

The First Three Minutes of the Universe

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Mathematics and Statistics
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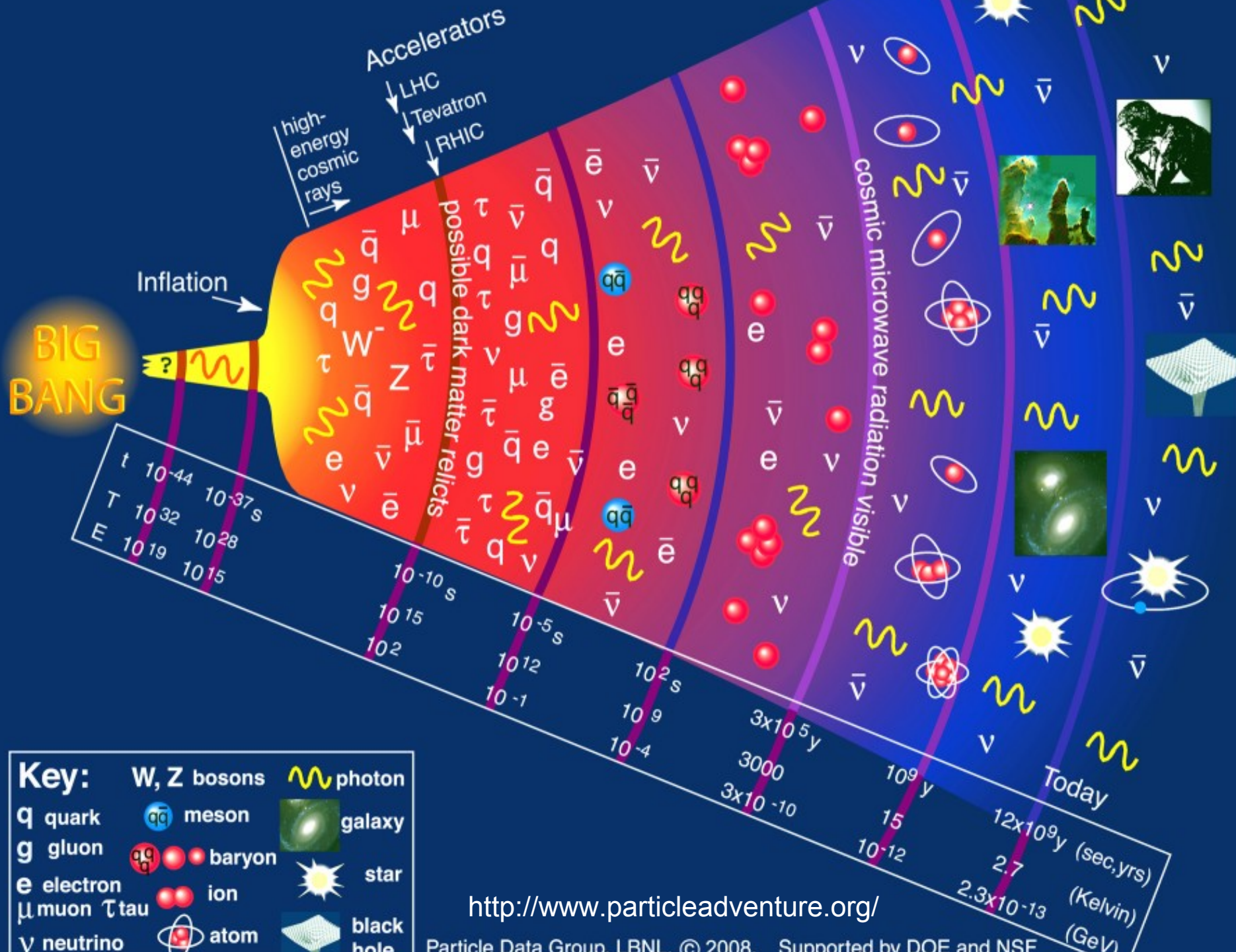
<http://www.dimijianimages.com/More-p20-Madagascar-p7/night-sky-from-Madagascar-gallery.htm>

The First Three Minutes, UNC Wilmington, College Day-2008

Known Universe

- ~ Age - 13.7 billion years old
- ~ Size - >93 billion lights years across
- ~ Density - 9.9×10^{-30} gms/cubic centimeter.
- ~ Appears to consist of
 - 73% dark energy,
 - 23% dark matter
 - 4% ordinary matter.
- ~ Black holes, white dwarfs, galaxies, stars, planets, comets, asteroids, WIMPs, MACHOs, ...

History of the Universe



In the beginning ...

<http://www.spaceandmotion.com/cosmology-history-astronomy-universe-space.htm>

~ **Thales of Miletus (624 BC – 546 BC)**

~ **Pythagoras (585-497 BC)**
Harmony of the Spheres

~ **Socrates (469-399 BC)**

~ **Democritus (460-370 BC)**

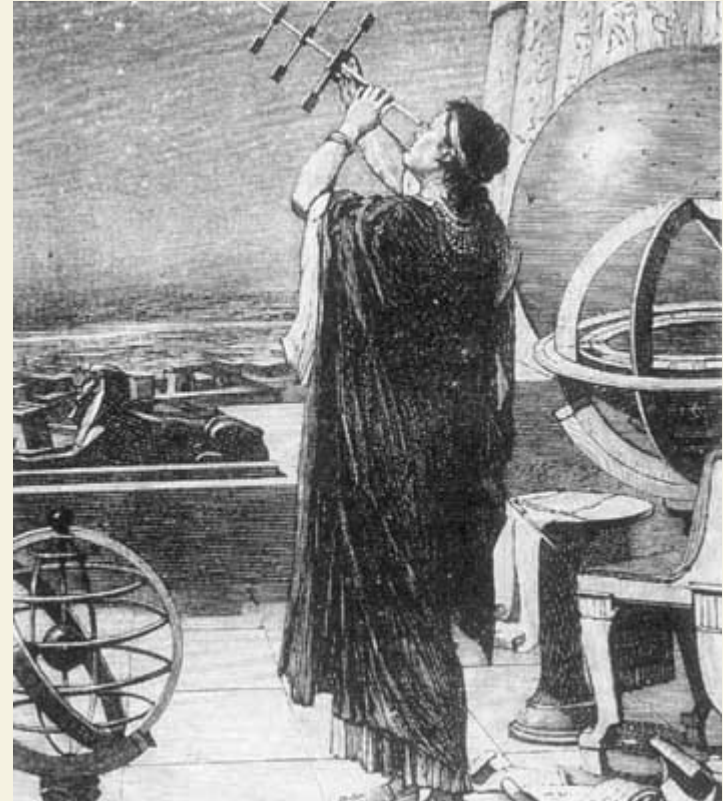
~ **Plato (427-347 BC)**

~ **Aristotle (384-322 BC)**

~ **Archimedes (287-212 BC)**

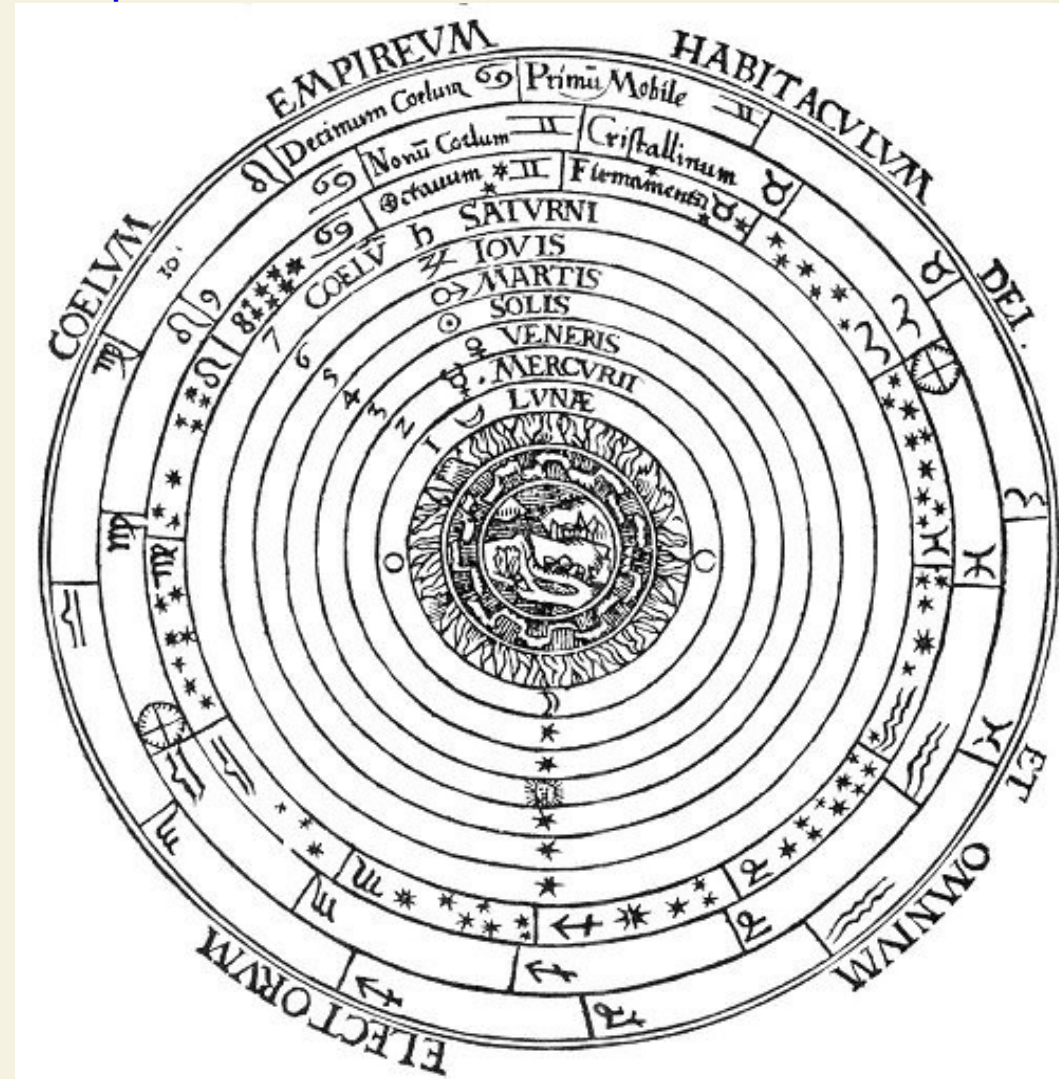
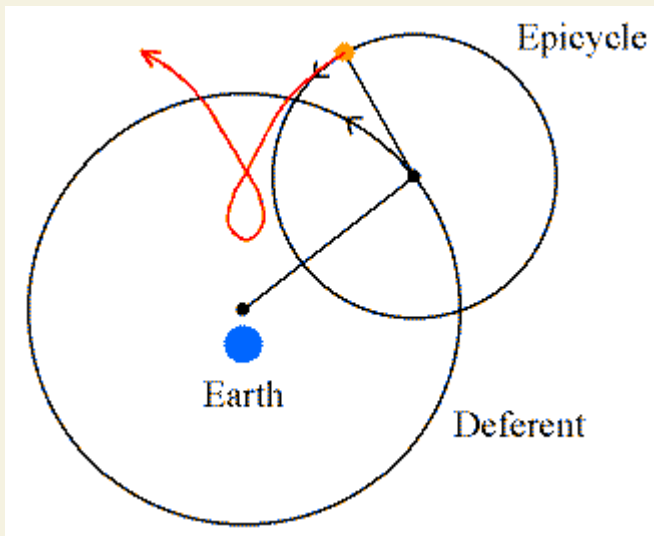
**Earth is fixed and immovable,
stars fixed in sky, planets = wanderers**

~ **Hipparchus (190-120 BC) - seasonal inconsistencies**



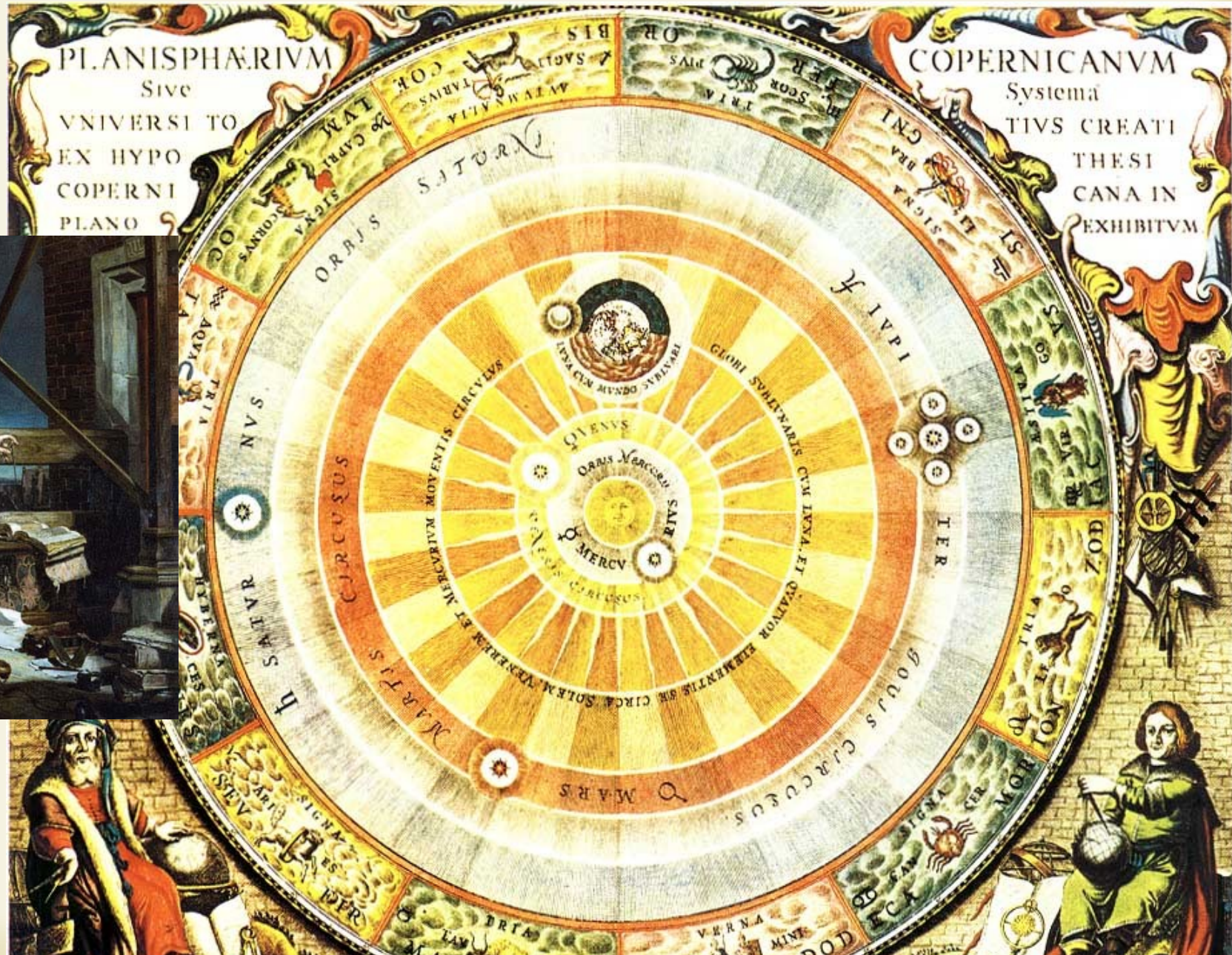
Geocentric System Ptolemy (85-165)

<http://www.sacred-texts.com/eso/sta/sta03.htm>



<http://www.daviddarling.info/images/epicycle.gif>

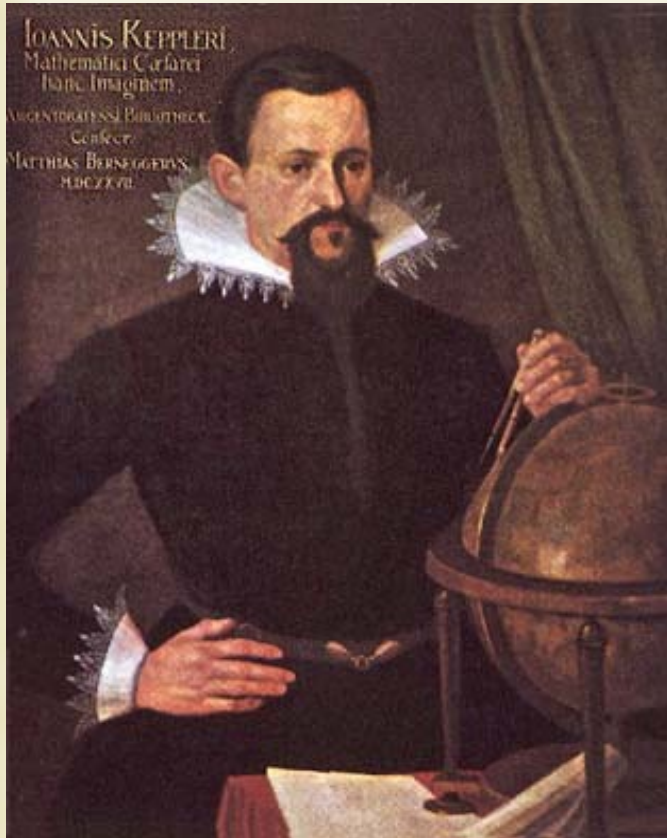
Heliocentric System Copernicus (1473-1543)



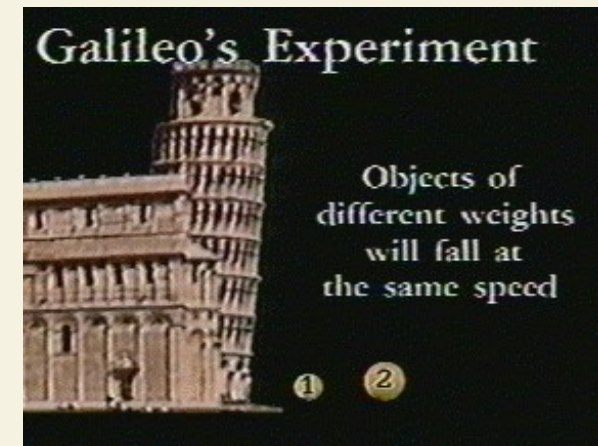
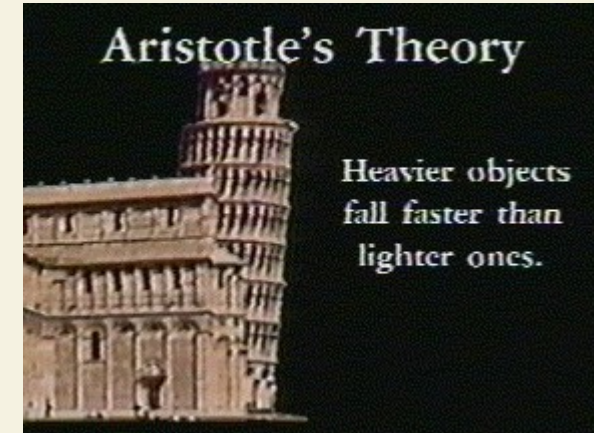
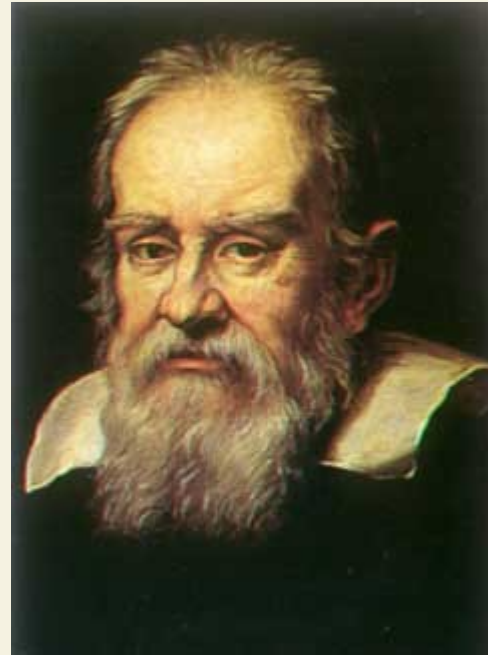
Scientific Deductions

Tycho Brahe
(1546-1601)

Johannes Kepler
(1571-1630)



Galileo Galilei
(1564-1642)



The Clockwork Universe

Sir Isaac Newton (1642-1727)

~ **Principia (1689)**

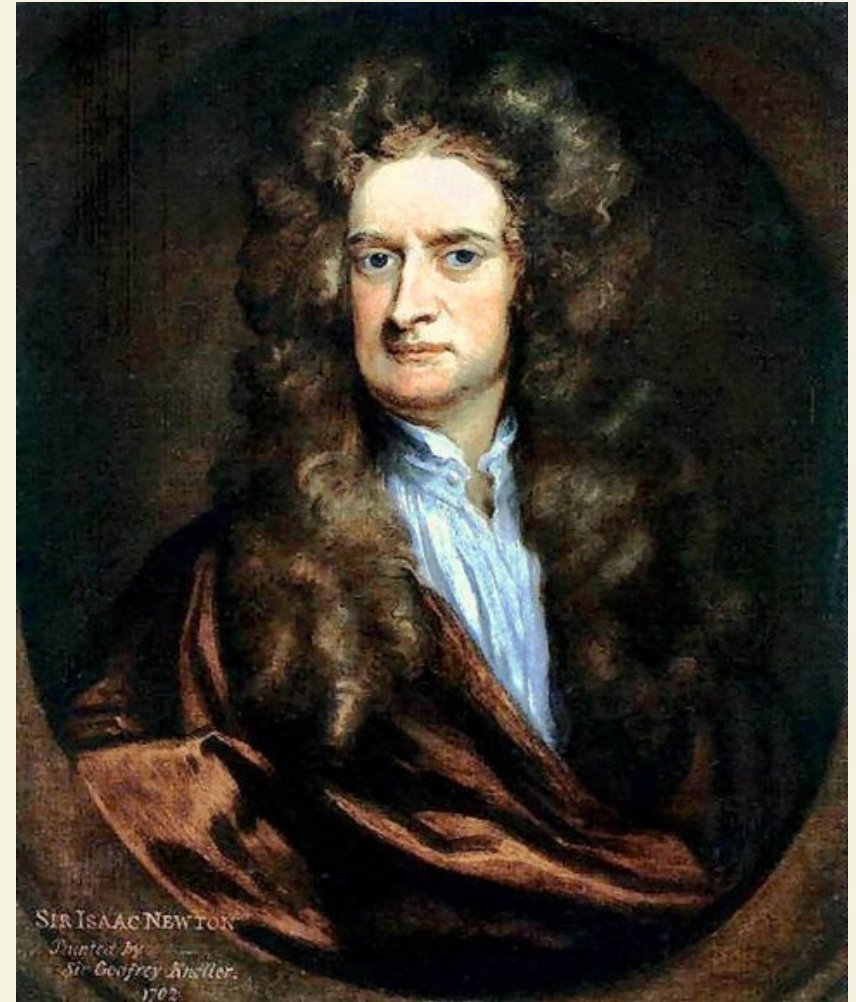
*Philosophiæ Naturalis Principia
Mathematica (Mathematical
Principles of Natural Philosophy)*
(1687)

Laws of Motion

Law of Gravitation

Kepler's Laws Explained

Calculus



Unification

... the force responsible for bodies falling on the Earth is the same as that causing the moon to follow its orbit.

