + r^2 \left(d\theta^2 + \sin^2 \theta d\phi^2\right)$



Spacetime

Schwarzschild

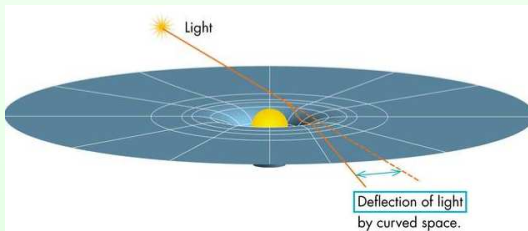
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Spacetime

Schwarzschild

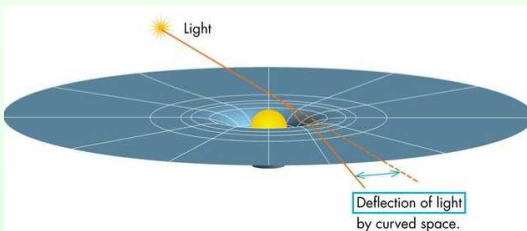
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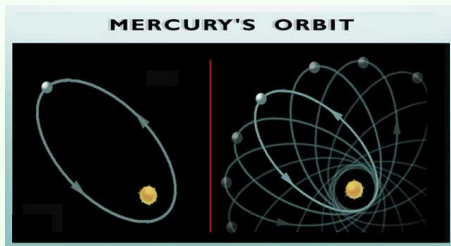
Spacetime

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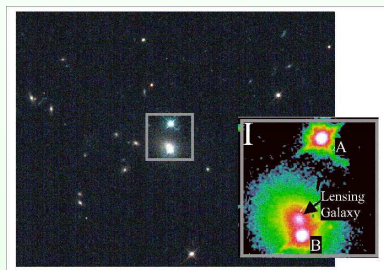
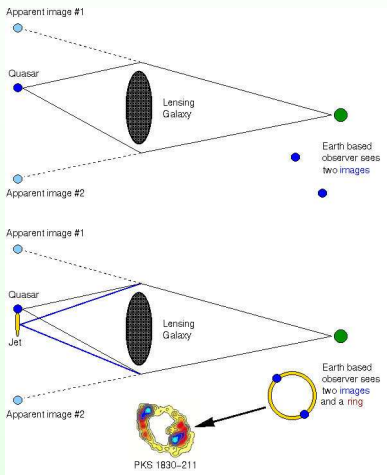
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- $ds^2 = -\left(1 + \frac{2\Phi}{c^2}\right) (cdt)^2 + \left(1 - \frac{2\Phi}{c^2}\right) dS^2$



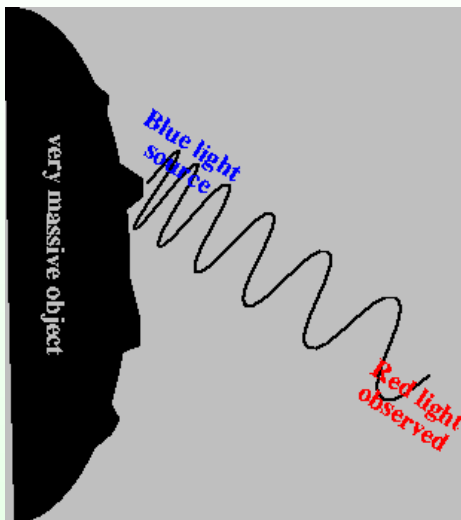
Classical Tests - Perihelion Shift



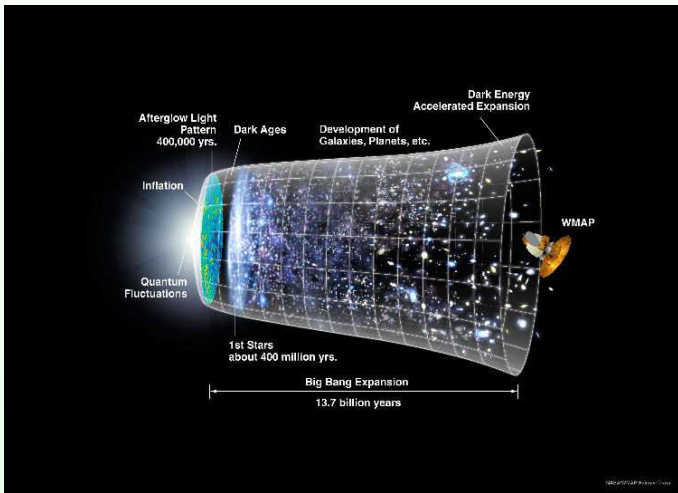
Classical Tests - Bending of Light



Classical Tests - Gravitational Red Shift

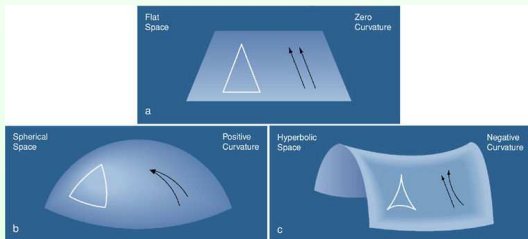


The Big Bang!