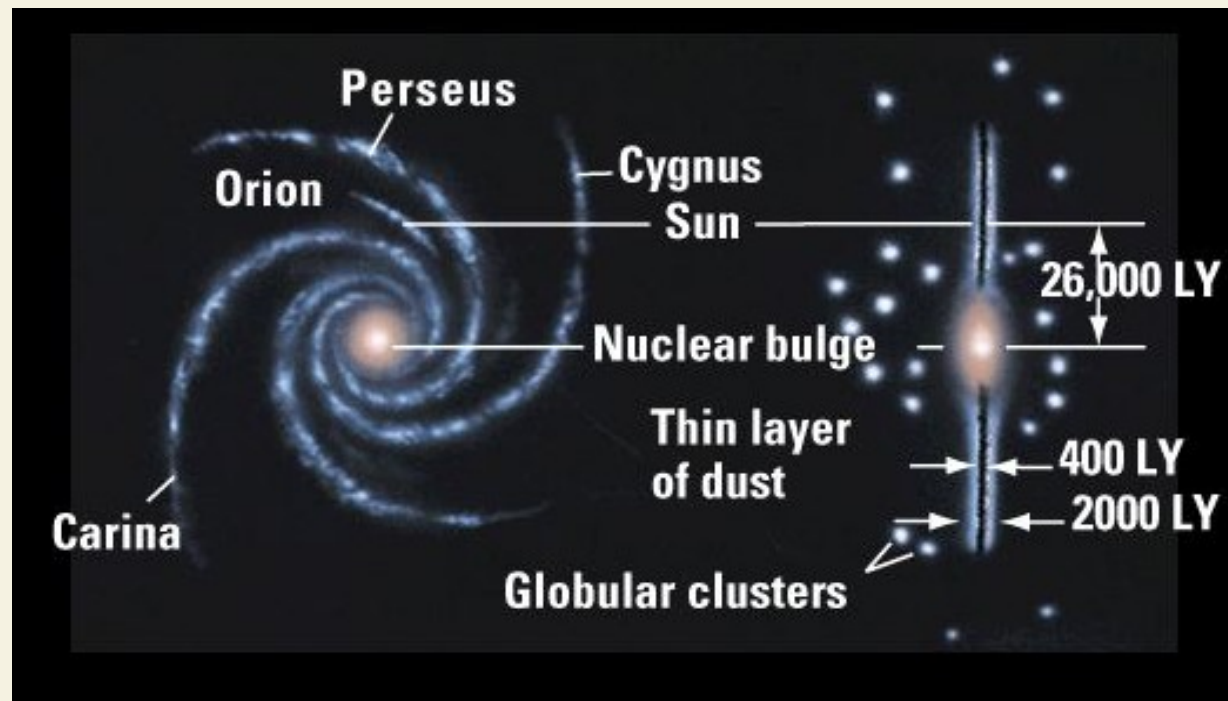
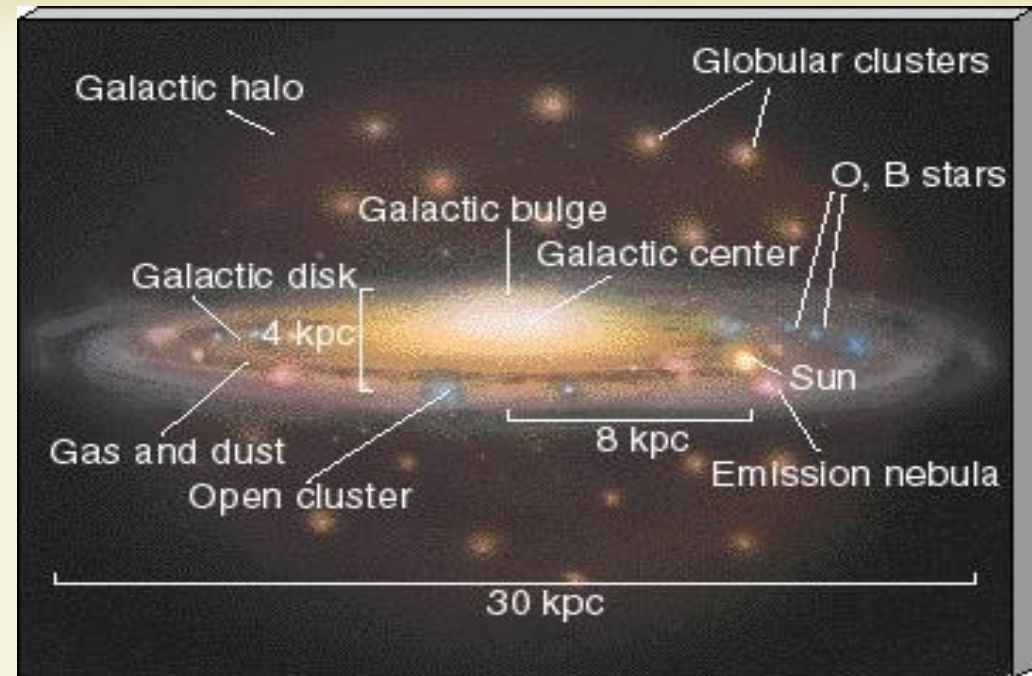
The Milky Way

~ 4×10^{11} stars

~ $> 6 \times 10^{11}$ sun mass

~ $M_{\text{ave}} = 0.3$ sun

How do we know?



Distances & the Milky Way Galaxy

~ Frederick William Herschel (1738-1822)



~ Built telescopes, discovered Uranus

~ Measured stellar distances

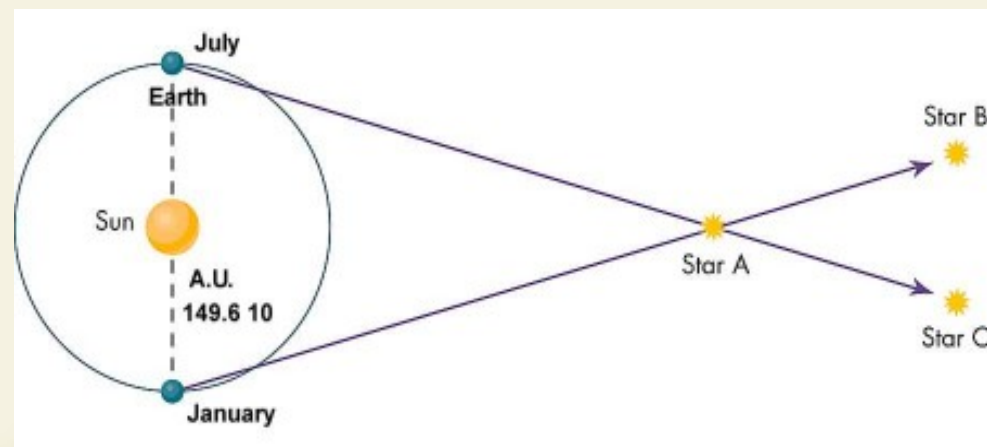
~ Stars distributed in a pancake shape – Milky Way

~ 1000 siriometers x 100 siriometers

~ Asteroids, infrared radiation ...

~ Friedrich Wilhelm Bessel (1784-1846) – 28 years

~ Used parallax to establish stellar distances in km



The Search for Nebulae

~ Charles Messier (1730-1817)

~ Catalog of 103

~ Crab Nebulae M1

~ Andromeda N. M31

Are they in Milky Way or beyond?

~ William/Caroline Hershel

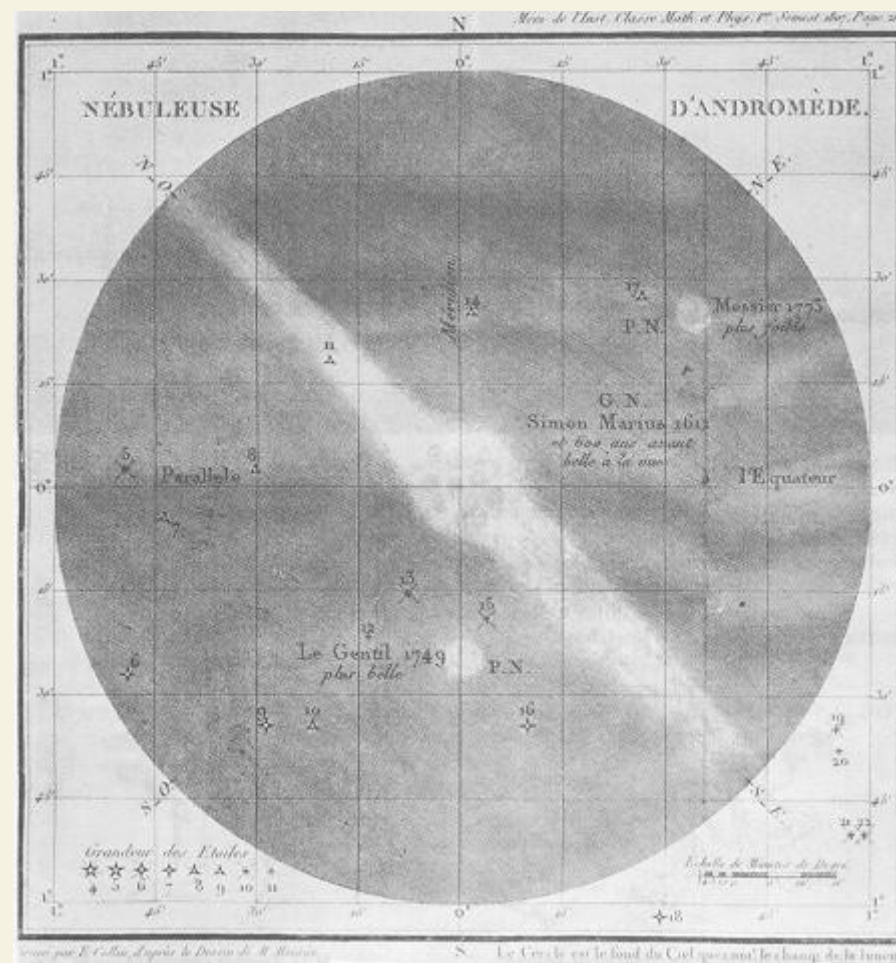
~ Cataloged 2500 nebulae

~ Sited a star in some
– perhaps solar system birth

~ Therefore, in Milky Way

~ Immanuel Kant (1724-1804)

~ Believed nebulae were beyond Milky Way



“The Great Debate”

Shapley-Curtis Debate - 1920

Are distant nebulae relatively small and within our galaxy, or are they large, independent galaxies?

~ Harlow Shapley (1885-1972)

~ Nebulae part of galaxy

~ Sun in outer regions of galaxy

~ Herbert Doust Curtis (1872-1942)

~ Nebulae outside galaxy

~ Sun at center

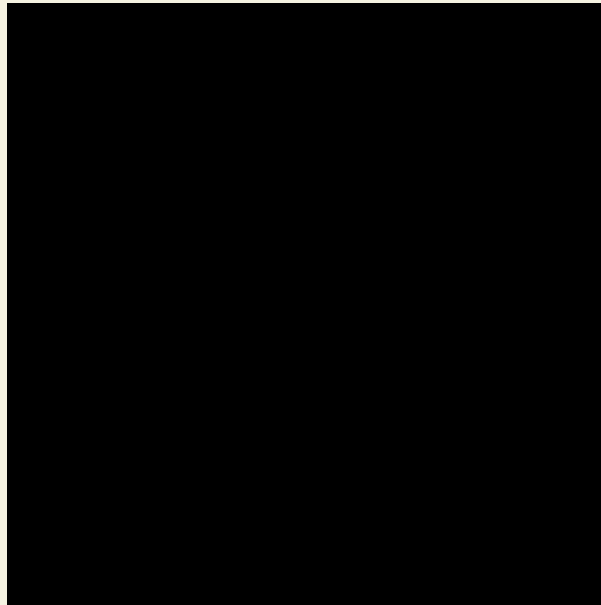
~ Walter Baade (1893-1960) – Milky Way is typical galaxy!



Before leaving the skies

Why is the night sky dark?

Olber's Paradox - 1823



http://en.wikipedia.org/wiki/Olbers'_paradox

Electricity and Magnetism

~ Magnetism

~ Lode stones

~ Compasses

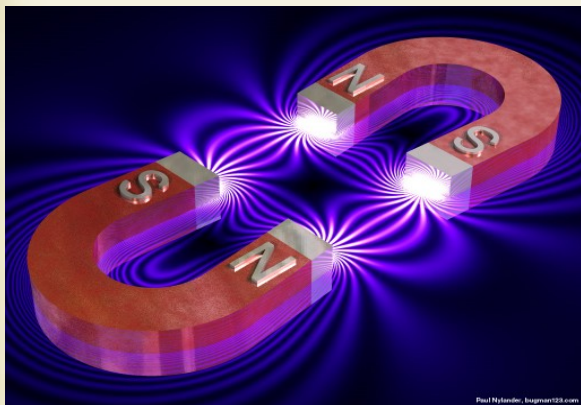
~ William Gilbert (1544-1603)

~ Thomas Browne (1605-1682)

~ Benjamin Franklin (1706-1790)

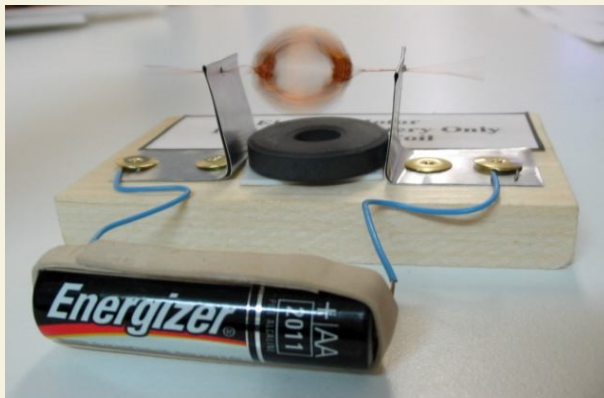
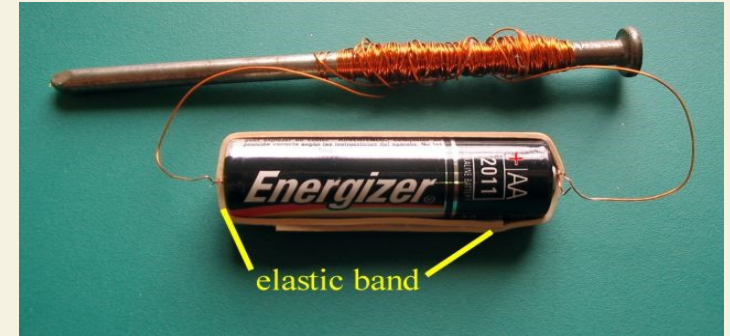
~ Luigi Galvani (1737-1798)

~ Alessandro Volta (1745-1827)



Electromagnetism

- ~ André-Marie Ampère, (1775 - 1836)
- ~ Hans Oersted (1777-1851)
current deflects compass needles
- ~ Georg Simon Ohm (1789-1854)
- ~ Joseph Henry (1797-1878)
electromagnetic induction, first motor,
telegraph
- ~ Michael Faraday (1791-1867)
electrolysis, motors, induction coils, ...



Electromagnetic Waves

~ James Clerk Maxwell (1831-1879)

- Theory of electromagnetism.
- Predicted the electromagnetic waves.
- Electromagnetic waves travel at

$$c = 299,792,458 \text{ m/s} = 186,000 \text{ mi/s}$$

~ Heinrich Hertz (1857-1894)

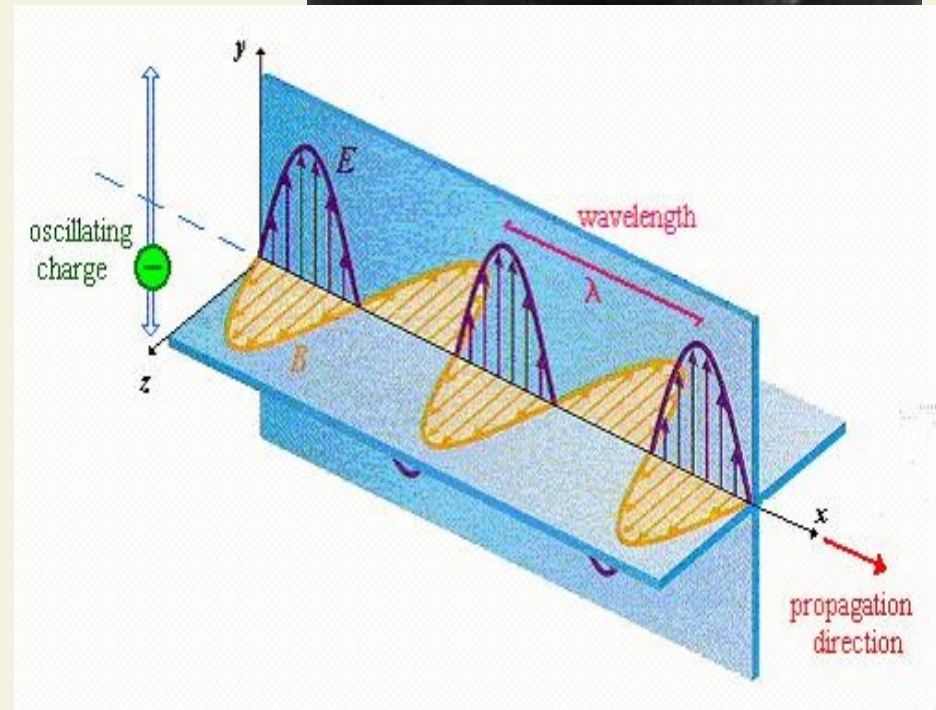
- sent the first radio waves (1888)

What is the medium?

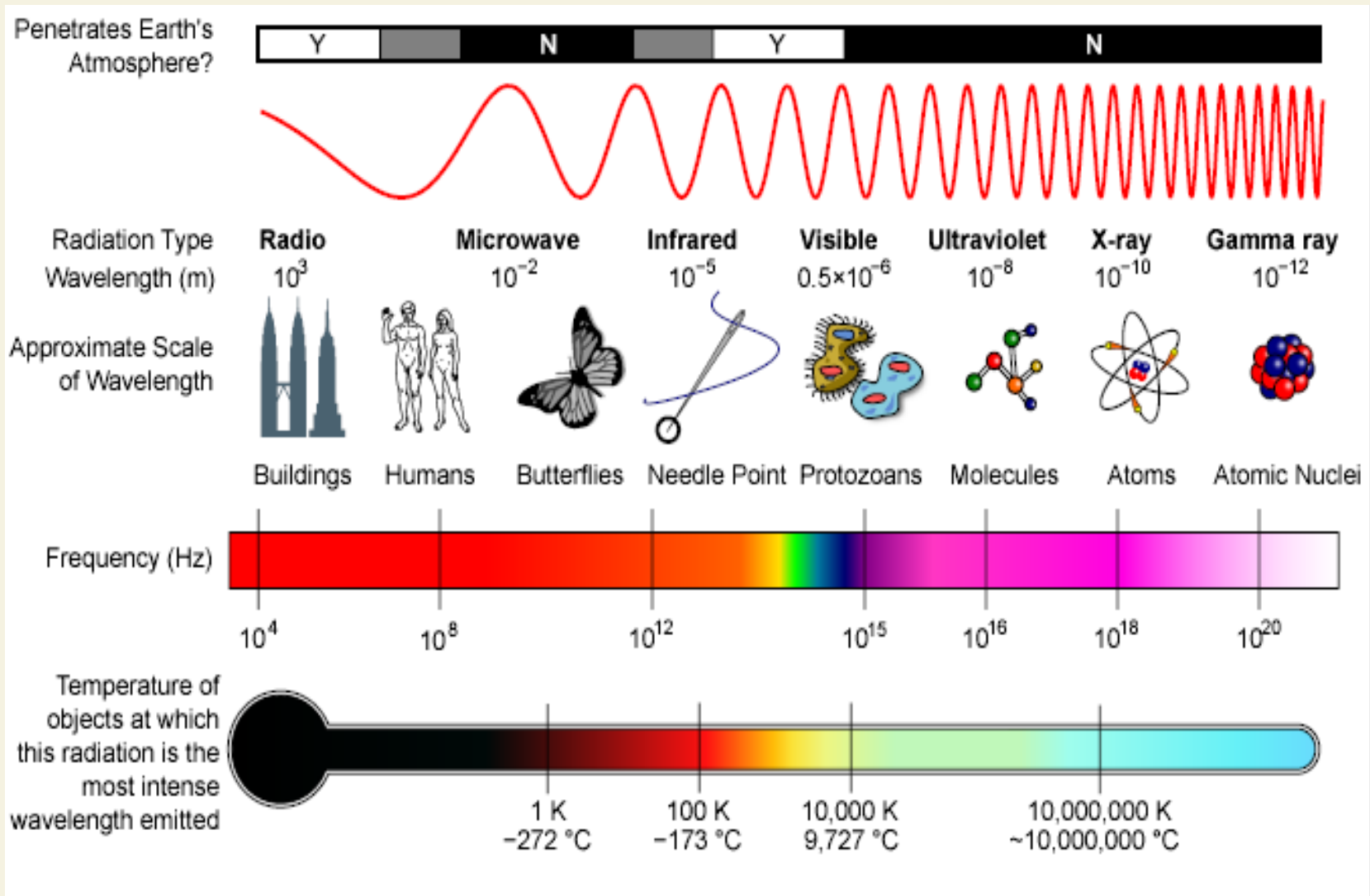
Luminiferous Aether

~ Michelson-Morley (1887)

- could not detect it.