Cosmology

FRW model

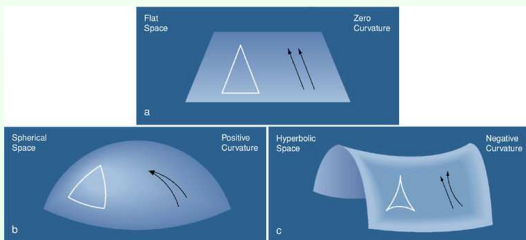


Cosmology

FRW model

- Line element

$$ds^2 = -dt^2 + a^2(t) \left[\frac{dr^2}{1-kr^2} + r^2 (d\theta^2 + \sin^2 \theta d\phi^2) \right]$$



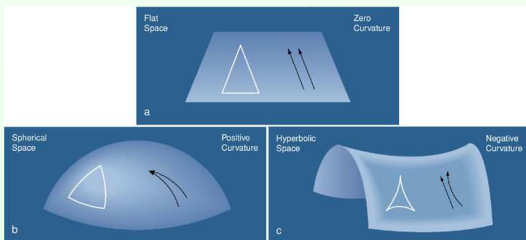
Cosmology

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- Friedman equation $\dot{a}^2 - \frac{8\pi\rho}{3}a^2 = -k$



Cosmology

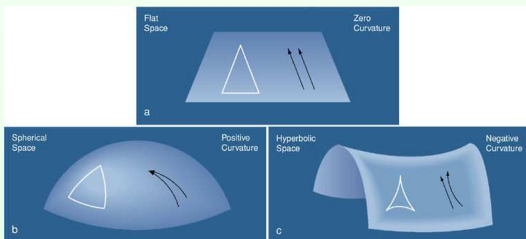
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- **Total Density** $\rho = \left(\Omega_v + \frac{\Omega_m}{a^3} + \frac{\Omega_r}{a^4} \right) \rho_{\text{crit}}$



Applications of General Relativity

There are many more applications ...

Global Positioning System (GPS)
 Cosmological redshift
 Big-bang
 Gravitational lensing
 Propagation of gravitational waves
 Spherical gravitational collapse
 Active Galactic Nuclei
 Frame-dragging by a rotating body
 Gravitational redshift
 Formation of black holes
 Neutron stars
 Cosmic Background Radiation

Lense-Thirring precession of a gyroscope
 Bending of light by the Sun
 Shapiro time delay
 The fate of the universe
 Determining parameters of binary pulsars
 X-ray sources
 Hawking radiation from black holes
 Expansion of the universe
 Accretion disks around compact objects
 Operation of gravitational wave detectors
 Precession of Mercury's perihelion
 ...

Syllabi - General Interest

The Narrative

Syllabi - General Interest

The Narrative

- Explain why we need GR, but do not explain GR.

Syllabi - General Interest

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Syllabi - General Interest

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Syllabi - General Interest

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Syllabi - General Interest

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 - Correct GR Simulation

Syllabi - General Interest

Syllabus

Topics	Input	GR Application
Tides Light Shifting	Equivalence Principle Doppler	GPS with corrections. Doppler shifts, Reference Frames
Light bending Black Holes CMB/Hubble	Why doesn't ISS fall? Orbits Spectral Lines, Escape Velocity Data	Correct bending, Lensing Maser data Cosmology

Syllabi - General Interest

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Possible Activities

Syllabi - General Interest

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Black Holes	Spectral Lines, Escape Velocity	Maser data
CMB/Hubble	Data	Cosmology

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- Taylor-Wheeler projects
- Same acceleration in vacuum for all
- Elevator acceleration
- Construct spacetime diagrams from a movie

Syllabi - Physics First

The Logical Order

Syllabi - Physics First

The Logical Order

- Assemble the necessary mathematical tools.

Syllabi - Physics First

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Syllabi - Physics First

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- Gets quickly to physical effects

Syllabi - Physics First

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- Gets quickly to physical effects
- More flexible timing

Syllabi - Physics First

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Advantages

- Gets quickly to physical effects
- More flexible timing
- More closely connected to curriculum - Mechanics!

Connections to Mechanics

First Integrals

Connections to Mechanics

First Integrals

- Test particles in Schwarzschild metric.

$$E = \frac{1}{2} \left(\frac{dr}{d\tau} \right)^2 + V_{\text{eff}}, \quad V_{\text{eff}} = -\frac{M}{r} + \frac{\ell^2}{2r^2} - \frac{M\ell^2}{r^3}$$

Connections to Mechanics

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- Light rays in Schwarzschild metric.

$$\frac{1}{b^2} = \frac{1}{\ell^2} \left(\frac{dr}{d\lambda} \right)^2 + W_{\text{eff}}, \quad W_{\text{eff}} = \frac{1}{r^2} \left(1 - \frac{2M}{r} \right)$$

Connections to Mechanics

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- FRW Cosmological Models.

$$\frac{1}{2} \left(\frac{da}{dt} \right)^2 + U_{\text{eff}} = \frac{\Omega_c}{2}, \quad U_{\text{eff}} = -\frac{1}{2} \left(\Omega_v a^2 + \frac{\Omega_m}{a} + \frac{\Omega_r}{a^2} \right)$$

Connections to Mechanics

First Integrals

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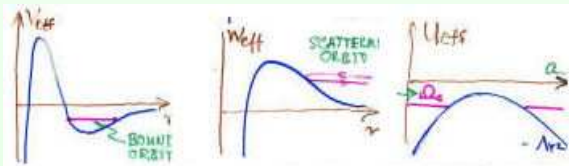
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Syllabi - Physics First

Things to be included

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Course Themes

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Course Themes

- Black Holes

Syllabi - Physics First

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Course Themes

- Black Holes
- Cosmology

Syllabi - Physics First

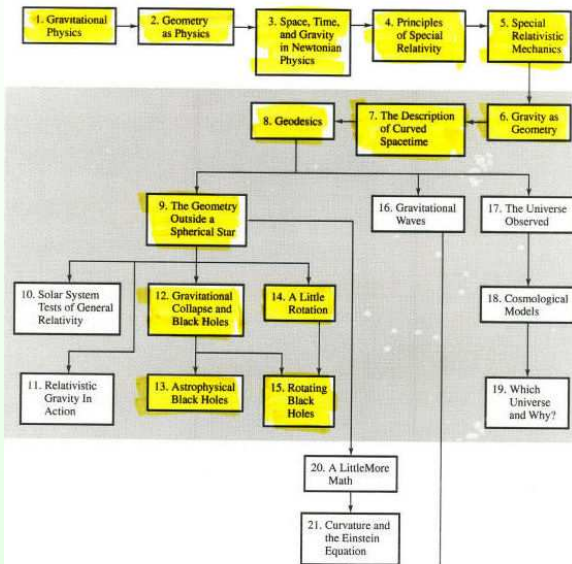
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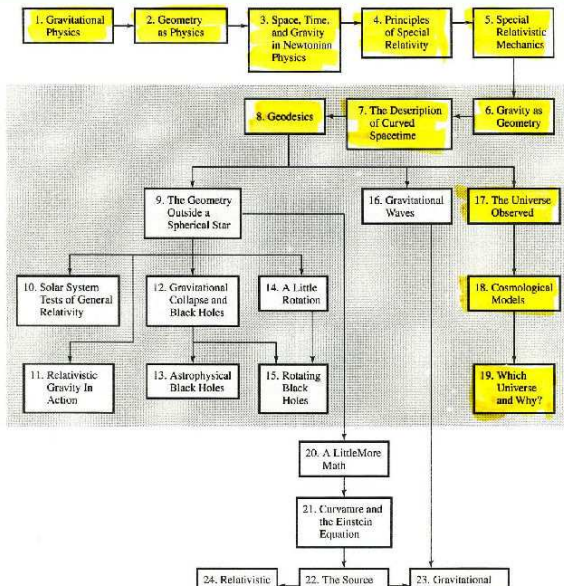
Course Themes