EM Spectrum



http://en.wikipedia.org/wiki/Electromagnetic_spectrum

Spectroscopy

Ionized gas gives off radiation

~ **Johann Balmer 1885**

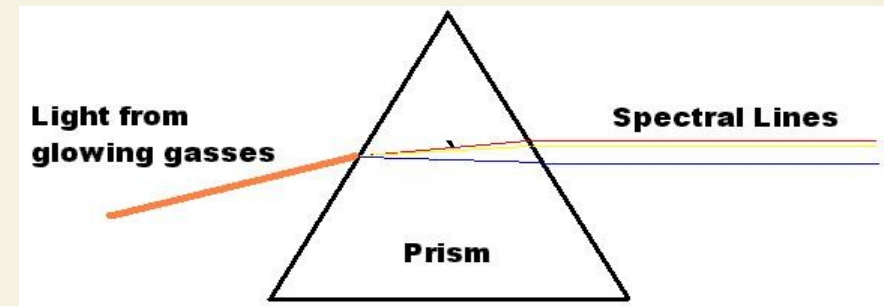
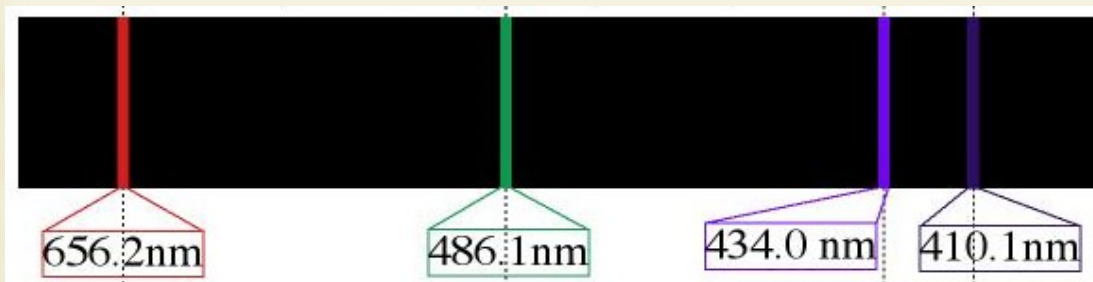
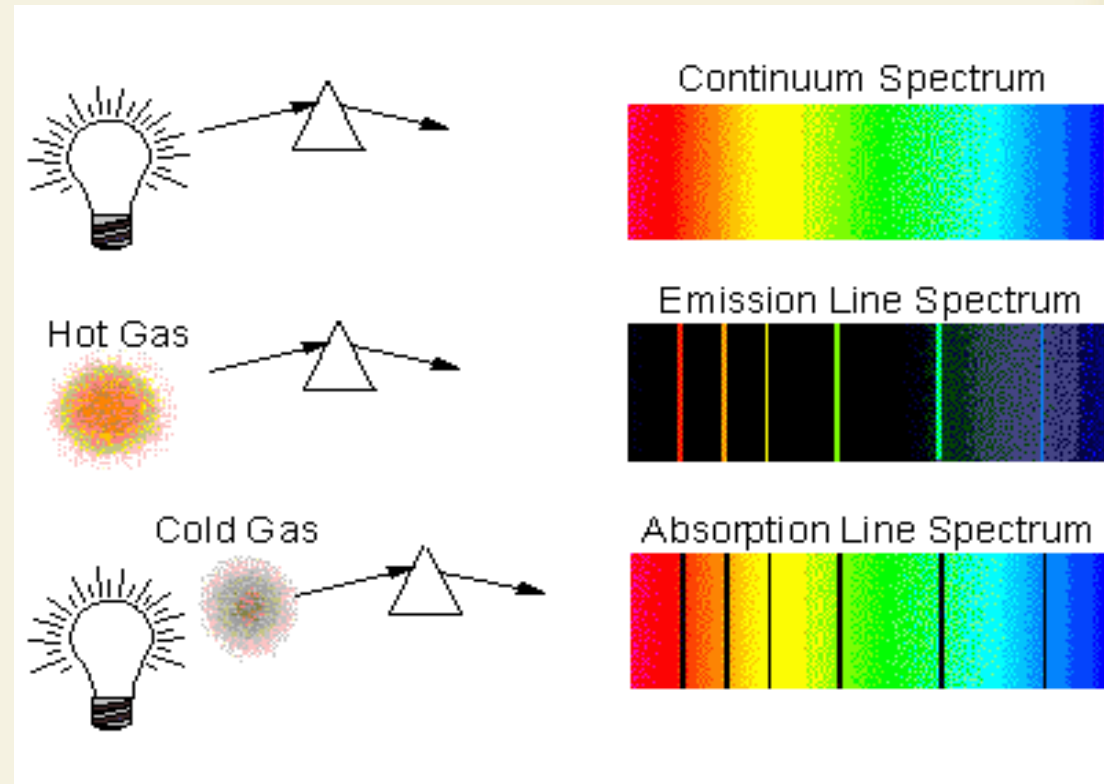
Spectral Lines: Hydrogen
410, 434, 486, 656 nm

Empirical Formula:

$$\lambda = R \left(\frac{1}{4} - \frac{1}{n^2} \right)$$

Predicted 5th-7th lines

~ **Lyman and Paschen Series**



Laws of Thermodynamics

~ James Joule (1818-1889)

Mechanical Equivalent of Heat

~ Engines: Watt, Carnot, Kelvin, Clausius, Carnot

~ Laws of Thermodynamics

1. Adding heat energy or doing work on a body increases internal energy.
2. A body will not spontaneously get hotter.

~ Joseph Stefan (1835-1893) and Ludwig Boltzmann (1844-1906)

Heated bodies Radiate - Stefan-Boltzmann Law

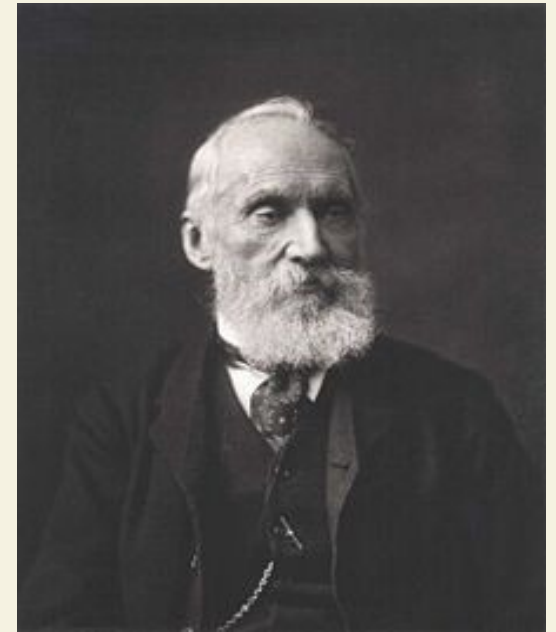
Radiation from blackbody proportional to T^4 .

Physics Revolutions

William Thomson, (1824 – 1907)

1st Baron Kelvin (**Lord Kelvin**)

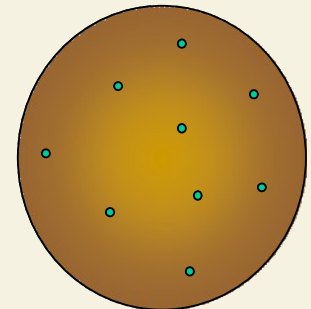
“There is nothing new to be discovered in physics now. All that remains is more and more precise measurement.” - 1900



1895 **Wilhelm Röntgen** discovers X-rays.

1898 **Marie and Pierre Curie** separate radioactive elements.

1898 **Joseph Thomson** measures electron, **“plum-pudding” model of the atom**
- a slightly positive sphere with small, raisin-like negative electrons.

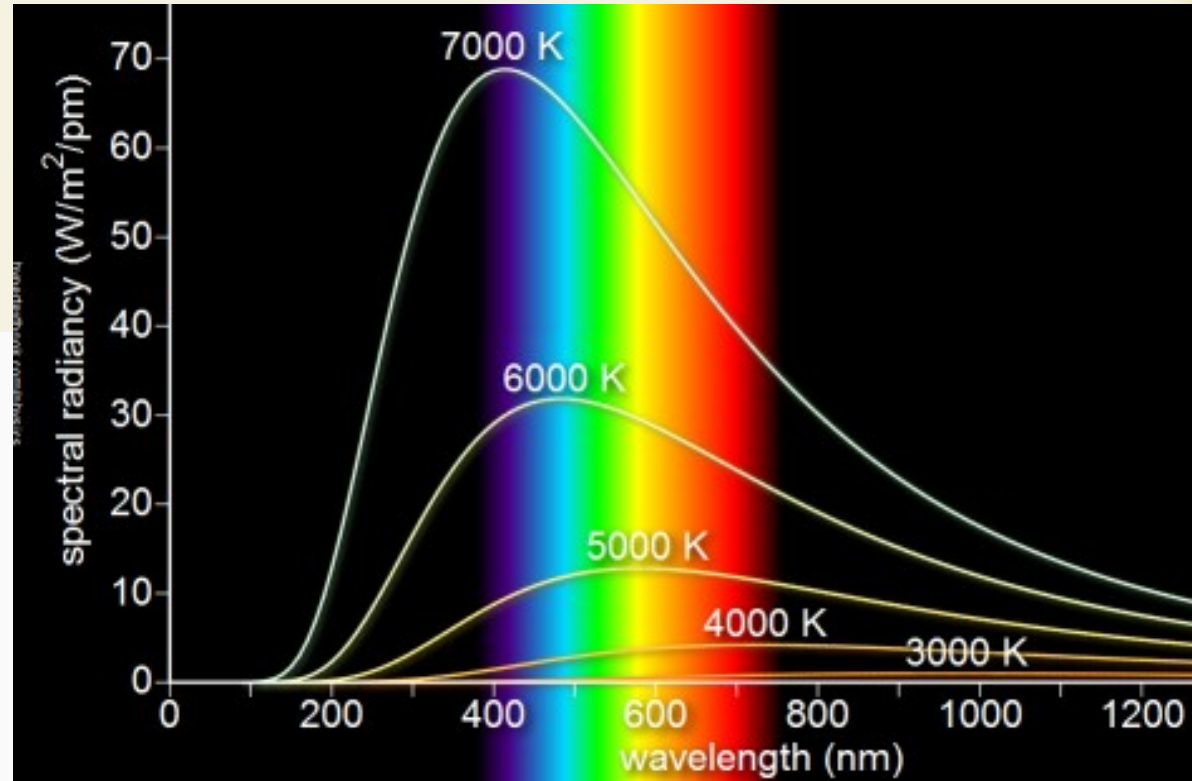
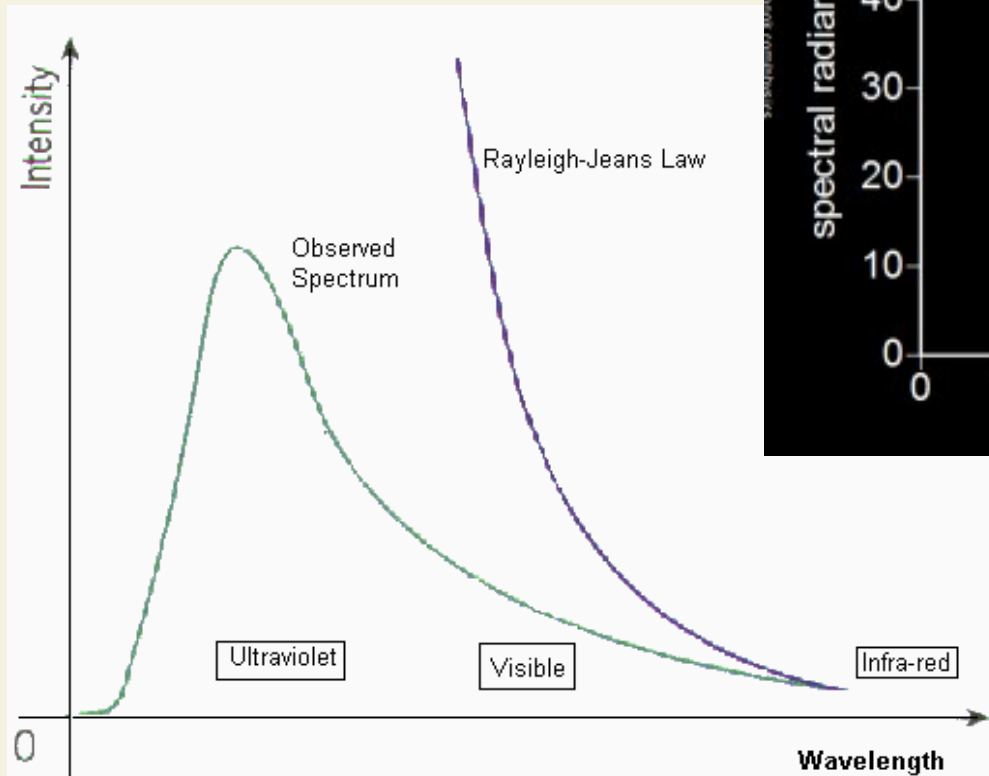


Blackbody Spectrum

Blackbody - a theoretical object that absorbs 100% of the radiation that hits it.

Wien's Law (1896)

Rayleigh - Jeans Law (1900)



Ultraviolet Catastrophe "... when you turn on your toaster, you are instantly fried by a massive gamma ray burst, since your little blackbody toaster should emit infinite energy at the shortest wavelengths."

Quantum Theory

Max Planck

(Karl Ernst Ludwig Marx Planck 1858-1947)

oscillators can only vibrate at discrete frequencies:

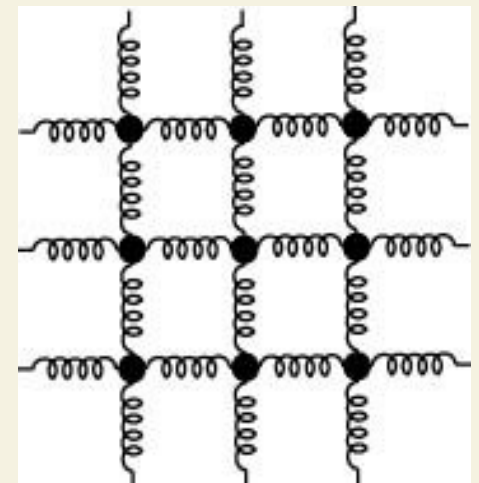
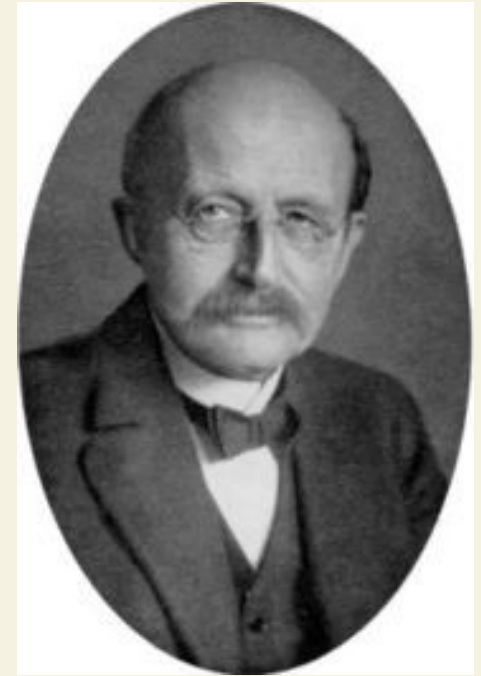
$$E_n = n(hf), n = 1, 2, 3 \dots$$

Thus, the energy difference

$$\Delta E = hf,$$

where Planck's constant is given by

$$h = 6.63 \times 10^{-34} \text{ Js}$$



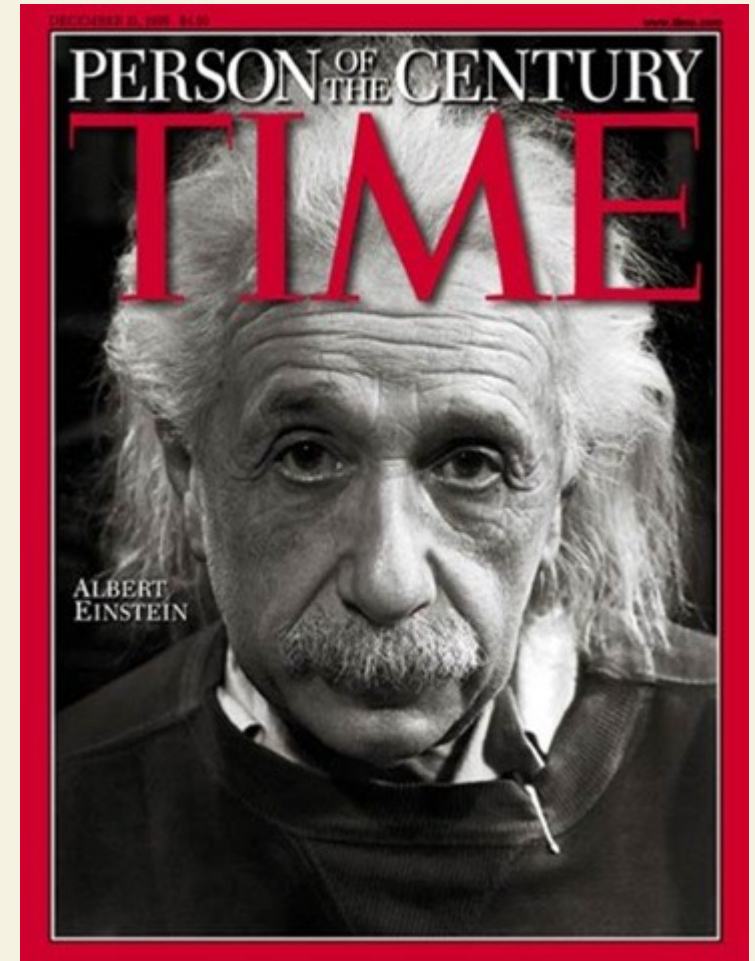
Albert Einstein(1879-1955)

Einstein's 1905 Papers

- **March – Photoelectric Effect**
- **May – Brownian Motion**
- **June – Theory of Relativity**
- **September – $E = mc^2$**

These papers lead to revolutions in physics that defined physics research for the entire century:

- **Confirming molecular theory.**
- **Questioning how we view space and time.**
- **Unifying electromagnetic theory with mechanics.**
- **Introducing wave-particle duality.**



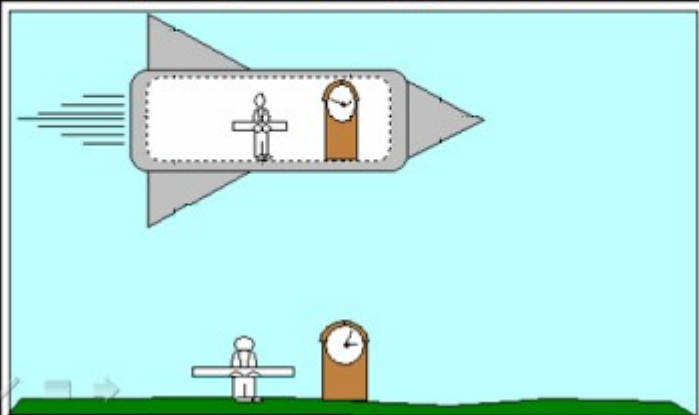
Relativity and Quantum Theory

Theory of Relativity

Physics looks the same to all observers moving at a constant velocity

The speed of light in a vacuum is the same for all observers

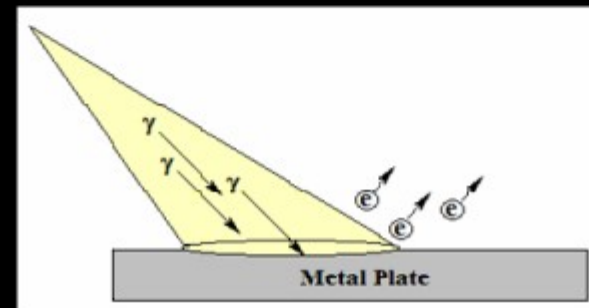
- There is no absolute time or position
- Moving clocks tick slower.
- Moving objects appear shorter.



Photoelectric Effect

Light can cause currents

- Electrons can be ejected from irradiated metal plates.
- Light can act like either particles (quanta) or waves.
- Extended Planck's ideas of energy quantization.
- Lead to explanation of electromagnetic spectra,
- Lead to the development of lasers, transistors and other applications.

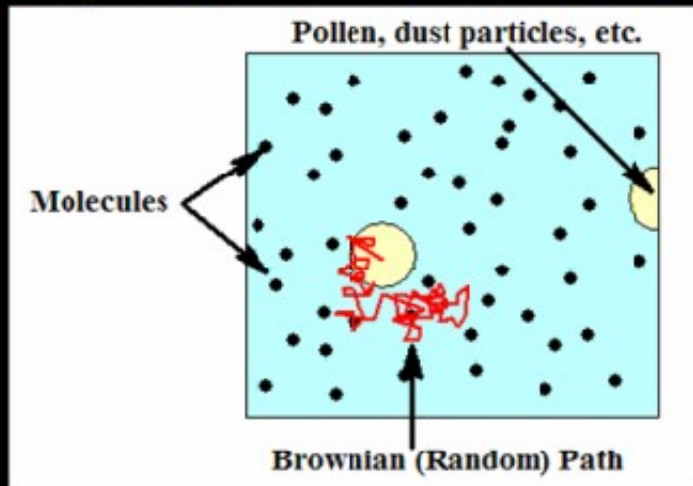


Other Papers

Brownian Motion

the random movement of particles suspended in a fluid

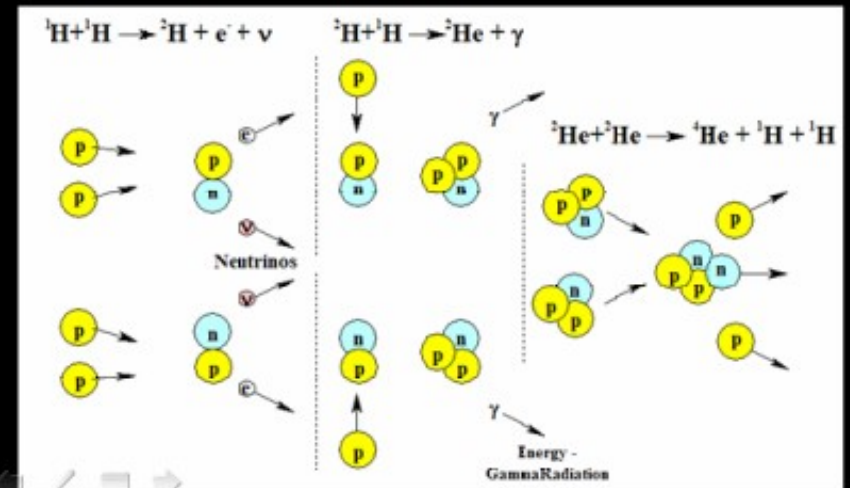
- Explained the observations credited to Robert Brown, 1827
- Predicted molecular motion and size through the effects of collisions with larger particles
- Einstein's work led to an acceptance of molecular theory

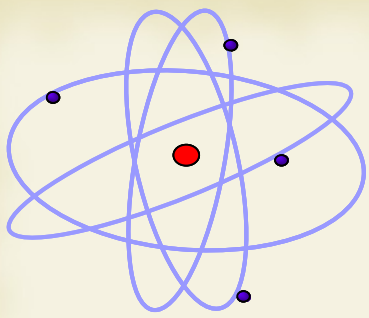


$E = mc^2$

Mass and Energy are different aspects of the same thing

- A consequence of special relativity
- Small bits of matter lead to large energy releases
- Lead to the atomic bomb
- Hydrogen Fusion in Sun:



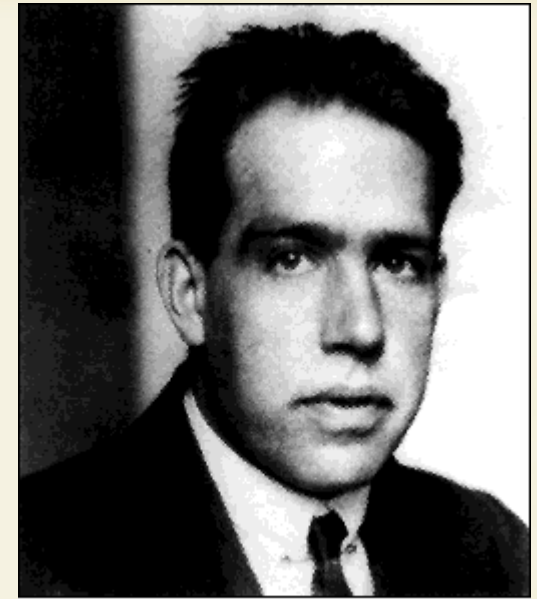


Bohr's Atom - 1913

Niels Bohr (1885-1962)

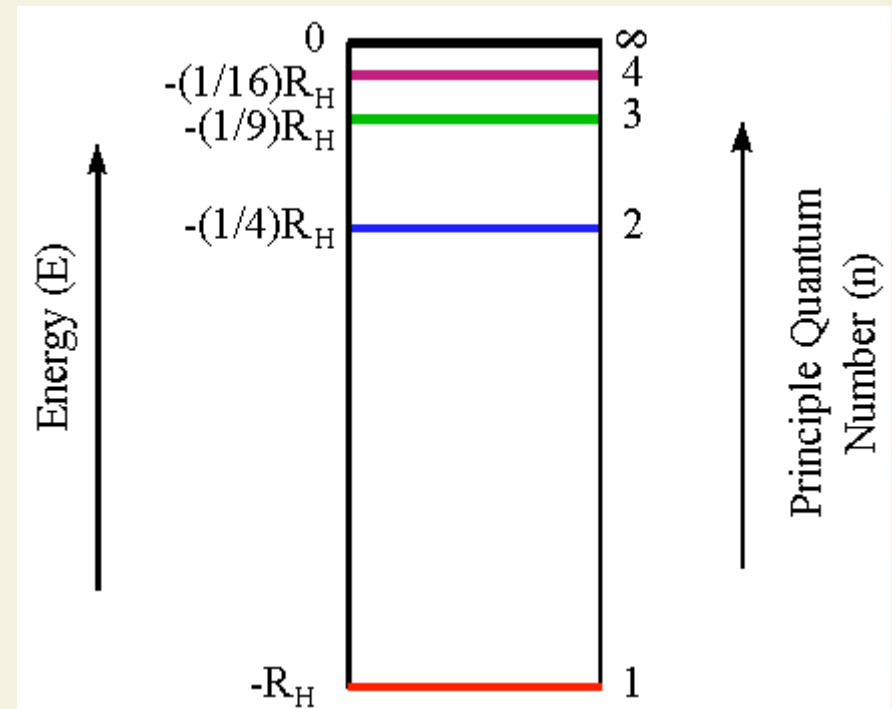
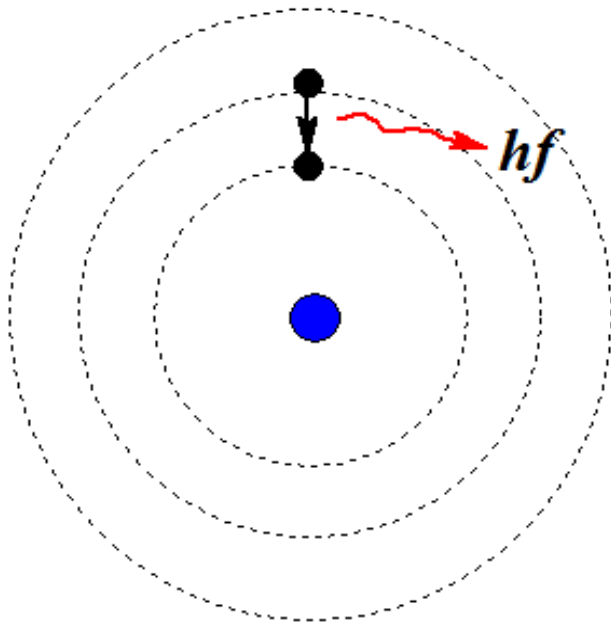
Electrons move in specific orbits.