- Black Holes
- Cosmology
- Gravitational Waves

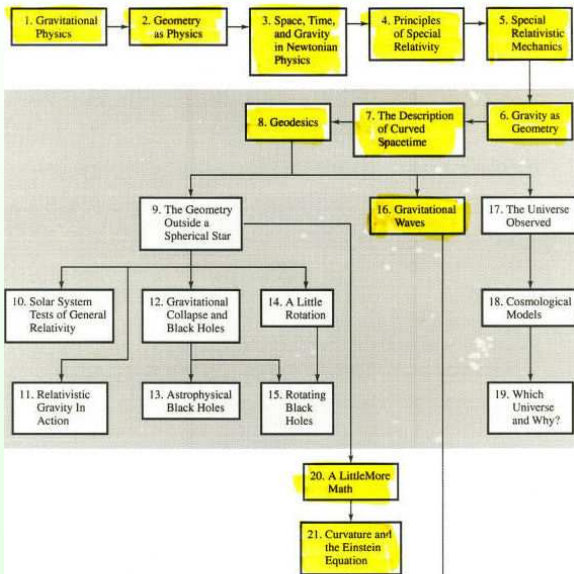
Syllabus - Black Hole Emphasis



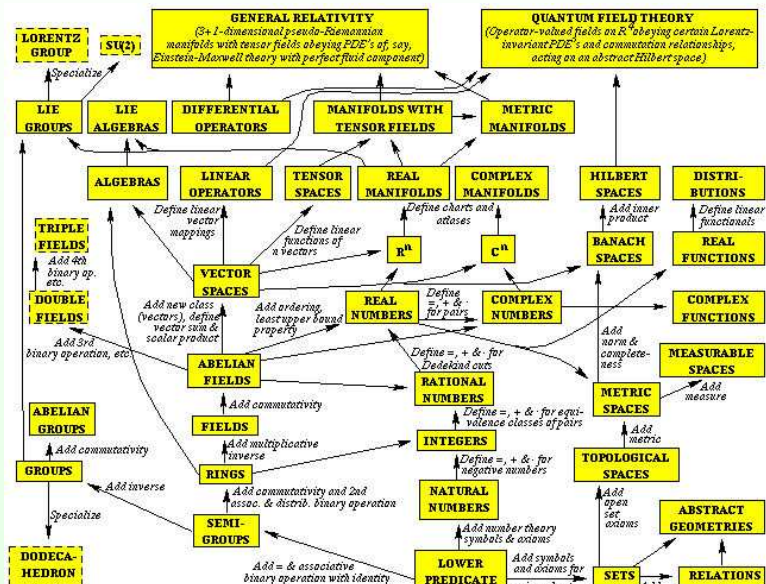
Syllabus - Cosmology Emphasis



Syllabus - Gravitational Wave Emphasis



Syllabi - Math Intensive



Syllabi - Math Intensive

The 7-Fold Way - Tom Moore

Syllabi - Math Intensive

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- Blend Math and Physics.

Syllabi - Math Intensive

The 7-Fold Way - Tom Moore

- Blend Math and Physics.
- Use lots of 2D Examples.

Syllabi - Math Intensive

The 7-Fold Way - Tom Moore

- Blend Math and Physics.
- Use lots of 2D Examples.
- Keep It Suitably Simple.

Syllabi - Math Intensive

The 7-Fold Way - Tom Moore

- Blend Math and Physics.
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Treat Tensors as Generalized Vectors

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- Tensors represent physical objects

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- Tensors represent physical objects
- Tensors have components relative to a basis

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Treat Tensors as Generalized Vectors

- Tensors represent physical objects
- Tensors have components relative to a basis
- Raise/lower indices to embed metric

Syllabi - Math Intensive

The 7-Fold Way - Tom Moore

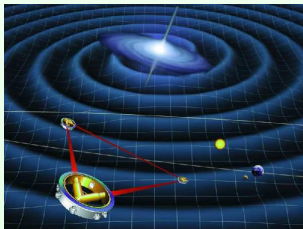
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Treat Tensors as Generalized Vectors

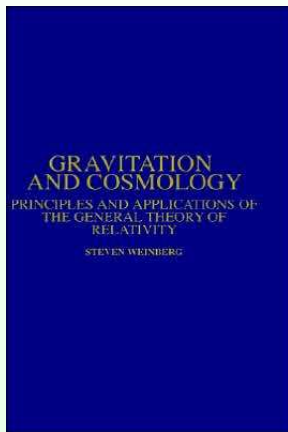
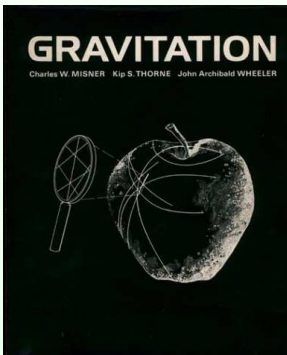
- Tensors represent physical objects
- Tensors have components relative to a basis
- Raise/lower indices to embed metric
- Tensors provide firm foundation for understanding