Accelerating electrons radiate at specific energies.



Niels Bohr

$$E_n = R_H(1/m^2 - 1/n^2)$$



Early History - Quantum Mechanics

1900 - Planck Explains Blackbody Radiation

1905 - Einstein - the Photoelectric Effect, Photons

1913 - The Bohr Model for Hydrogen

1916 - Confirmation of photon, Millikan

1923 - Compton Effect – X-Ray Scattering

1924 – de Broglie - Particles Behave Like Waves

1925 - Matrix Mechanics - Heisenberg

1926 - Derivation of Planck's Law – Dirac

1926 – Wave Mechanics - Schrödinger

1927 - The Uncertainty Principle - Heisenberg

1927 - Davisson-Germer Verified deBroglie's idea

1928 - Relativistic Quantum Mechanics - Dirac



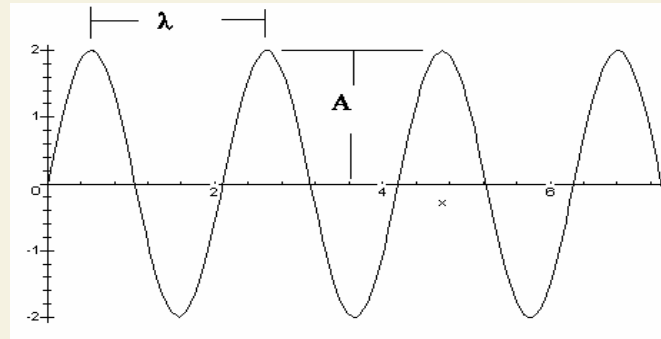
What are Waves?

Characteristics

Wavelength

Frequency

Wavespeed

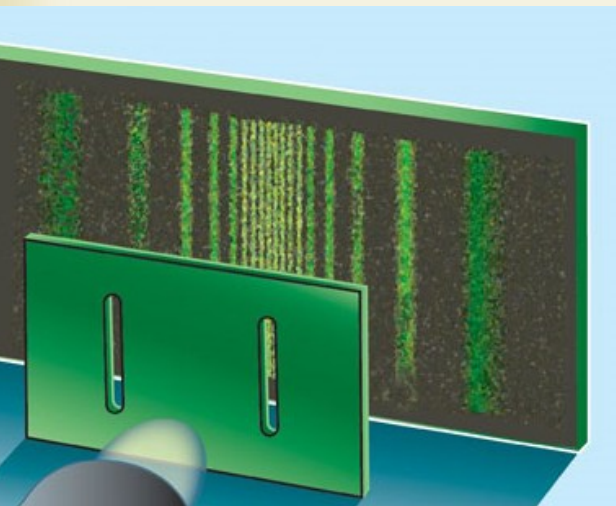
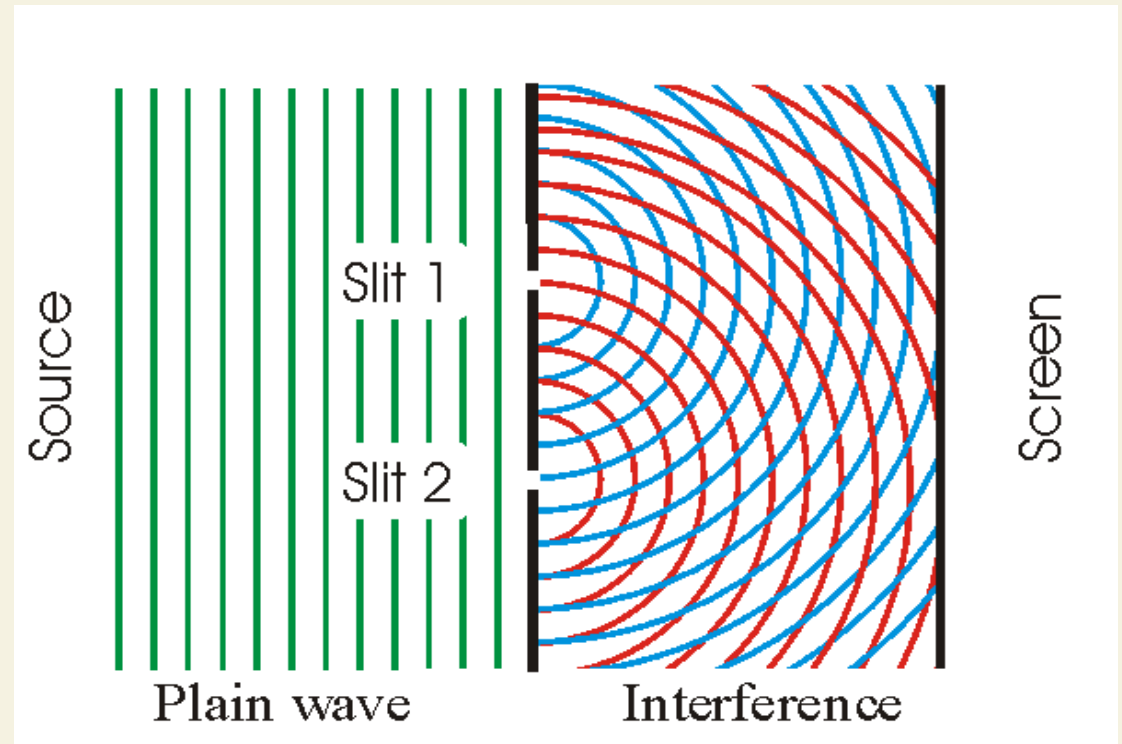


Behavior

Superposition

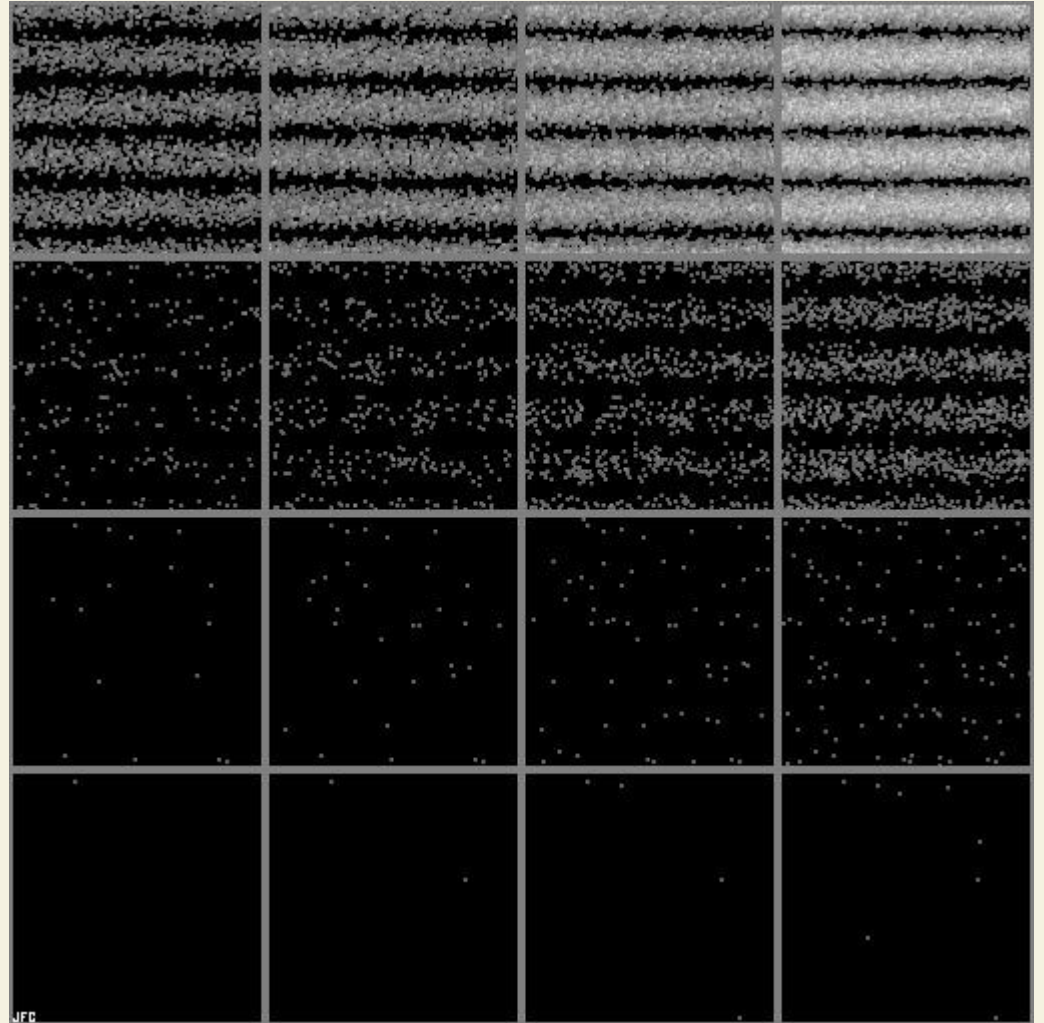
Interference

Diffraction



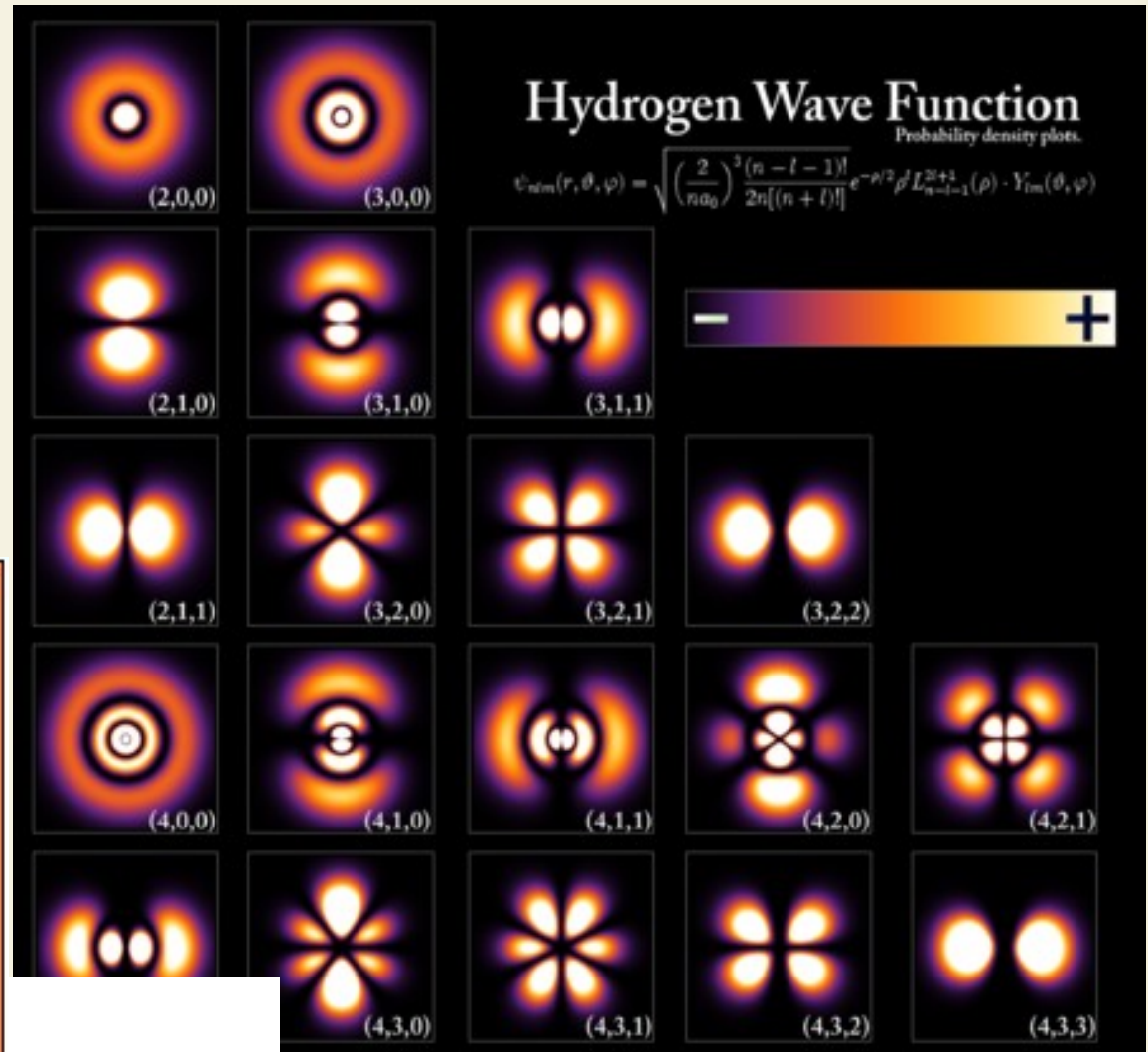
Electron Diffraction - 1929

- ~ George Paget Thomson
(1892-1975)
- ~ Clinton Joseph Davisson
(1881-1958)
- ~ Lester Halbert Germer
(1896-1971)



Erwin Schrödinger (1887-1961)

- ~ Matter Waves
- ~ Predicted Hydrogen Spectrum
- ~ Needed Interpretation
- ~ Leads to ...



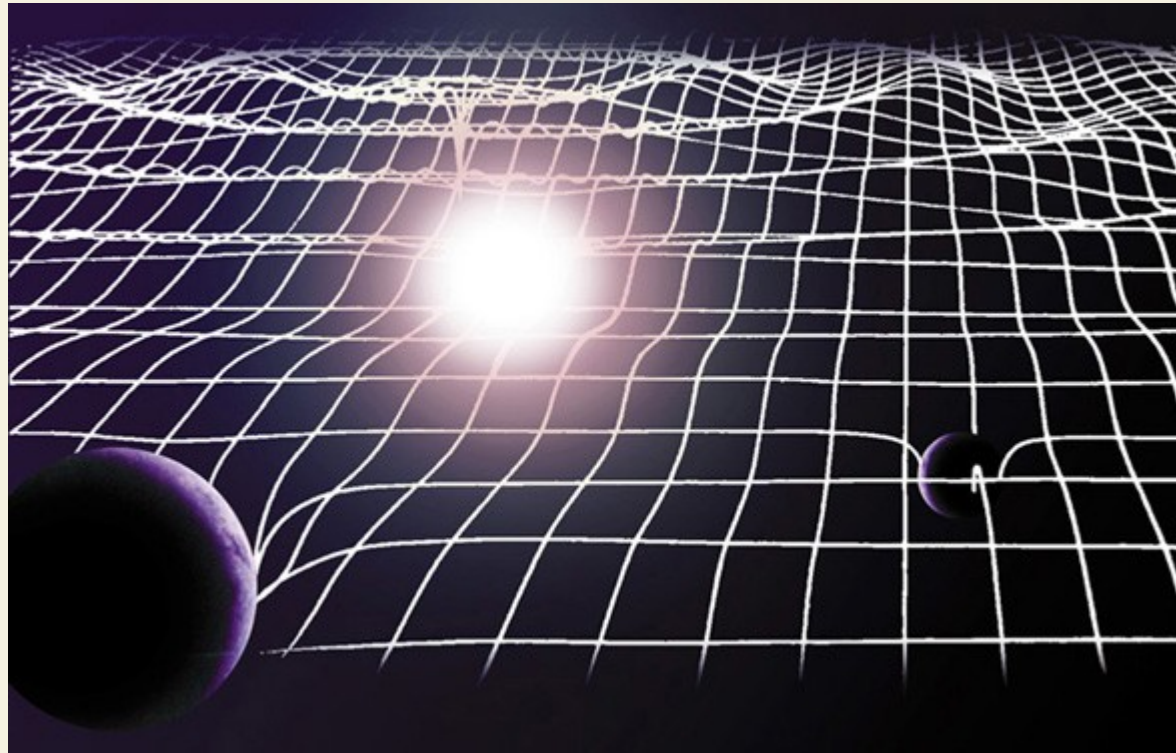
$$i\hbar \frac{\partial}{\partial t} \Psi(x,t) = \left[-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} + V(x) \right] \Psi(x,t) = E\Psi(x)$$

Time evolution Time independent equation

Paradigm Shifts

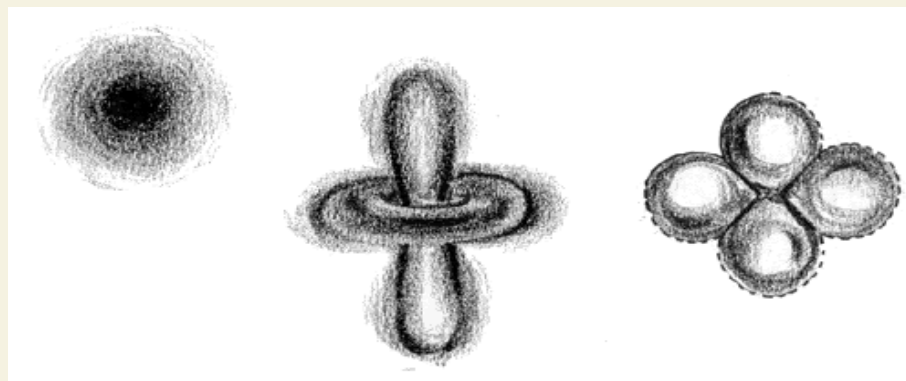
Relativity

Space and Time not absolute, not Euclidean



Quantum Mechanics

Loss of Determinism

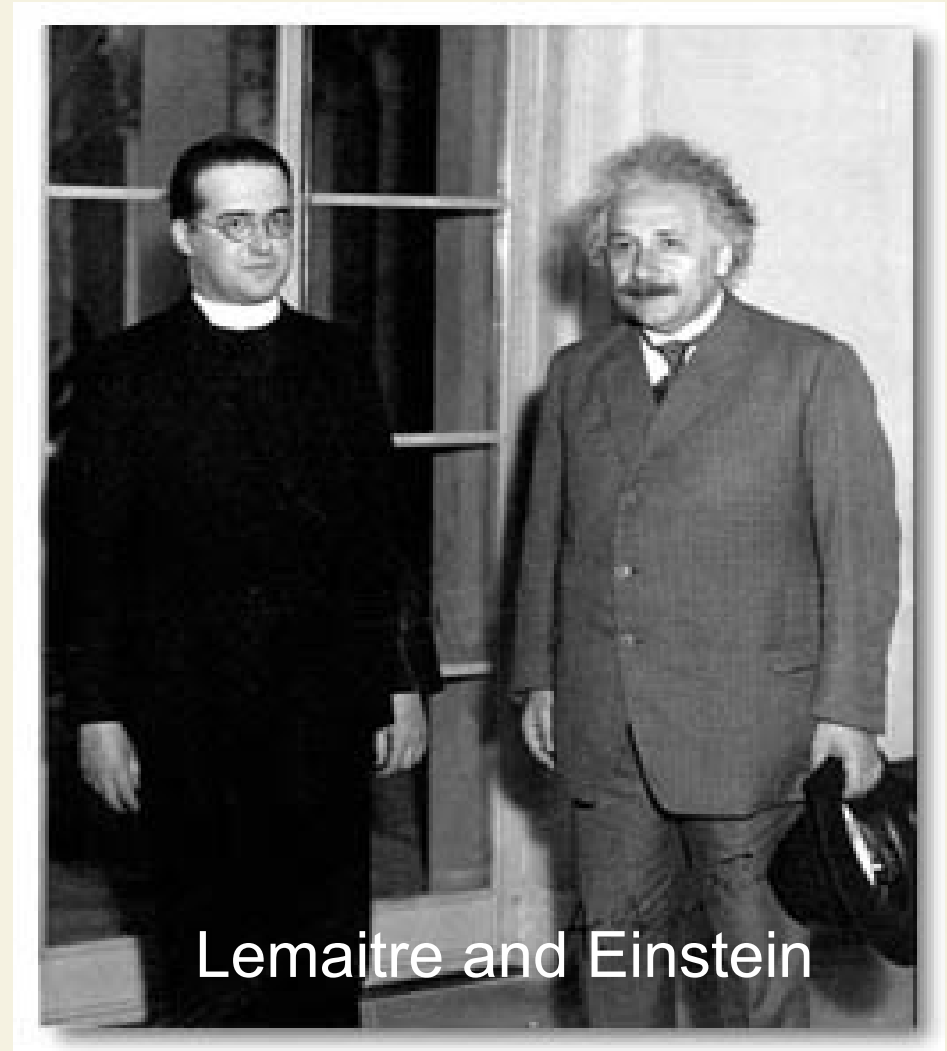


The First Three Minutes, UNC Wilmington, College Day-2008

32

Pre-Modern Cosmology

- 1915 General Relativity
- 1916 Schwarzschild
- 1917 de Sitter
- 1922 Friedman
- 1927 Lemaitre
- 1929 Hubble
- 1932 Einstein-de Sitter
- 1948 Gamow - CMB
- 1950 Hoyle – Steady State
- 1965 Penzias and Wilson



Lemaitre and Einstein

<http://www.catholiceducation.org/articles/science/sc0022.html>

General Relativity

~ Einstein - 1915

Newton's gravitational attraction replaced

Curvature of spacetime tells bodies how to move

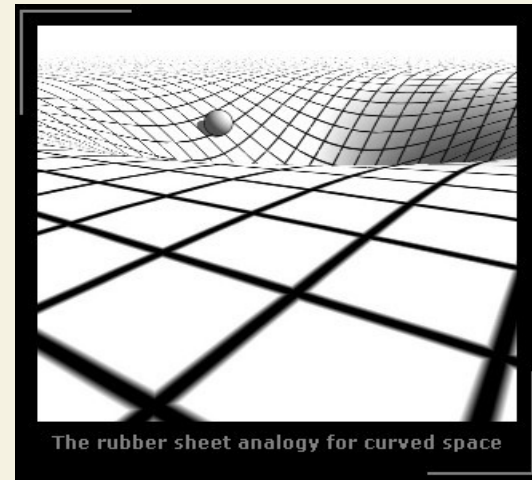
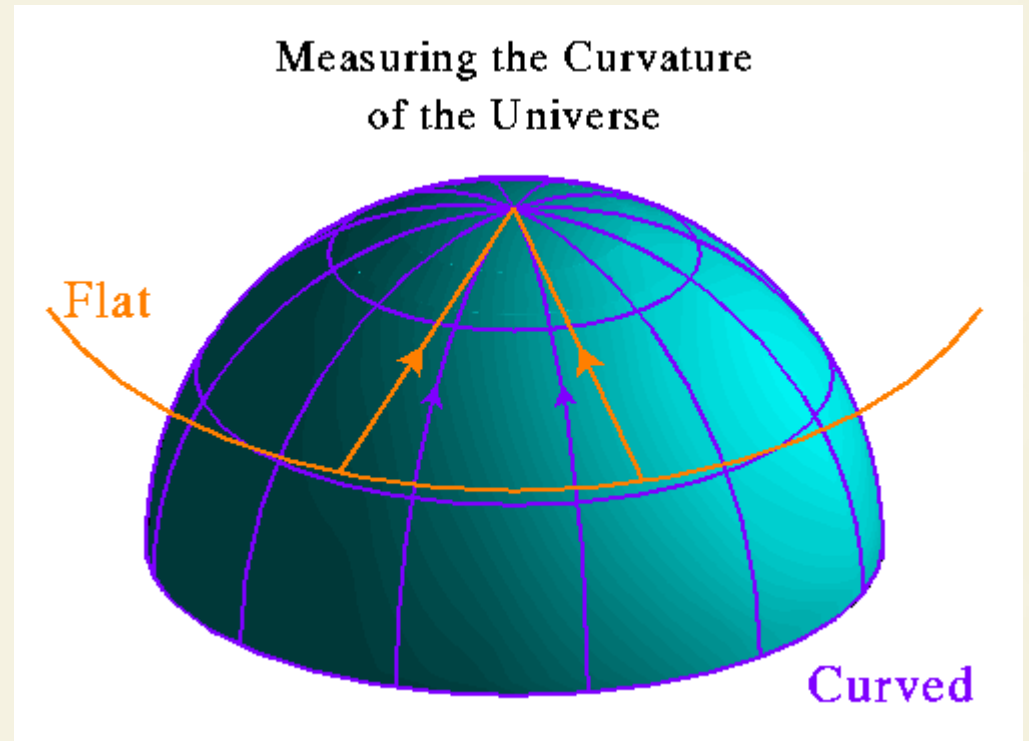
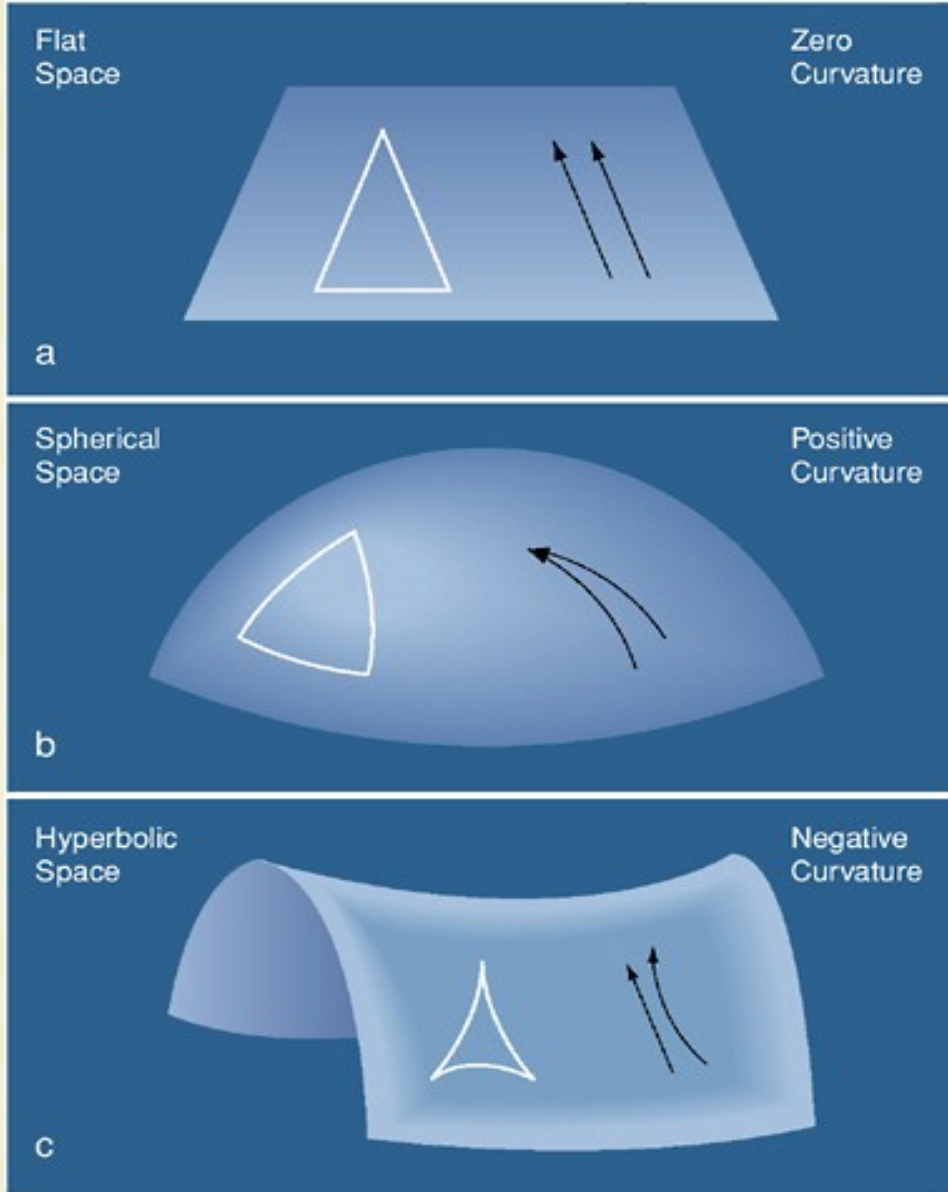
Bodies tell spacetime how to curve

~ Karl Schwarzschild (1873-1916)

Papers on spherical solution sent from WWI front

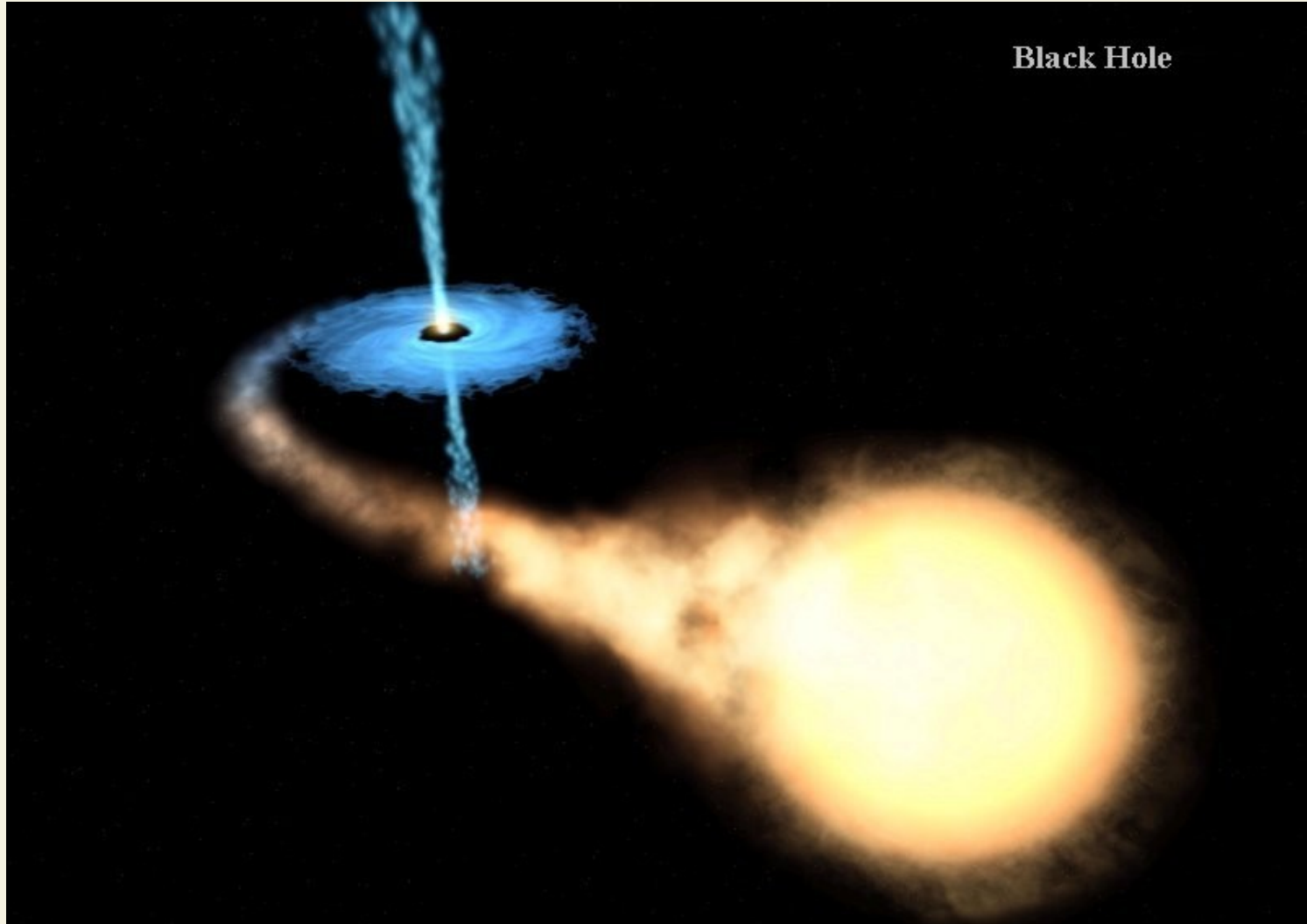
Einstein presented Feb 24, 1916

Curved Space





Black Holes



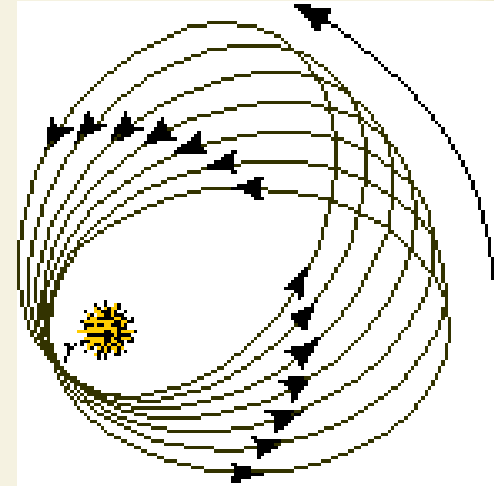
Testing of General Relativity

New Theories need to derive known results & predict new results

~ Mercury's Orbit Precession

1858 Urbain Le Verrier – $531/574$ arcsec/century

Nov 18, 1915 Einstein – GR gives 574!



~ Bending of Light

Erwin Freundlich – 1912

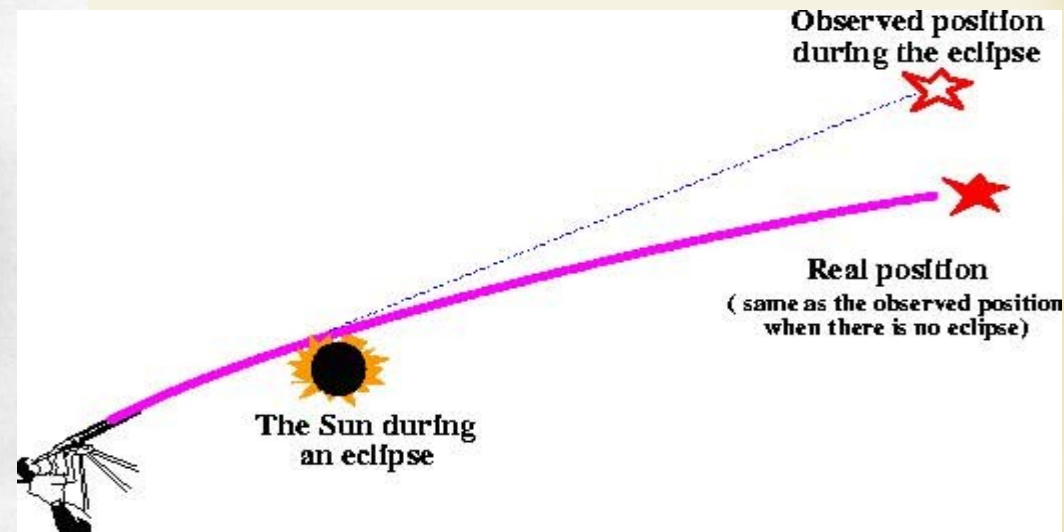
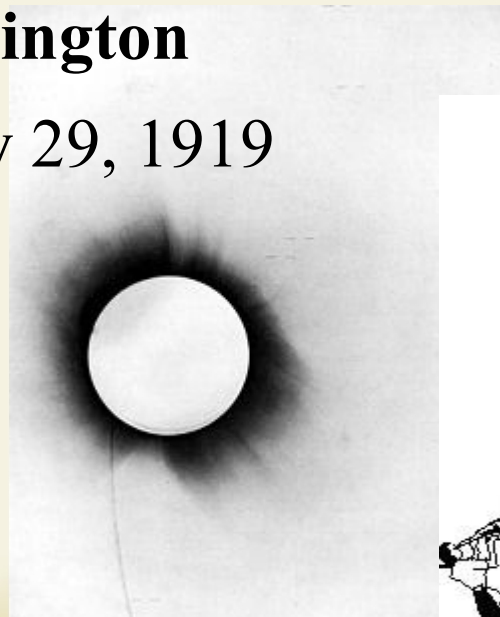
Crimea – Aug 21, 1914

Sir Arthur Eddington

Brazil – May 29, 1919

LIGHTS ALL ASKEW
IN THE HEAVENS

Men of Science More or Less
Agog Over Results of Eclipse
Observations.



1917- Cosmological Considerations

Cosmological Considerations of the General Theory of Relativity, Einstein

Cosmological Principle

The universe is the same everywhere

Homogeneous

The universe looks the same from every point

Isotropic

The universe looks the same in every direction

Einstein's Model – Not Static!

All bodies attract leading to collapse – unstable universe!

Fudge Factor - “his greatest blunder”

Einstein adds **cosmological constant, Λ**

Provides a repulsion of masses

New Theories

~ Aleksandr Friedmann (1888-1925) - 1922

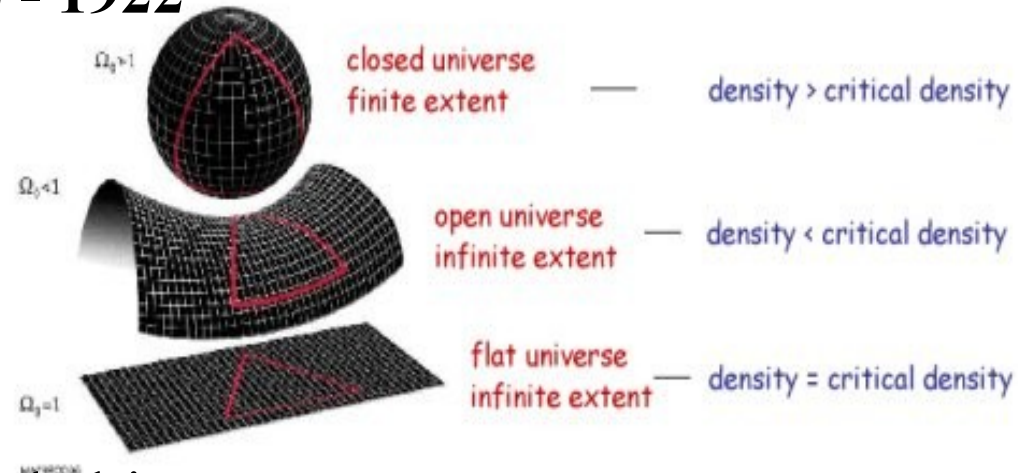
Gave universe an initial kick

Initial density gave three scenarios

Low density – forever expands

High density – re-contracts

Critical density – slows without halting



~ George Lemaitre (1894-1966) - 1927

Physicist and Priest, worked with Eddington

Rederived Friedmann's work

Consequence - traced back in time to moment of creation

Proposed cosmic rays came from early universe

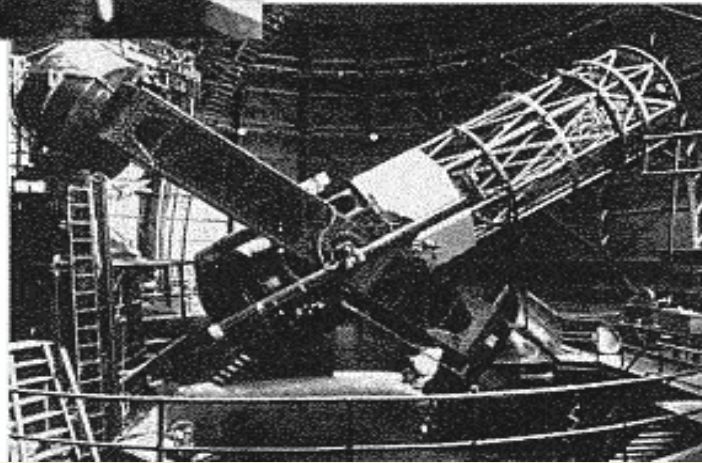
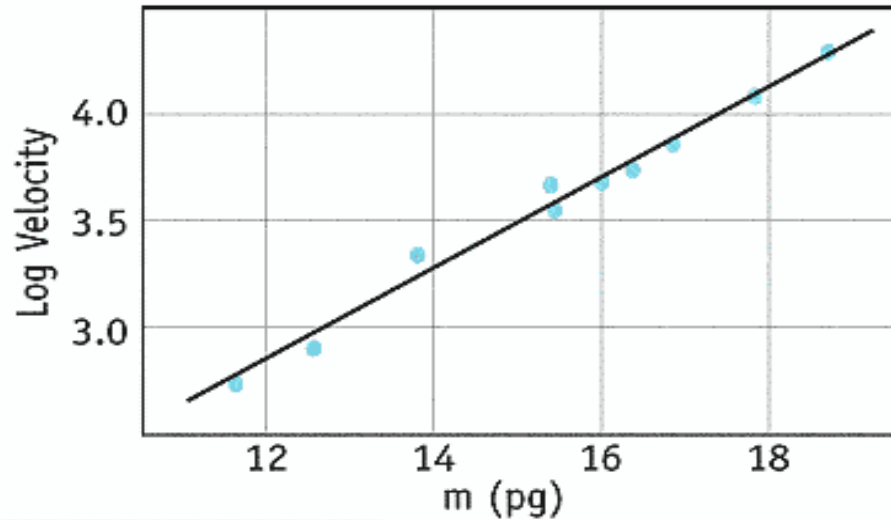
~ Both rejected by Einstein!

Edwin Powell Hubble (1889-1953)

DISCOVERY OF EXPANDING UNIVERSE



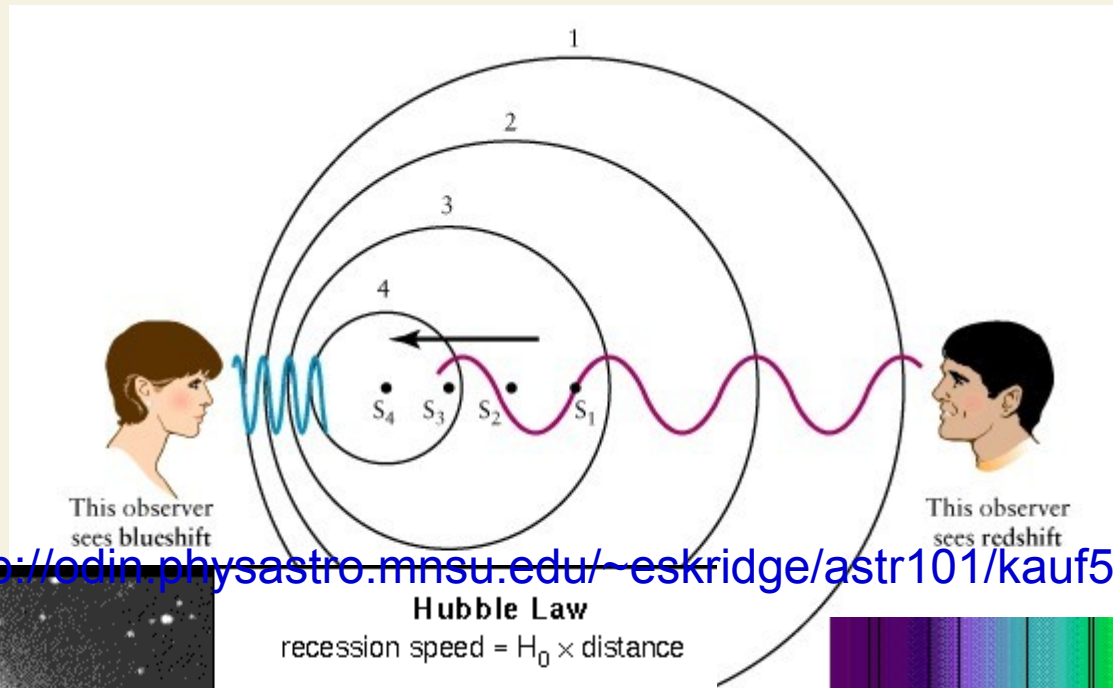
Edwin Hubble



Mt. Wilson
100 Inch
Telescope

Doppler Effect

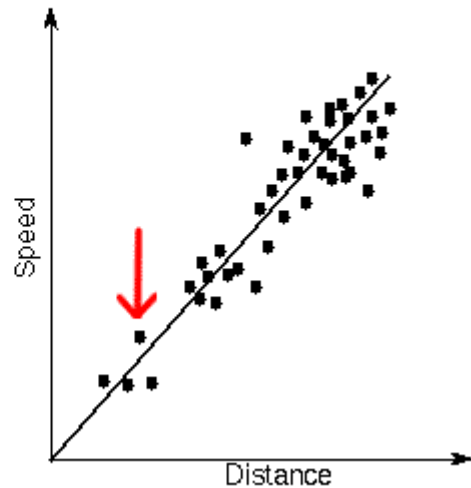
Christian Andreas Doppler (1803-1853)