Syllabi - Math Intensive

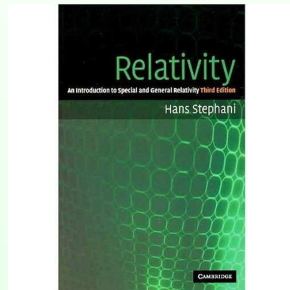
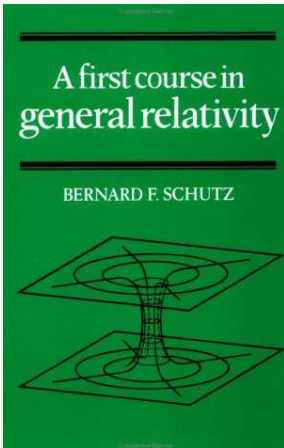
1	Conceptual Overview	Review of Relativity	Four-Vectors
2	Index Notation	Arbitrary Coordinates	Tensor Equations
3	Maxwell's Equations	Geodesics	The Schwarzschild Metric
4	Particle Orbits	Perihelion Precession	Photon Orbits
5	Gravitational Lenses	Event Horizon	Alternative Coordinates
6	BH Thermodynamics	The Kerr Metric	Kerr Particle Orbits
7	Ergoregion and Horizon	Negative Energy Orbits	The Penrose Process
8	The Absolute Gradient	Geodesic Deviation	The Riemann Tensor
9	Stress Energy Tensor	The Einstein Equation	Interpreting the Equation
10	Schwarzschild Solution	The Observed Universe	A Cosmic Metric
11	Evolution of the Universe	Cosmic Implications	The Early Universe
12	Linearized Gravity	Gauge Freedom	Gravitational Waves
13	"Energy" in GWs	Generation of GWs	Applications



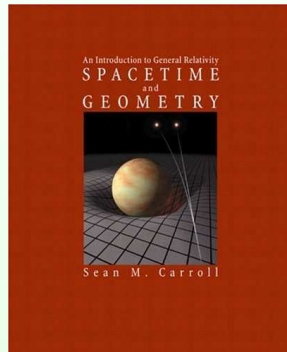
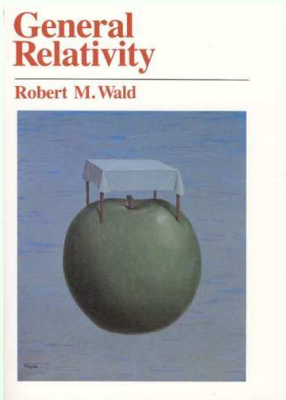
Texts - Old



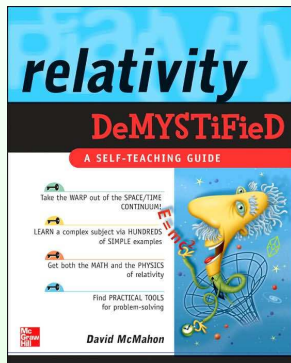
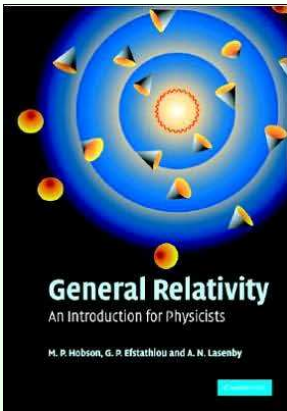
Texts - Semi-Old



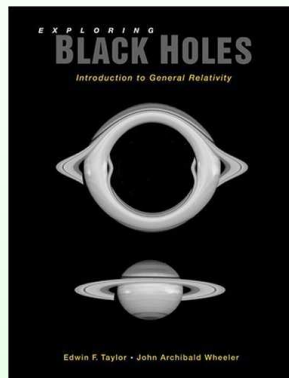
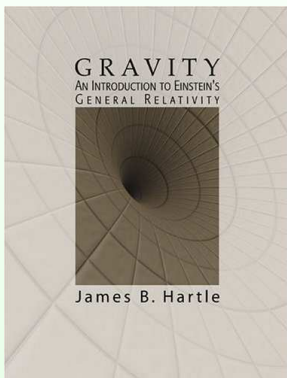
Texts - Math Intensive



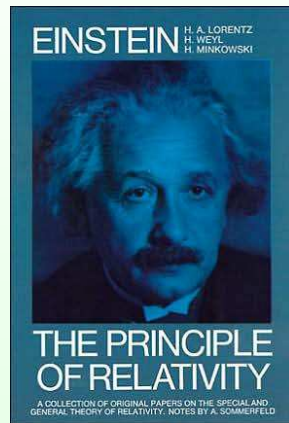
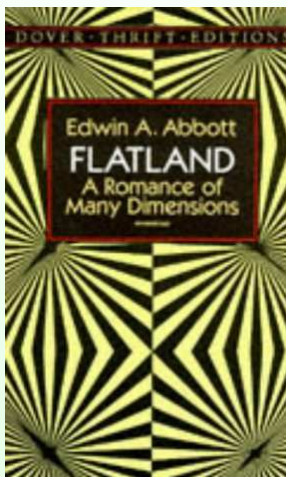
Texts - Math Intensive



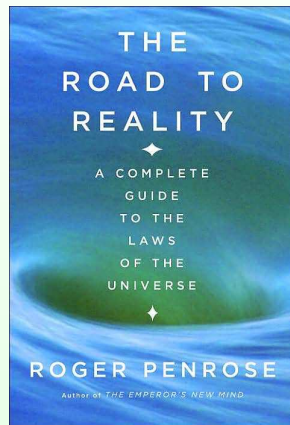
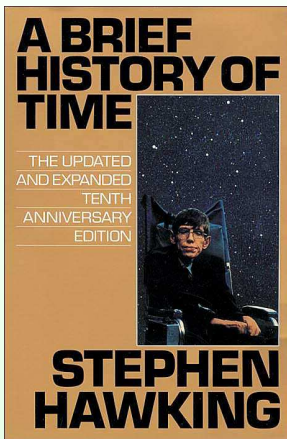
Texts - Physics First