http://odin.physastro.mnsu.edu/~eskridge/astr101/kauf5_23.JPG

Hubble Law

$$\text{recession speed} = H_0 \times \text{distance}$$



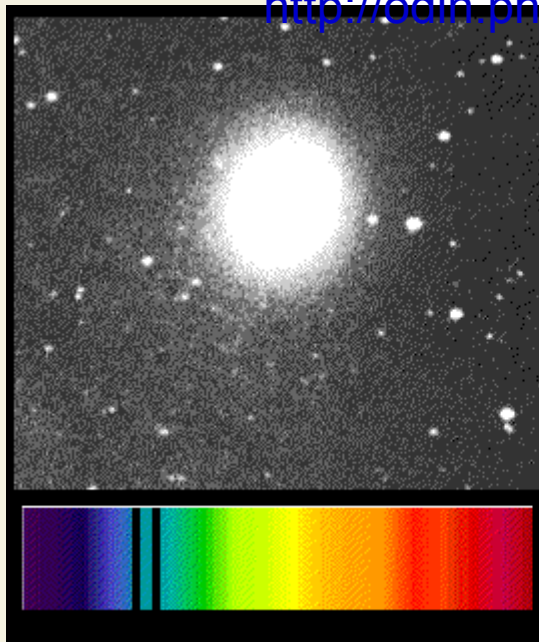
moving toward you: blueshift



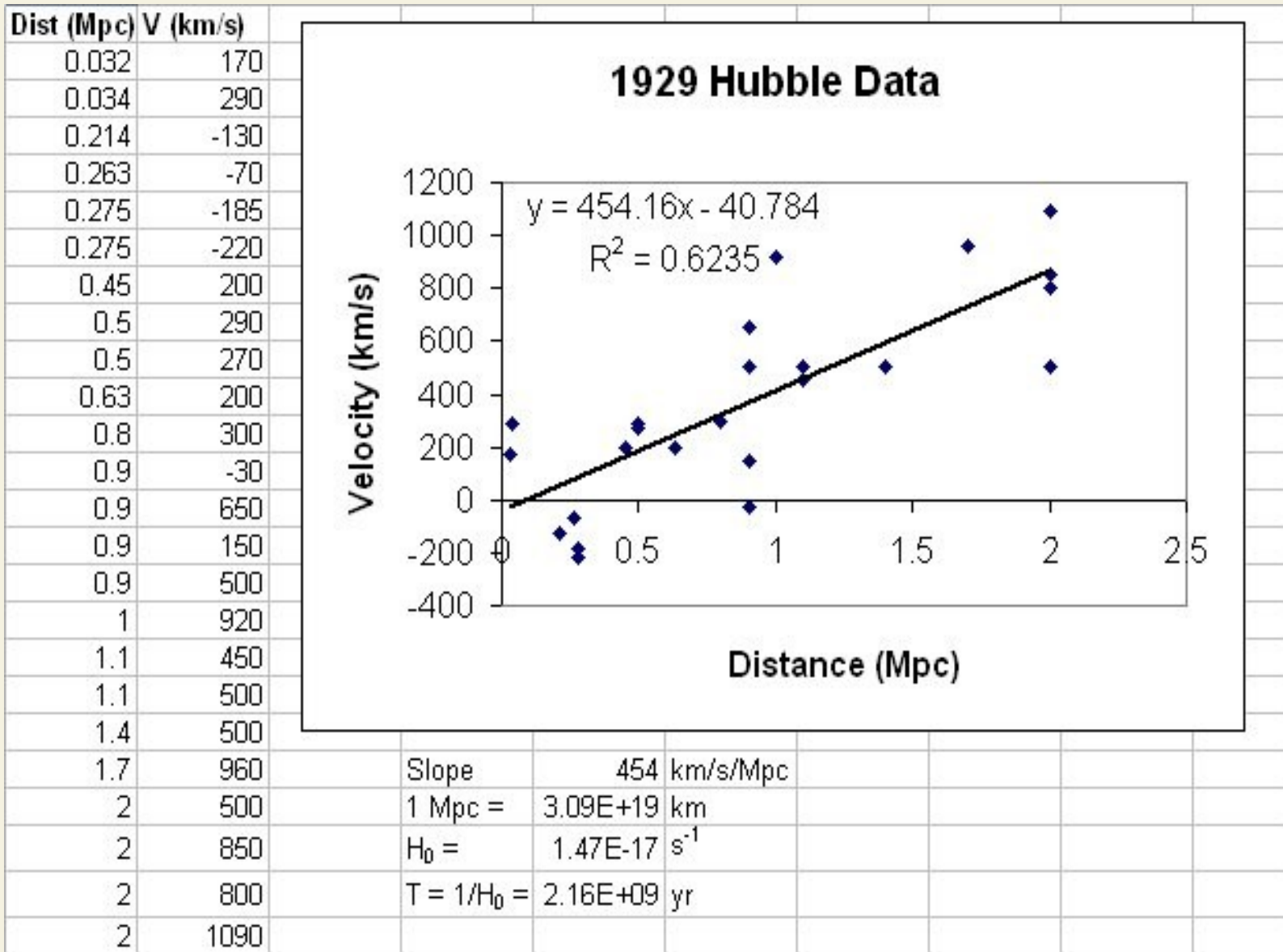
at rest



moving away from you: redshift



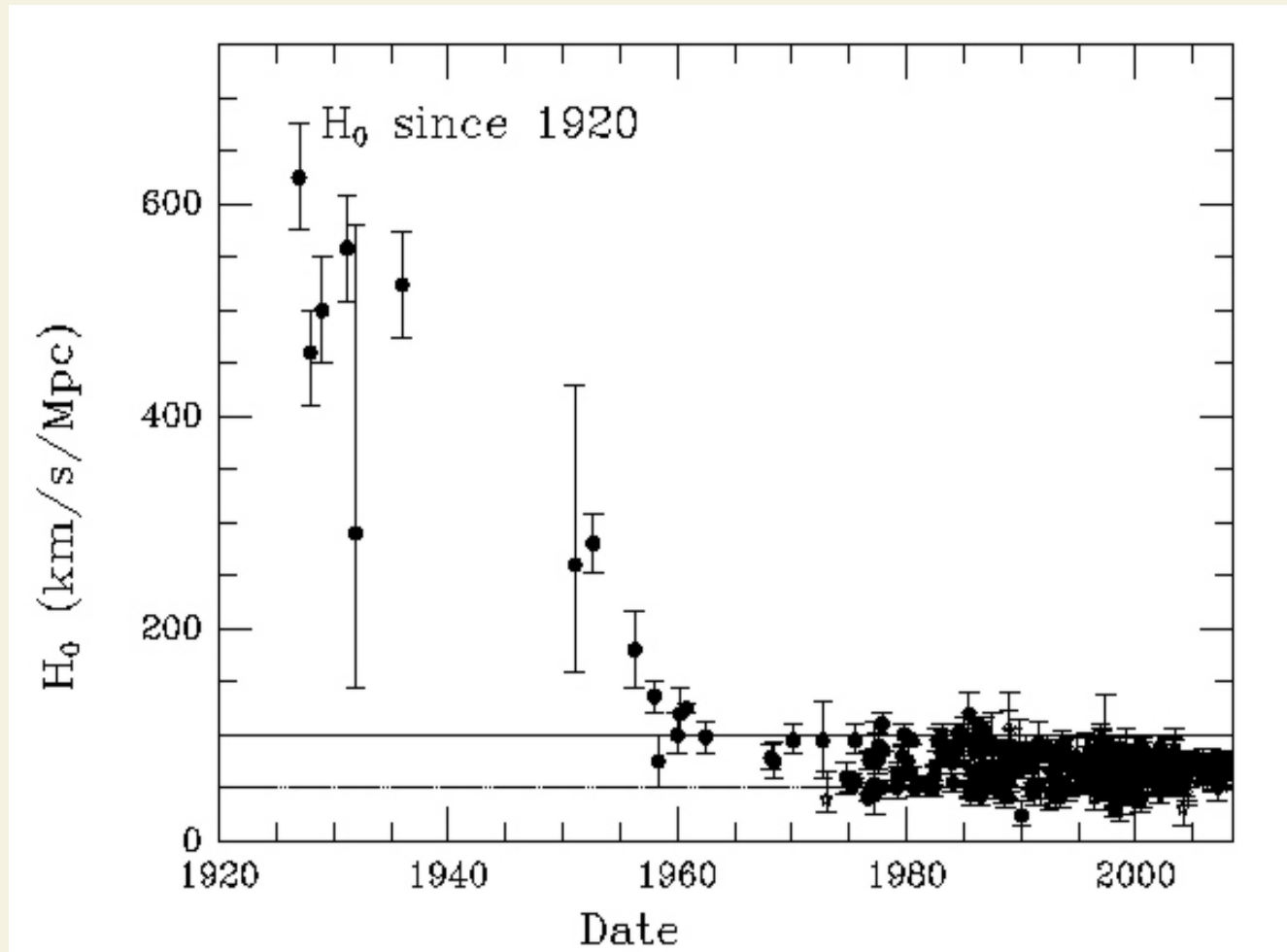
Hubble's Constant: $v = H_0 d$



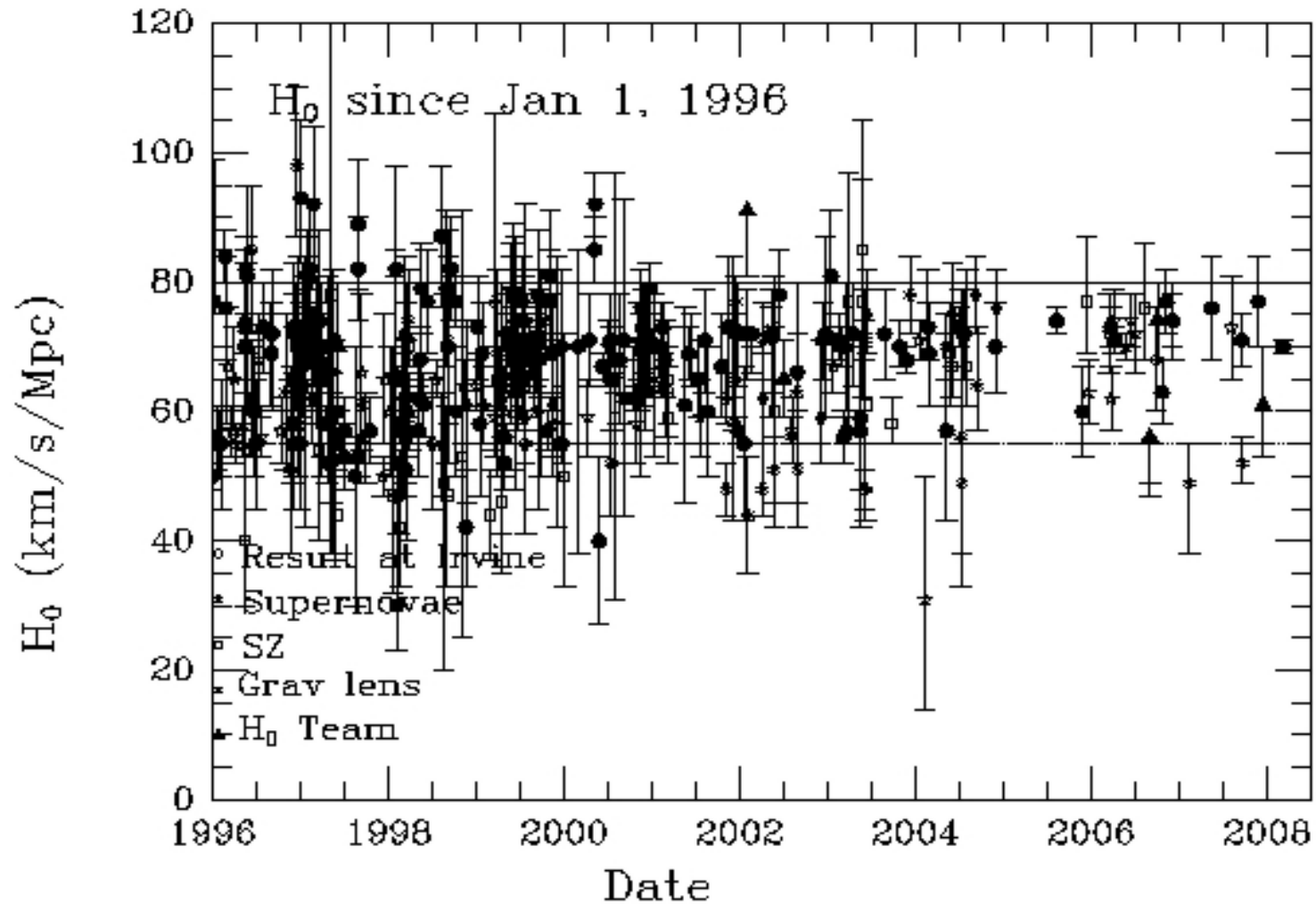
H₀ and the Age of the Universe

~ **Age** = $1/H_0 = 1/454 \text{ km/s/Mpc} = 2 \text{ billion yrs}$

~ **BUT** – 1930 - Geologists, “Earth 3 billion yrs old!”



And the winner is ...



Age of Universe

Current Value H_0 : **72 +/- 8 km/s/Mpc**

1 Mpc = 3.086×10^{22} m – try Google!

1 km/s/Mpc = 3.24×10^{-20} 1/s

$$\begin{aligned} 1/H_0 &= 4.286 \times 10^{17} \text{ s} \\ &= \mathbf{13.6 \text{ Gyr}} \end{aligned}$$

WMAP – 13.7 +/- 0.13 Gyr

If flat and matter dominated – $2/(3H_0)$ – 9 Gyr

Big Bang vs Steady State Models

Gamow, Alpher, Herman - 1948

Expansion and cooling of universe

Initial state - infinite density and temperature.

"Ylem" = protons, neutrons, and electrons in an ocean of radiation.

Computer calculation of nuclear processes

Gave off radiation => the universe is now 5K

Hoyle, Bondi, Gold - 1950

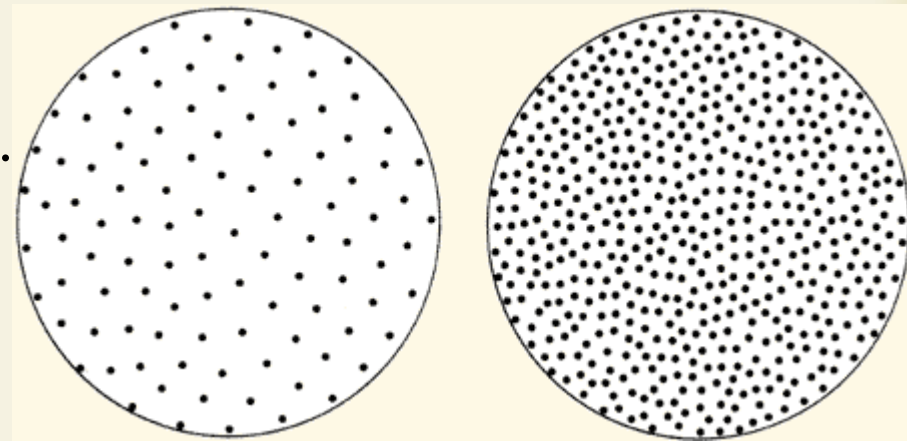
In a steady state universe the density would remain constant.

1 – Age of Universe

2 - The rate of expansion of the universe.

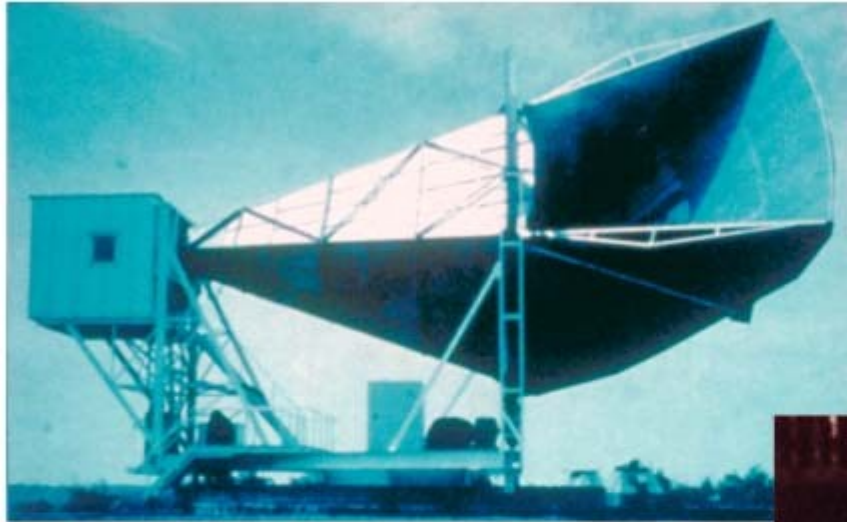
Big Bang - rate would slow

Steady State-rate would remain constant.



Arno Penzias and Robert Wilson - 1965

DISCOVERY OF COSMIC BACKGROUND



Microwave Receiver



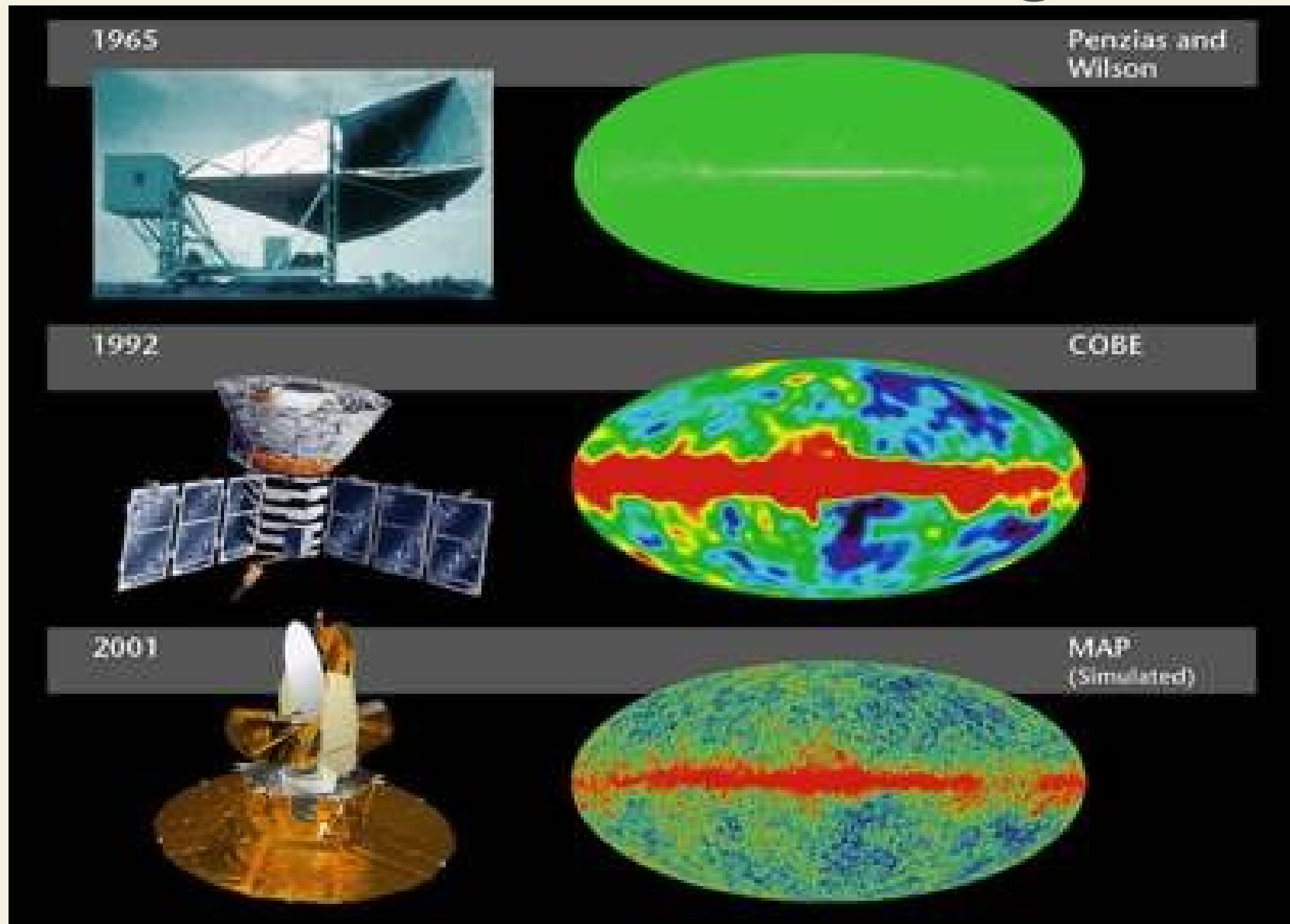
MAP990045

Robert Wilson



Arno Penzias

Cosmic Microwave Background

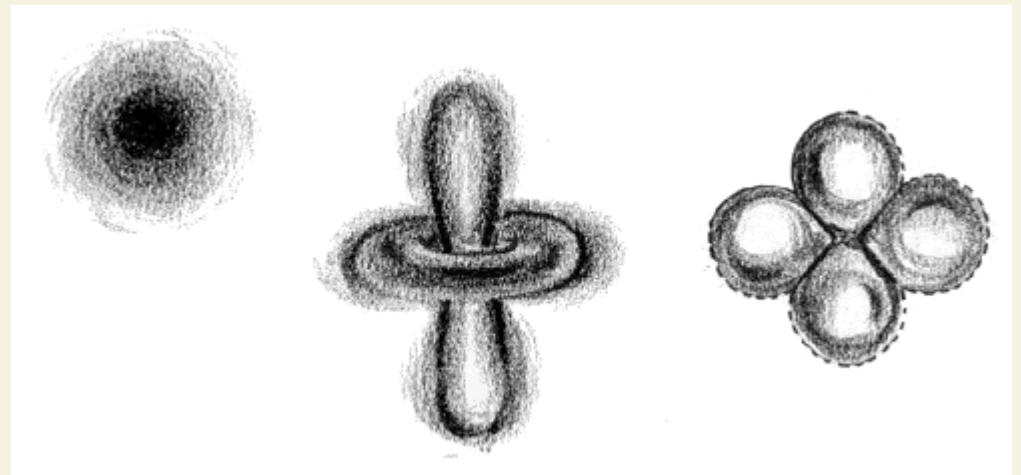
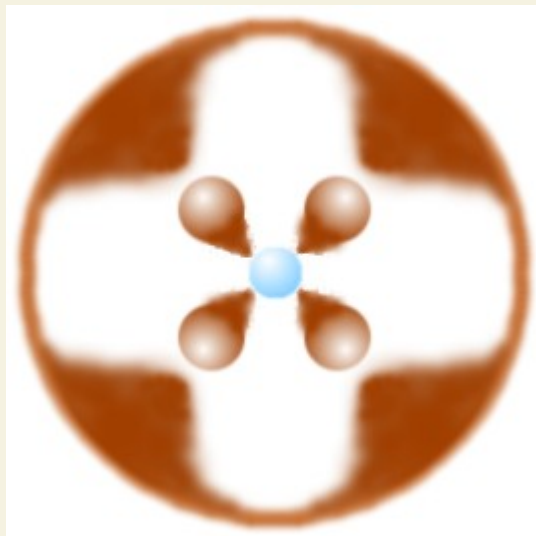
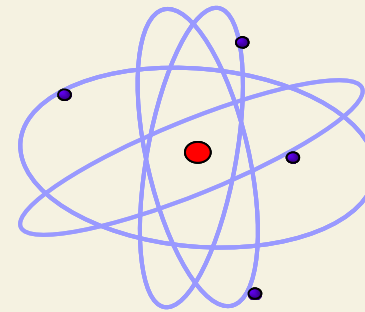
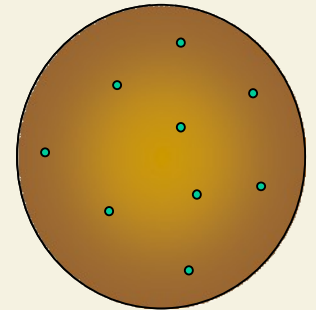


http://www.space.com/scienceastronomy/map_mission_basics_030211.html

The First Three Minutes, UNC Wilmington, College Day-2008

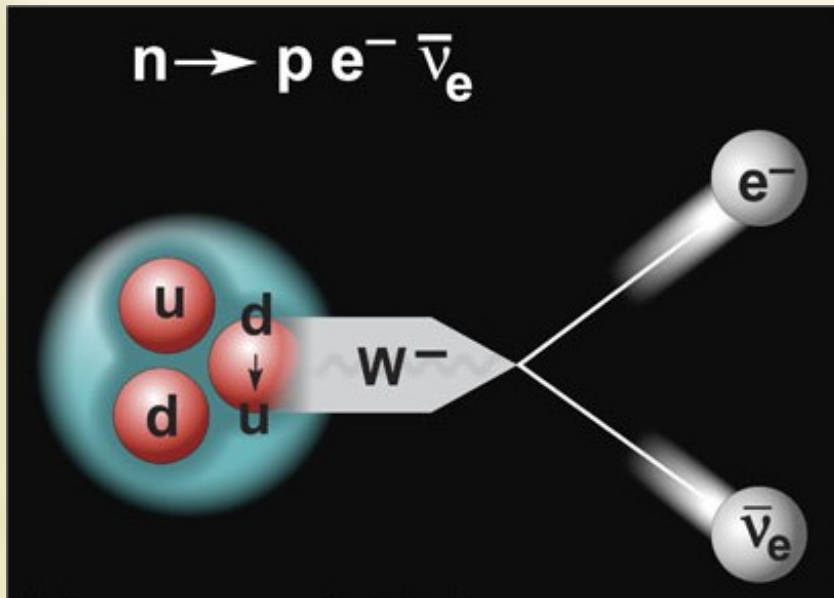
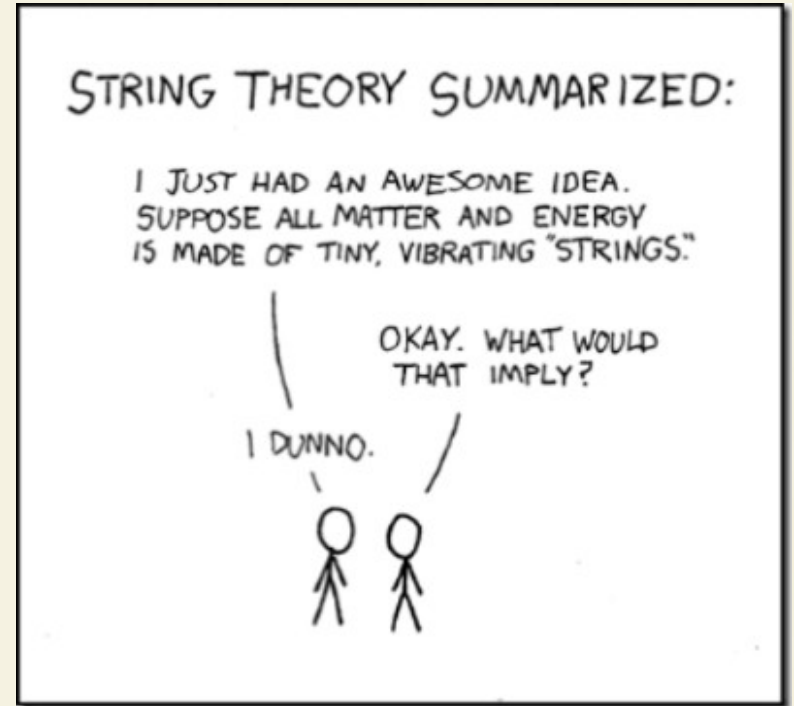
What is the Universe Made Of?