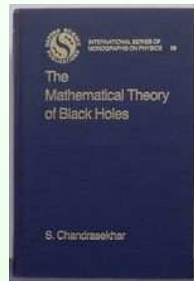
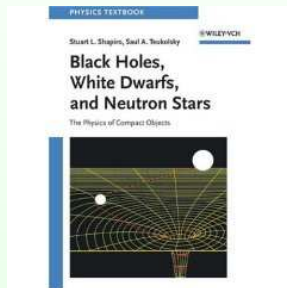
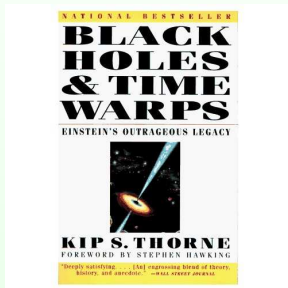
Texts - General Interest



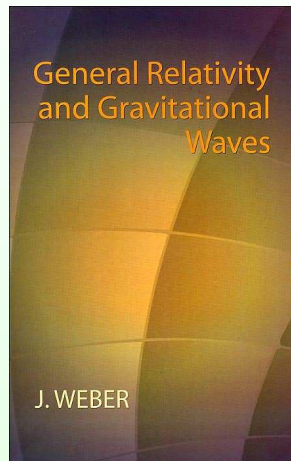
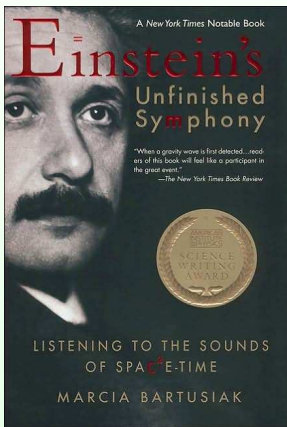
Texts - General Interest



Texts - Black Holes



Texts - Gravitational Waves



Poster Presentation

AAPT - Teaching General Relativity, Syracuse, NY- 07/2006

Lessons on Teaching Undergraduate GR

Lessons on Teaching Undergraduate General Relativity and Differential Geometry Courses

Russell L. Herman and Gabriel Lugo

University of North Carolina Wilmington, Wilmington, NC

Abstract

We describe the course content and lessons learned teaching simultaneously offered courses to undergraduate physics and mathematics majors. A subset of students took both courses. The general relativity course was offered in the physics curriculum and focused more on the physics with standard mathematics prerequisites. The differential geometry course was aimed at the geometry of curves and surfaces ending with a study Cartan's equations and applications to computing curvatures in general relativity.

GR Syllabus

AAPT - Teaching General Relativity, Syracuse, NY- 07/2006

Lessons on Teaching Undergraduate GR

General Relativity

Geometry on a Sphere

Special Relativity

Four Vectors

Dynamics

Principle of Equivalence

Newtonian Gravity

Metrics

Light Cones

Local Inertial Frames

Curved Spacetime

Geodesic Equation

Symmetries and Conservation Laws

Schwarzschild Solution

Gravitational Redshift

Perihelion Shift

Black Holes

Cosmology

Einsteins Equation

Prerequisites: Multivariate Calculus, Classical Mechanics, Modern Physics, Jr-Sr Standing.

DG Syllabus

AAPT - Teaching General Relativity, Syracuse, NY- 07/2006

Lessons on Teaching Undergraduate GR

Differential Geometry

Linear Algebra

Tangent Vectors

Curves

Fundamental Theorem of Curves

Surfaces

Curvature of Curves and Surfaces

1-Forms

Tensors

Higher Rank Forms

Exterior Derivatives

Hodge * Operator

Frames

Curvilinear Coordinates

Covariant Derivatives

Cartan Equations

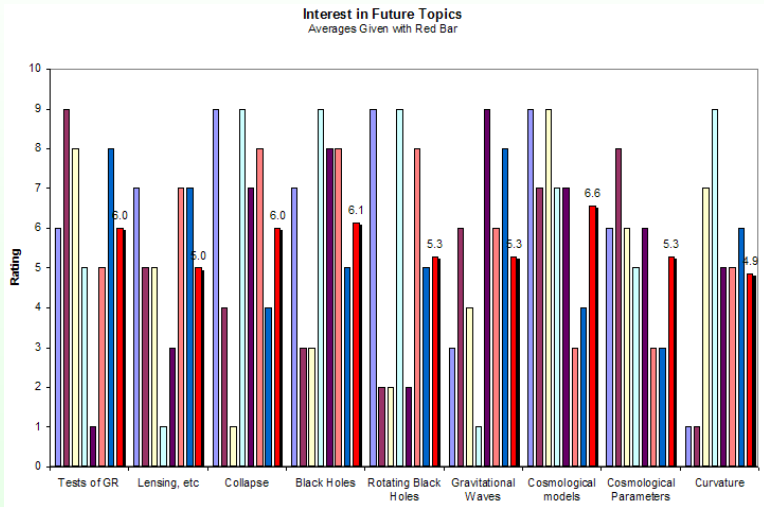
Manifolds

Fundamental Forms

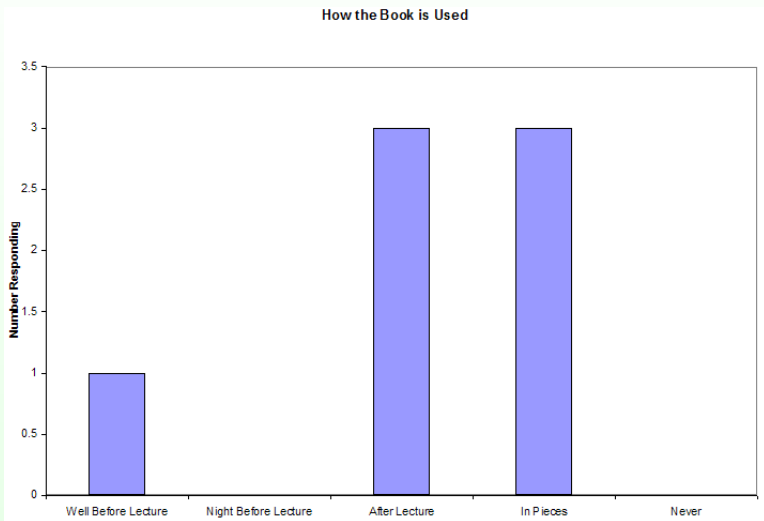
Curvature and Einsteins Equation

Prerequisites: Linear Algebra, Multivariate Calculus, Jr-Sr Standing.

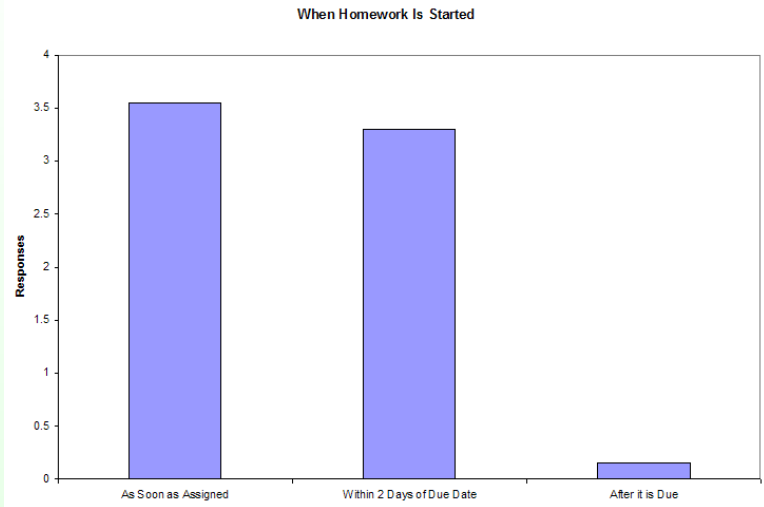
Midterm Survey - Topics Interest



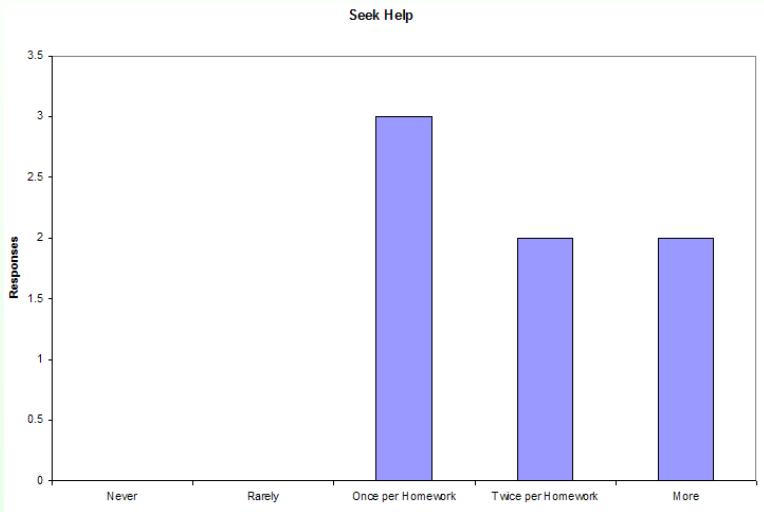
Midterm Survey - When Do Students Read Text?