- ~ Atoms -
- ~ Electrons
- ~ Nucleus - Nucleons
- ~ Antiparticles
- ~ And ... quarks?



What Holds it Together?

- ~ Gravitational Force
- ~ Electromagnetic Force
- ~ Strong Force
- ~ Weak Force



Unification of Forces –

Electricity/Magnetism = EM

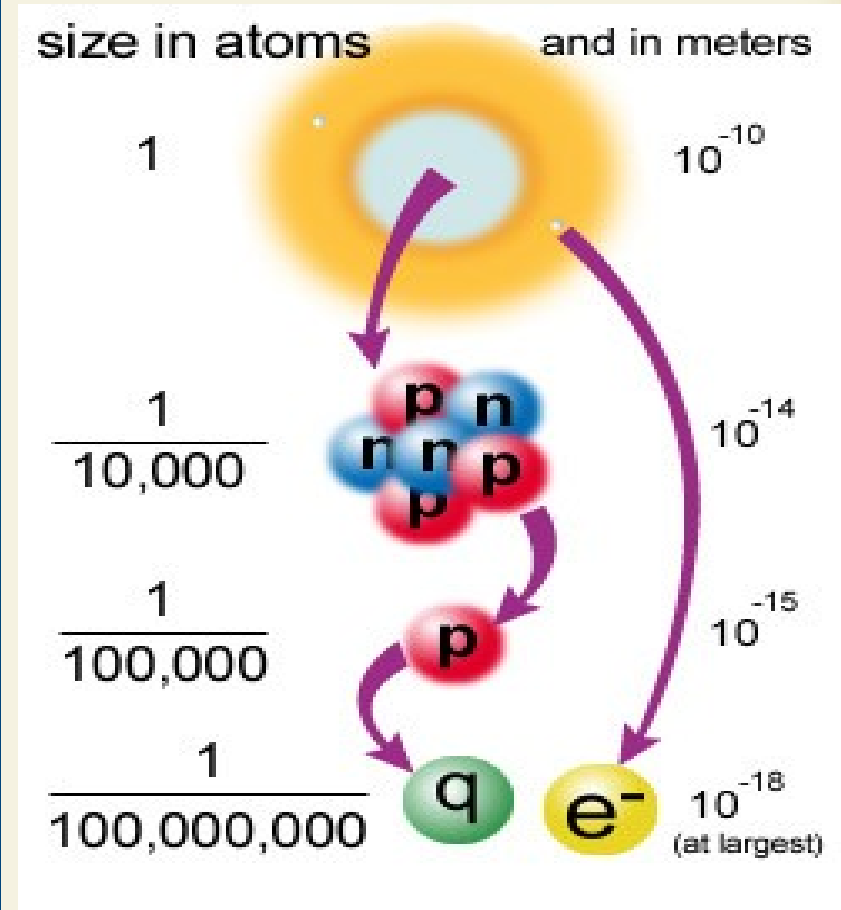
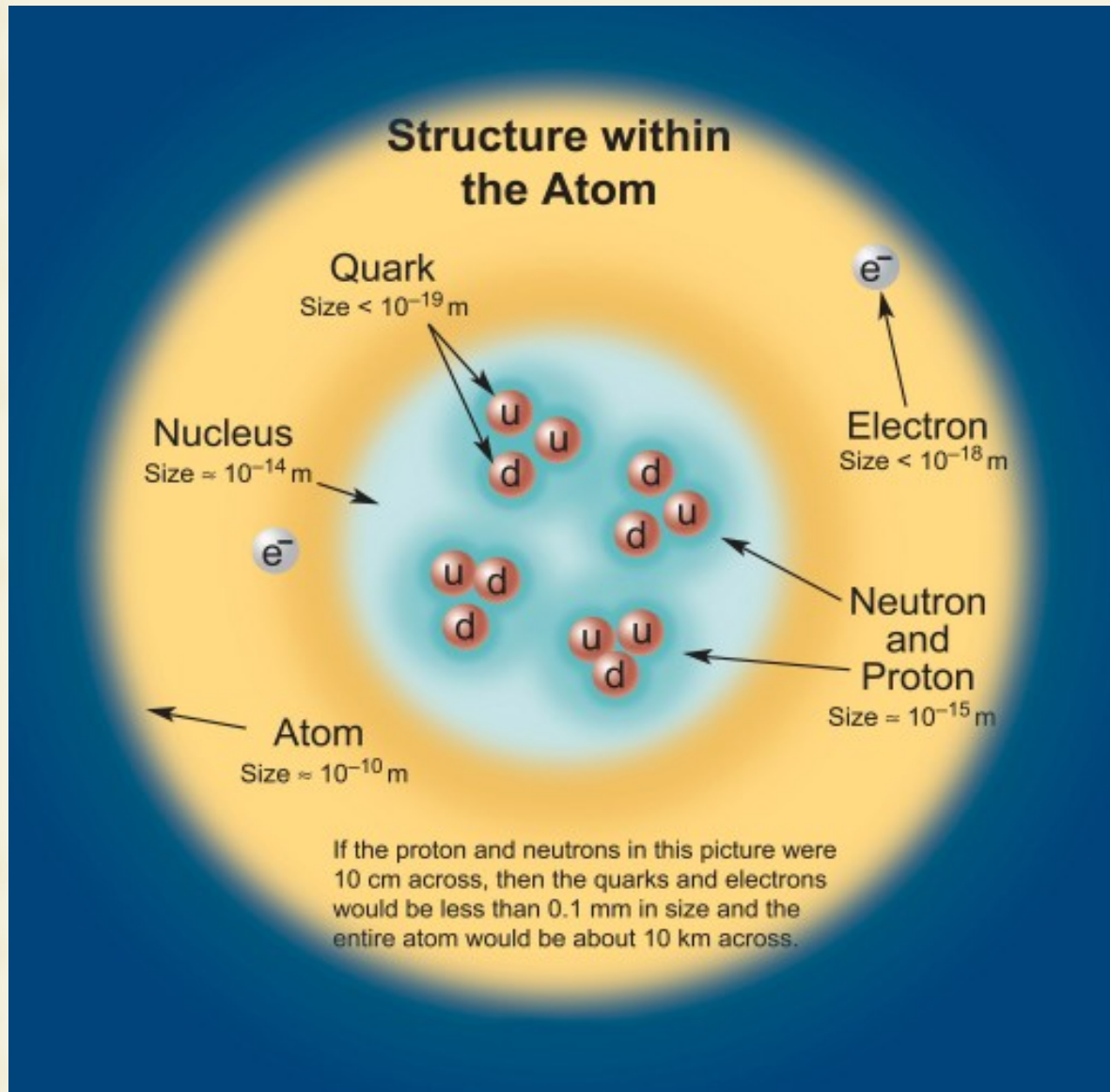
Quantum and EM = QED

QED and Weak = Electroweak

Quantum & Strong = QCD

QCD & Electroweak = **Standard Model**

The Standard Model



Particle Discoveries

1930 **Wolfgang Pauli** “**neutrino**”

1931 **Paul Dirac** **positrons/antiparticles**

1931 **James Chadwick** **neutron**.

1933-34 **Enrico Fermi** - theory of beta decay

1933-34 **Hideki Yukawa** “**pions**” between protons and neutrons.

1937 **Muon** is discovered in cosmic rays.

1946-47 “**lepton**” is introduced

1947 Pion found in cosmic rays.

1949 Discovery of **K⁺** via its decay.

1950 The neutral pion is discovered.

1951 λ^0 and the K^0 .

1952 **delta** particle: (δ^{++} , δ^+ , δ^0 , and δ^- .)

1953 The beginning of a “**particle explosion**”

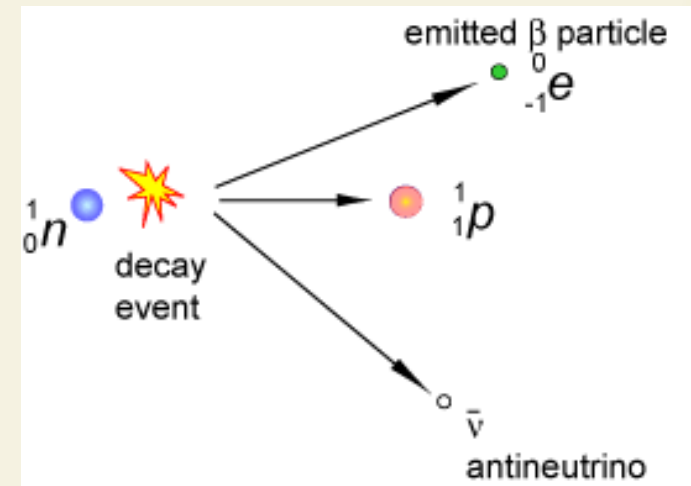
1953-57 **internal structure** for protons and neutrons

1957 **Julian Schwinger** **unification of weak and electromagnetic interactions.**

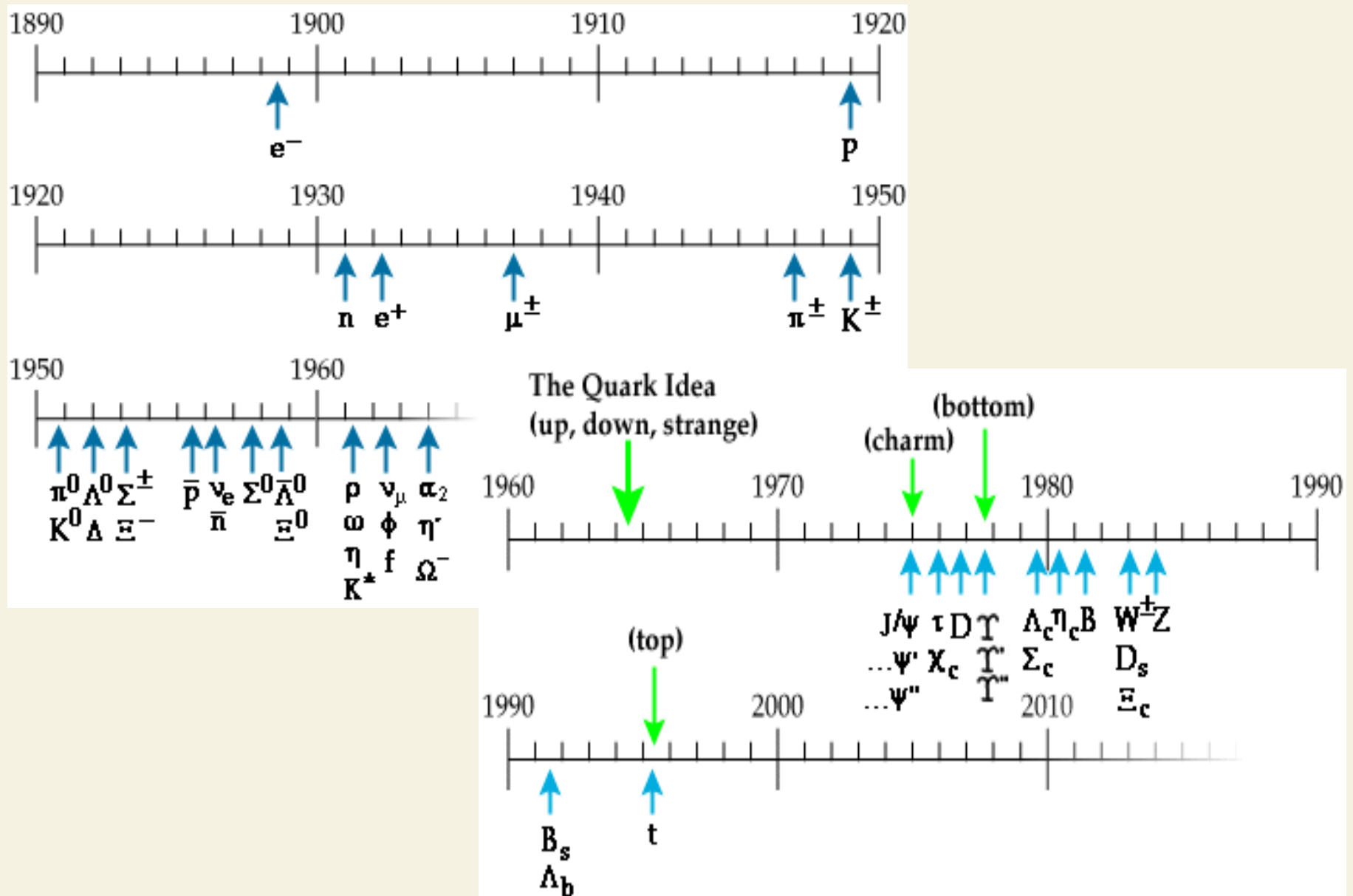
1957-59 **Julian Schwinger, Sidney Bludman, and Sheldon Glashow,**

weak interactions are mediated **W⁺** and **W⁻**

1962 Experiments verify two distinct types of **neutrinos** (electron and muon neutrinos).



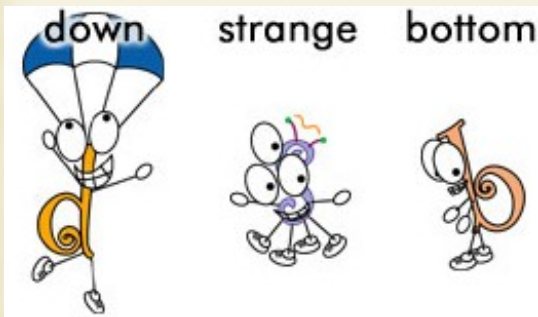
The Particle Explosion



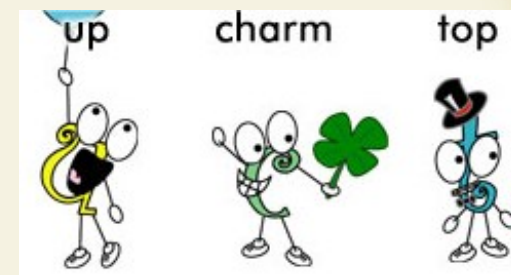
The First Three Minutes, UNC Wilmington, College Day-2008

Timeline -Quarks

- 1964 **Murray Gell-Mann** and **George Zweig** tentatively put forth **quarks**.
mesons and baryons are composites of three quarks or antiquarks:
up, down, strange
- 1964 Leptons suggest fourth quark, **charm** - **Sheldon Glashow** and **James Bjorken**
- 1965 **O.W. Greenberg**, **M.Y. Han**, and **Yoichiro Nambu** introduce **color charge**.
- 1967 **Steven Weinberg** and **Abdus Salam**
Unified electromagnetic and weak interactions, predict Higgs Boson
Theory needs neutral, weakly interacting boson that mediates weak interaction
- 1968-69 Stanford Linear Accelerator - electrons are scattered off protons,
Electrons appeared to be bouncing off small hard cores inside proton.
James Bjorken and **Richard Feynman** analyzed as particles inside proton
- 1970 **Sheldon Glashow**, **John Iliopoulos**, and **Luciano Maiani**
recognize the importance of a fourth type of quark in **Standard Model**.
- 1973 **Donald Perkins**, re-analyzes old CERN data, finds indications of **Z⁰** exchange.
- 1973 A quantum field theory of strong interaction - **quantum chromodynamics (QCD)**.



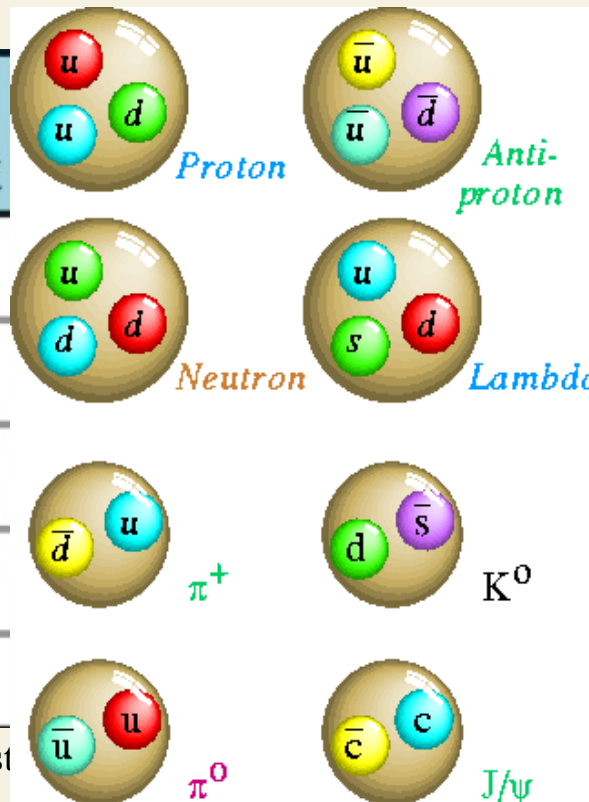
The First Three Minutes, UNC Wilmington, College Day-2008



Timeline - Standard Model

- 1974 **Burton Richter** and **Samuel Ting**, - “**J/psi**” particle, a charm-anticharm meson.
- 1976 **Gerson Goldhaber** and **Francois Pierre** find the **D⁰** meson (anti-up and charm).
- 1976 The **tau** lepton is discovered by **Martin Perl** and collaborators at SLAC.
- 1977 **Leon Lederman** and his collaborators at Fermilab discover the **bottom** quark.
- 1978 **Charles Prescott** and **Richard Taylor** observe a **Z⁰** mediated weak interaction
- 1983 Find **W[±]** and **Z⁰** intermediate bosons using the CERN synchrotron
using p and anti-p techniques of **Carlo Rubbia** and **Simon Van der Meer**
- 1995 The **top** quark found at the unexpected mass of 175 GeV

Symbol	Name	Quark content
p	proton	uud
p̄	antiproton	ūūd̄
n	neutron	udd
Λ	lambda	uds
Ω⁻	omega	sss



Symbol	Name	Quark content
π⁺	pion	u\bar{d}
K⁻	kaon	s\bar{u}
ρ⁺	rho	u\bar{d}
B⁰	B-zero	d\bar{b}
η_c	eta-c	c\bar{c}