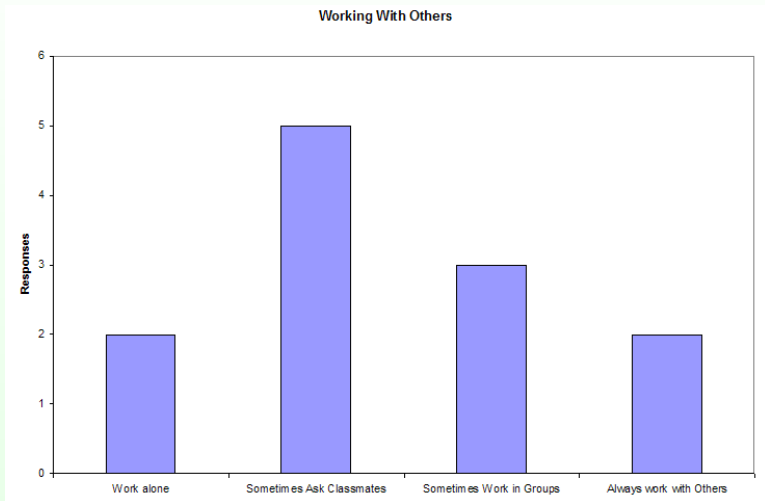
Midterm Survey - When Do They Start Assignments?



Midterm Survey - Do They Seek Help?

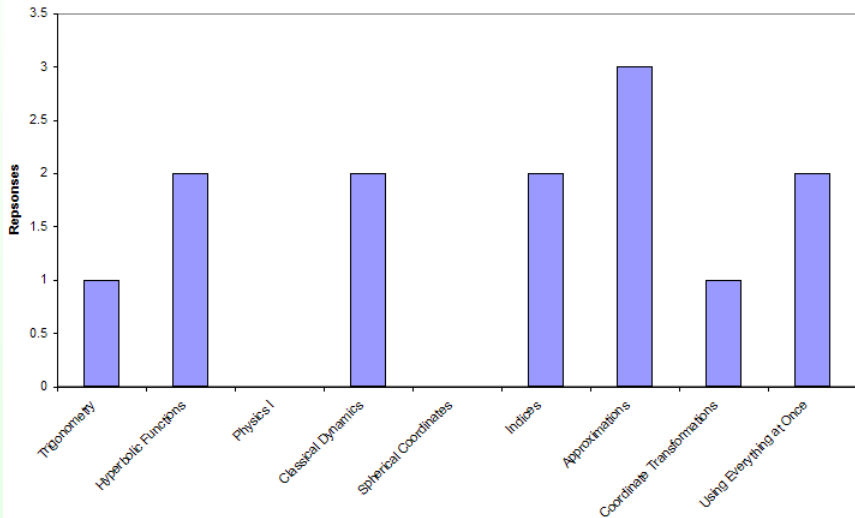


Midterm Survey - Do They Work Together?



Midterm Survey - What Gives Them Difficulty?

Problem Topics



Lessons Learned

AAPT - Teaching General Relativity, Syracuse, NY- 07/2006

Lessons on Teaching Undergraduate GR

Lessons Learned

Undergraduates need

1. more linear algebra emphasizing linear transformations, the spectral theorem and applications
2. more exposure to using approximations based on binomial expansions
3. more geometric insight
4. more exposure to indexed quantities
5. more practice doing homework in physics classes
6. lessons on how to read physics and mathematics texts
7. to learn how to transfer knowledge between courses

Decision Questions

You and Your Department

- How much time is available? 15 wks? 30 wks
- Is this one shot or ongoing?
- What is your background in GR?
- What subtopic is of most interest to you?

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The Students

- What is their background? Major/Non-major? Mechanics? E& M? Astrophysics? Beginning string theory?
- What is Math Background? Advanced Calculus? Differential Geometry?
- What are their motivations? - General interest, astrophysics, gravitational waves, intro to strings, just to know tensors?

Decision Questions

The Content

- Is the course to be a prerequisite?
- What do students need/want to hear? Solving Einstein equation or doing specific application?
- What is the purpose? What do you know well?
- Is there a focus? GWs, BHs, Cosmology, Formalism?

Summary

- 1 TGRU Workshop
 - Web Site
 - Why? Who? What?
- 2 Special Relativity
 - Postulates
 - Spacetime Diagrams
- 3 General Relativity
 - Principles
 - Flat Space
 - Curved Space
- 4 Course Syllabi
 - General Interest
 - Physics First
 - Math Intensive
- 5 Textbooks
- 6 Poster
- 7 Summary