The First

Illeg Day-2008

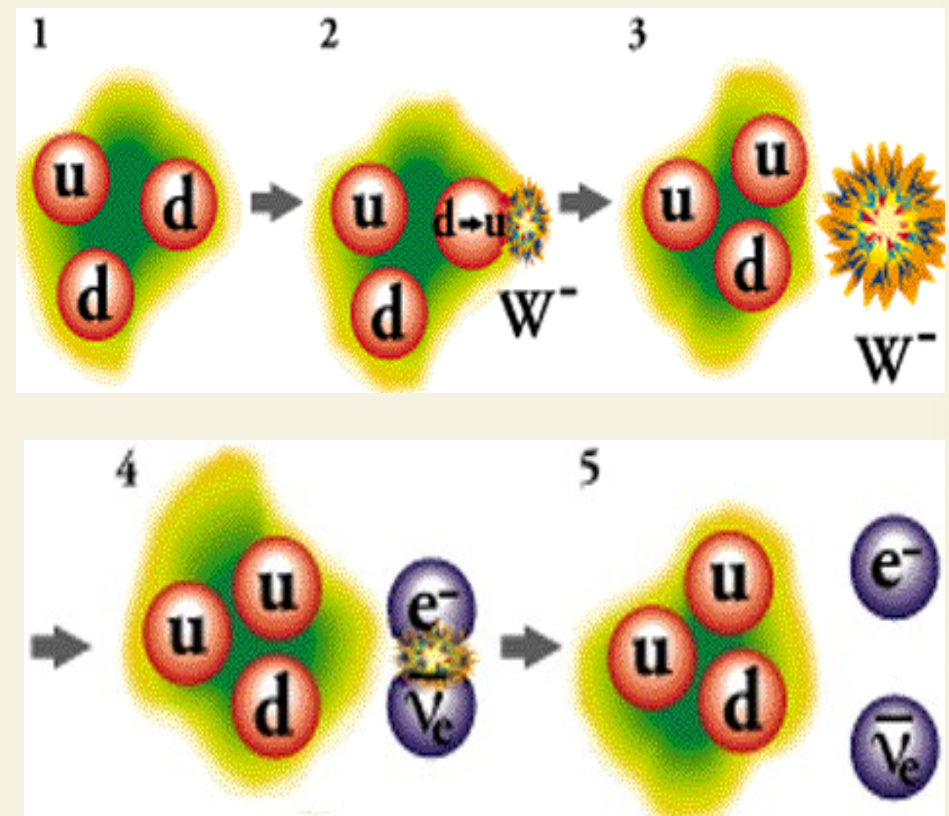
Quarks and Leptons

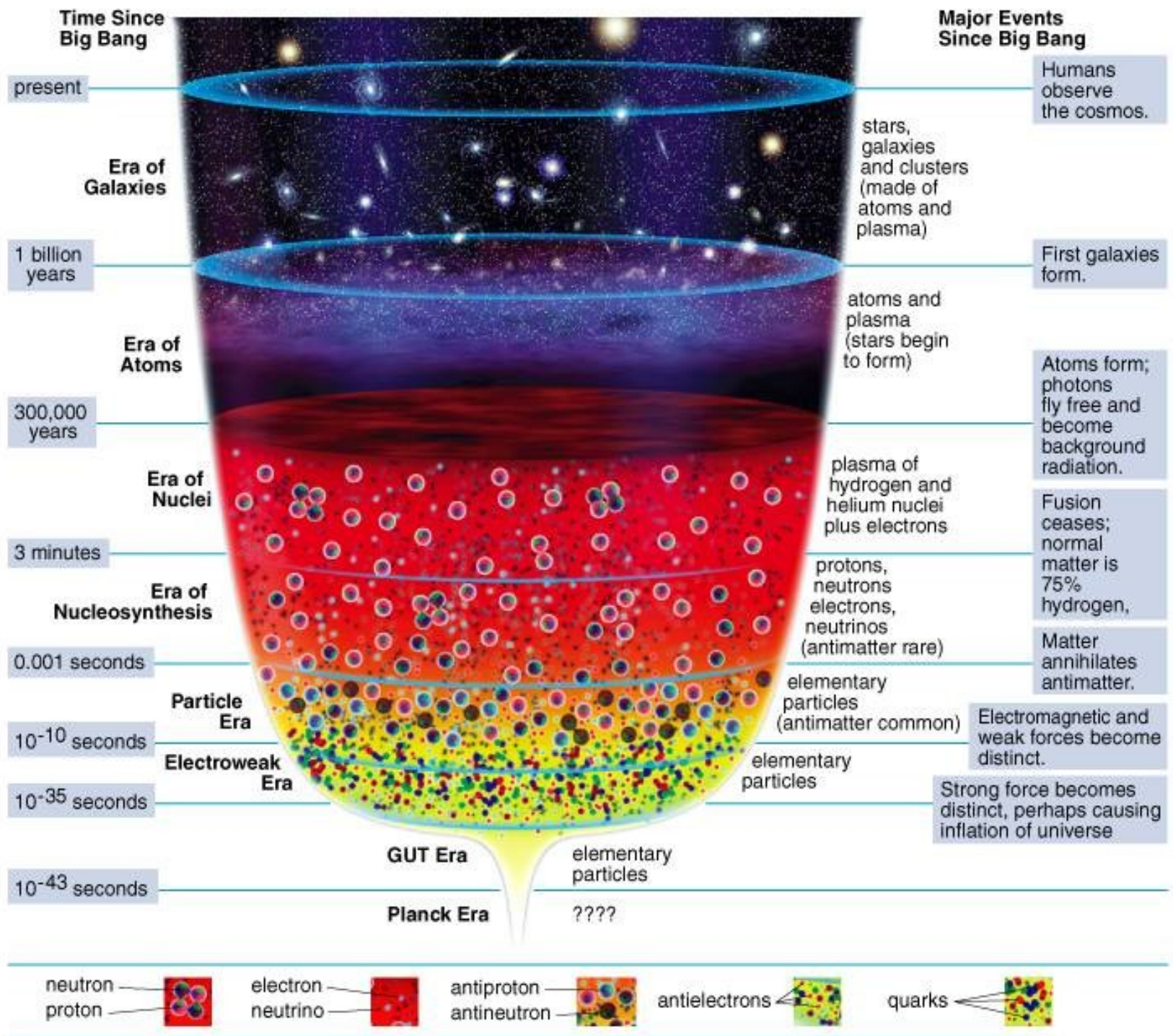
Quarks	<i>u</i>	<i>c</i>	<i>t</i>	
	up	charm	top	
	<i>d</i>	<i>s</i>	<i>b</i>	
	down	strange	bottom	
	Leptons	ν_e	ν_μ	ν_τ
		e- Neutrino	μ - Neutrino	τ - Neutrino
<i>e</i>		μ	τ	
	electron	muon	tau	
	I	II	III	
	The Generations of Matter			

Neutron - udd

Proton - uud

$$n \Rightarrow p + e^- + \bar{\nu}_e$$

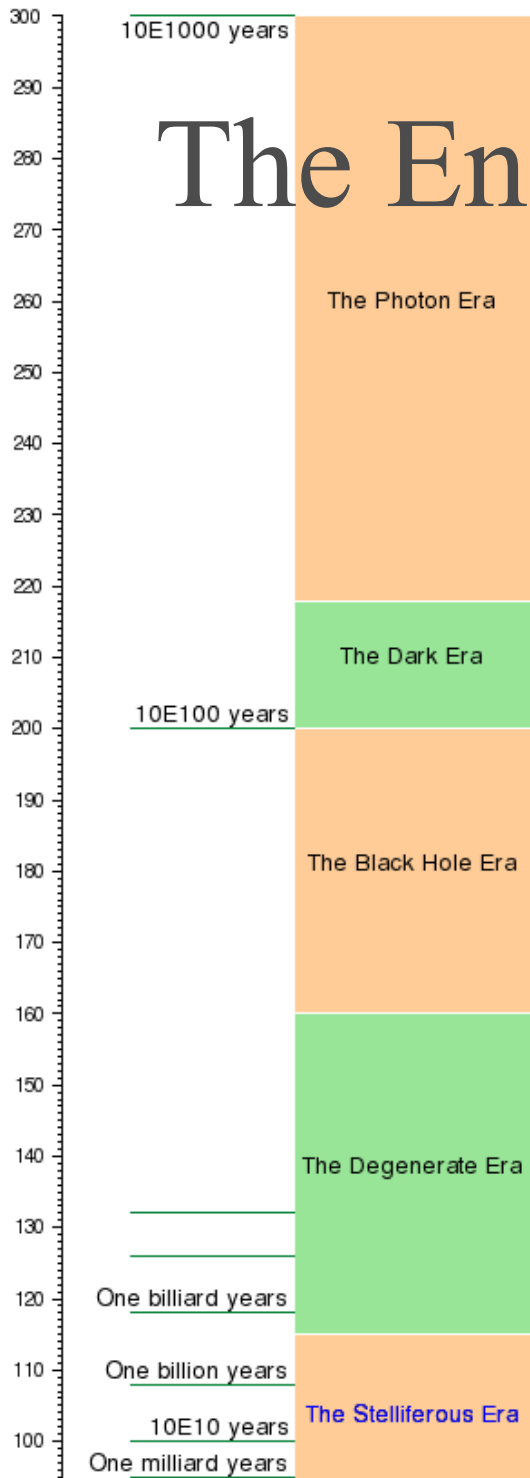




Double-logarithmic time:
100*log log year

Heat Death

The End of the Universe



- The last supermassive black holes have evaporated
- Small and medium sized black holes have evaporated
- All proton decay. The matter that stars and life are built of no longer exist.
- Proton decay to 1/2
- Galaxies no longer exist. Stars flung out of orbit or consumed by larger bodies as black holes.
- Formation of new stars ceases. The last star dies down. Death of unintelligent water-based life.
- 200 milliard years. The Sun dies down.
- 13.7 milliard years: this present day
- 8.7 milliard years. Formation of the Sun
- 600 million years. Formation of the first galaxy
- 100 million years. First star began to shine
- Life on Earth
- Reionization

Universe Accelerating?



The expansion of the universe appears to be accelerating. Is this due to Einstein's Cosmological Constant? If not, will experiments reveal a new force of nature or even extra (hidden) dimensions of space?

More Mysteries

Why No Antimatter?



Matter and antimatter were created in the Big Bang. Why do we now see only matter except for the tiny amounts of antimatter that we make in the lab and observe in cosmic rays?

Dark Matter?



Invisible forms of matter make up much of the mass observed in galaxies and clusters of galaxies. Does this dark matter consist of new types of particles that interact very weakly with ordinary matter?

Origin of Mass?



In the Standard Model, for fundamental particles to have masses, there must exist a particle called the Higgs boson. Will it be discovered soon? Is supersymmetry theory correct in predicting more than one type of Higgs?

Further Reading

The First Three Minutes, S. Weinberg

A Briefer History of Time, S. Hawking and L. Mlodinow

Many Worlds in One, A. Vilenkin

Endless Universe: Beyond the Big Bang, P. J. Steinhardt and N. Turok

Big Bang: The Origin of the Universe (P.S.), S. Singh

Dark Side of the Universe: Dark Matter, Dark Energy, and the Fate of the Cosmos, I. Nicolson

The Elegant Universe or The Fabric of the Cosmos, B. Greene

Parallel Worlds: A Journey Through Creation, Higher Dimensions, and the Future of the Cosmos, M. Kaku

Warped Passages: Unraveling the Mysteries of the Universe's Hidden Dimensions, L. Randall

The Trouble with Physics, L. Smolin

A Brief History of Time, S. Hawking

What Don't We Know?



Galaxy Cluster Abell 1689

Hubble Space Telescope • Advanced Camera for Surveys

NASA, N. Benitez (JHU), T. Broadhurst (The Hebrew University), H. Ford (JHU), M. Clampin (STScI), G. Hartig (STScI), G. Illingworth (UCO/Lick Observatory), the ACS Science Team and ESA
STScI-PRC03-01a



Questions
are
guaranteed in
life;
Answers
aren't.

More information:

<http://people.uncw.edu/hermanr/>

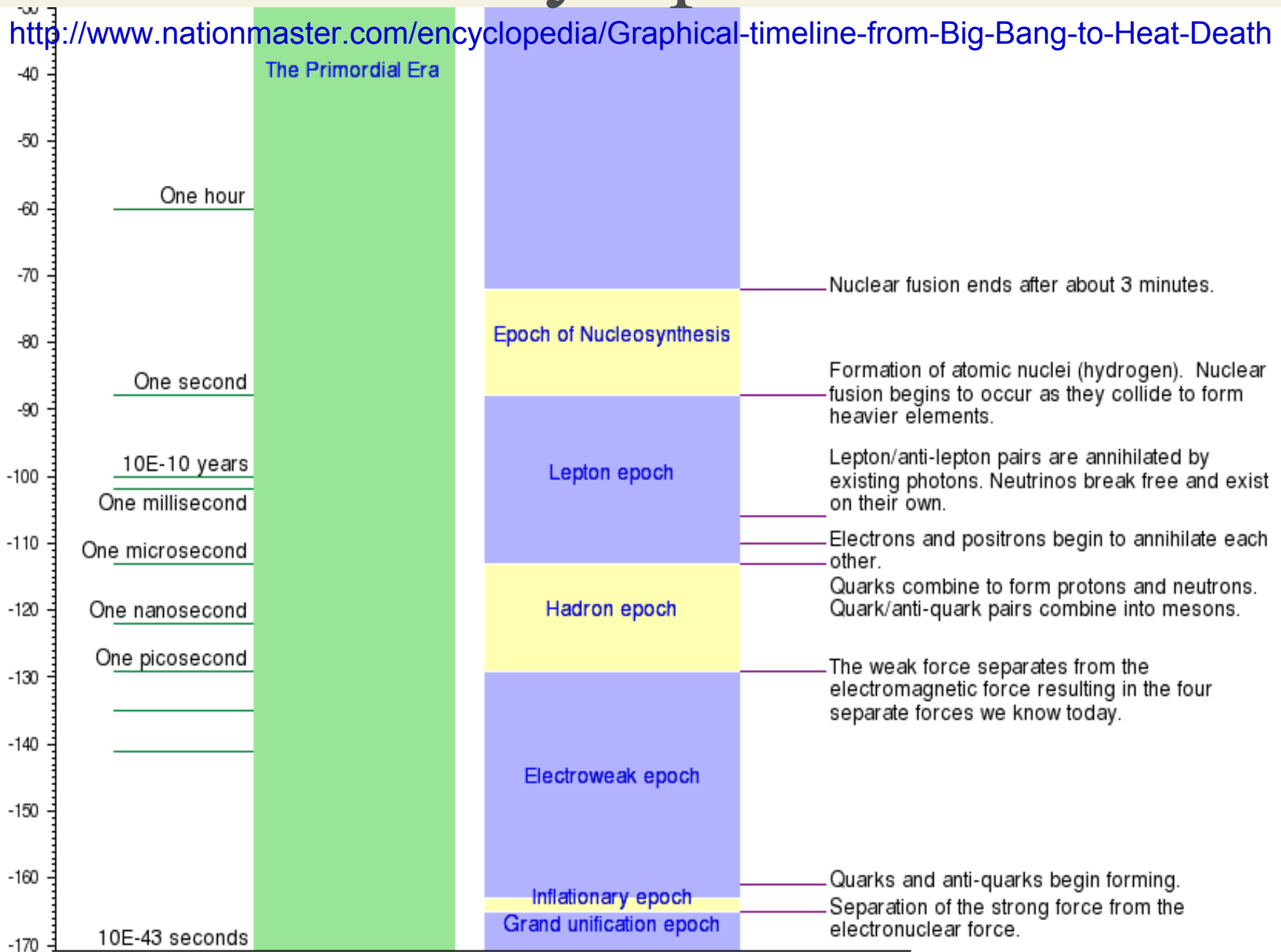
The History of the Universe in 200 Words or Less

Copyright 1996-1997 by Eric Schulman.

Quantum fluctuation. Inflation. Expansion. Strong nuclear interaction. Particle-antiparticle annihilation. Deuterium and helium production. Density perturbations. Recombination. Blackbody radiation. Local contraction. Cluster formation. Reionization? Violent relaxation. Virialization. Biased galaxy formation? Turbulent fragmentation. Contraction. Ionization. Compression. Opaque hydrogen. Massive star formation. Deuterium ignition. Hydrogen fusion. Hydrogen depletion. Core contraction. Envelope expansion. Helium fusion. Carbon, oxygen, and silicon fusion. Iron production. Implosion. Supernova explosion. Metals injection. Star formation. Supernova explosions. Star formation. Condensation. Planetesimal accretion. Planetary differentiation. Crust solidification. Volatile gas expulsion. Water condensation. Water dissociation. Ozone production. Ultraviolet absorption. Photosynthetic unicellular organisms. Oxidation. Mutation. Natural selection and evolution. Respiration. Cell differentiation. Sexual reproduction. Fossilization. Land exploration. Dinosaur extinction. Mammal expansion. Glaciation. Homo sapiens manifestation. Animal domestication. Food surplus production. Civilization! Innovation. Exploration. Religion. Warring nations. Empire creation and destruction. Exploration. Colonization. Taxation without representation. Revolution. Constitution. Election. Expansion. Industrialization. Rebellion. Emancipation Proclamation. Invention. Mass production. Urbanization. Immigration. World conflagration. League of Nations. Suffrage extension. Depression. World conflagration. Fission explosions. United Nations. Space exploration. Assassinations. Lunar excursions. Resignation. Computerization. World Trade Organization. Terrorism. Internet expansion. Reunification. Dissolution. World-Wide Web creation. Composition. Extrapolation?

Early Epochs

<http://www.nationmaster.com/encyclopedia/Graphical-timeline-from-Big-Bang-to-Heat-Death>



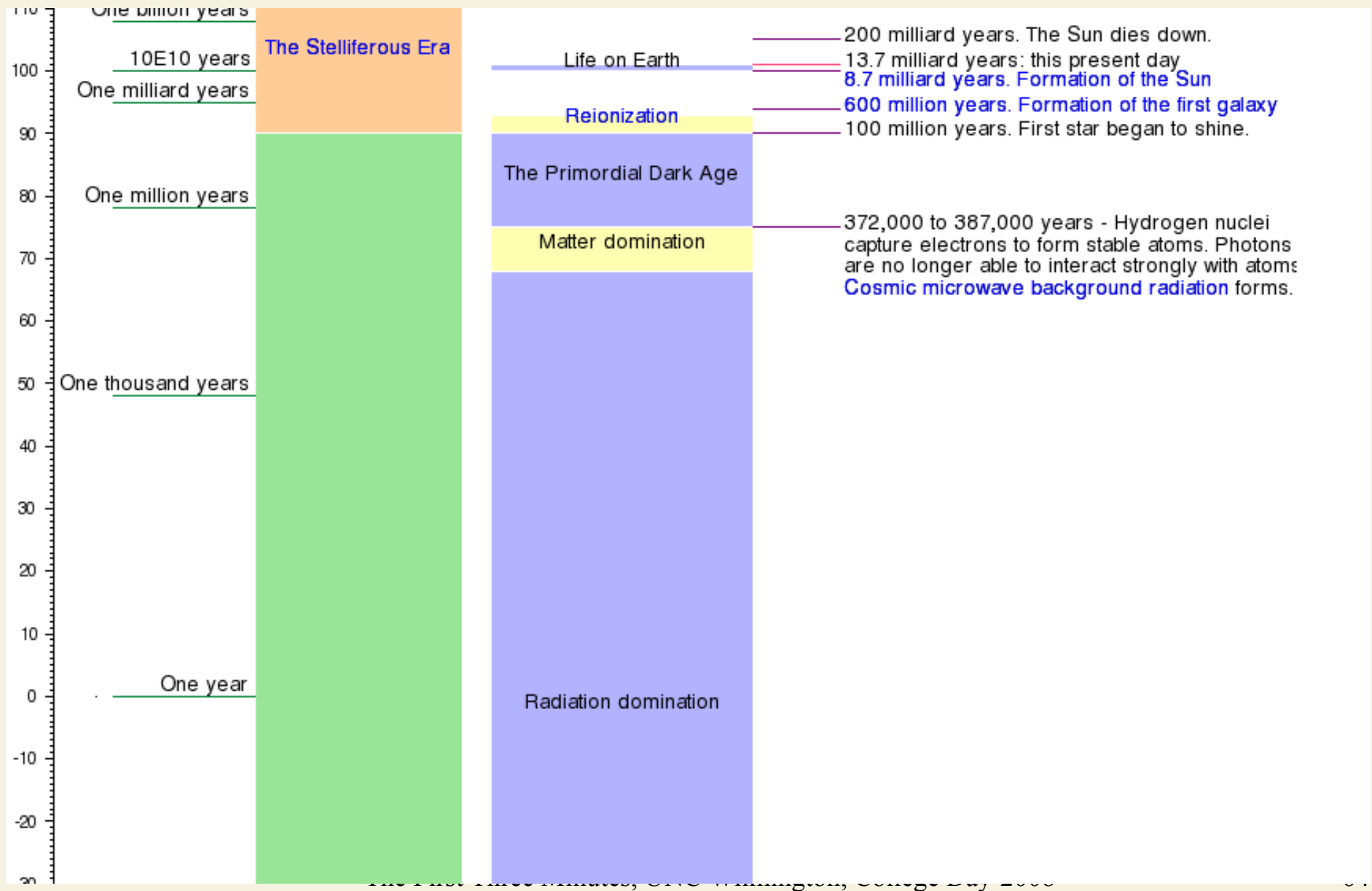
0: Linear time

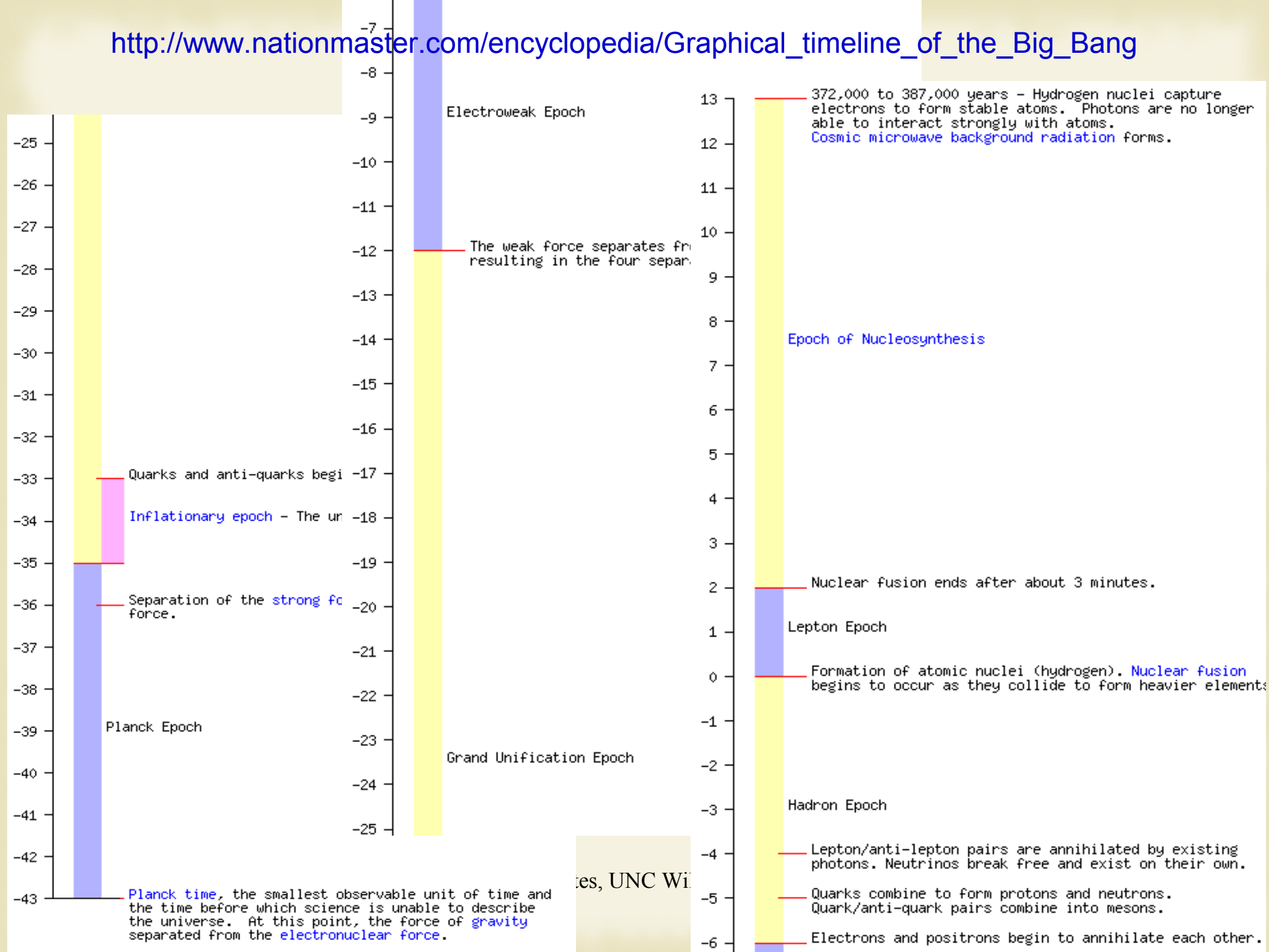
Big Bang

Planck epoch

Planck time, the smallest observable unit of time and the time before which science is unable to describe the universe. At this point, the force of gravity separated from the electronuclear force.

Later Epochs





Supernova 1987A



<http://www.oberlin.edu/physics/dstyler/Astronomy/S>



<http://science.nationalgeographic.com/staticfile>